

Towards a new paradigm of dust structure in AGN: Dissecting the mid-IR emission of Circinus galaxy

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with

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illustration: M. Kornmesser (ESO)

“I don't always study AGN. But when I do, I prefer Circinus and NGC 1068”

– The Most Interesting Man in the World

TORUS 2018

The many faces of the AGN obscuration

Puerto Varas, Chile

10-14 December 2018



Models of thick turbulent gas disks with magnetocentrifugal winds in AGN and their application to **Circinus** and NGC 1068

Towards a new paradigm of dust structure in AGN: Dissecting the mid-IR emission of **Circinus** galaxy

The first IR/X-ray model of the circumnuclear environment of the **Circinus** Galaxy

Circumnuclear Multi-phase Gas in the **Circinus** Galaxy Revealed with ALMA

The compact molecular torus in the **Circinus** galaxy constrained by ALMA

X-Ray Spectral Model from Clumpy Torus and Its Application to **Circinus** Galaxy

The many 'faces' of the molecular torus of NGC **1068**

The counter rotating molecular torus in NGC **1068**

ALMA reveals a rotating dense molecular torus in NGC **1068**

NUclei of GALaxies (NUGA) resolved by ALMA

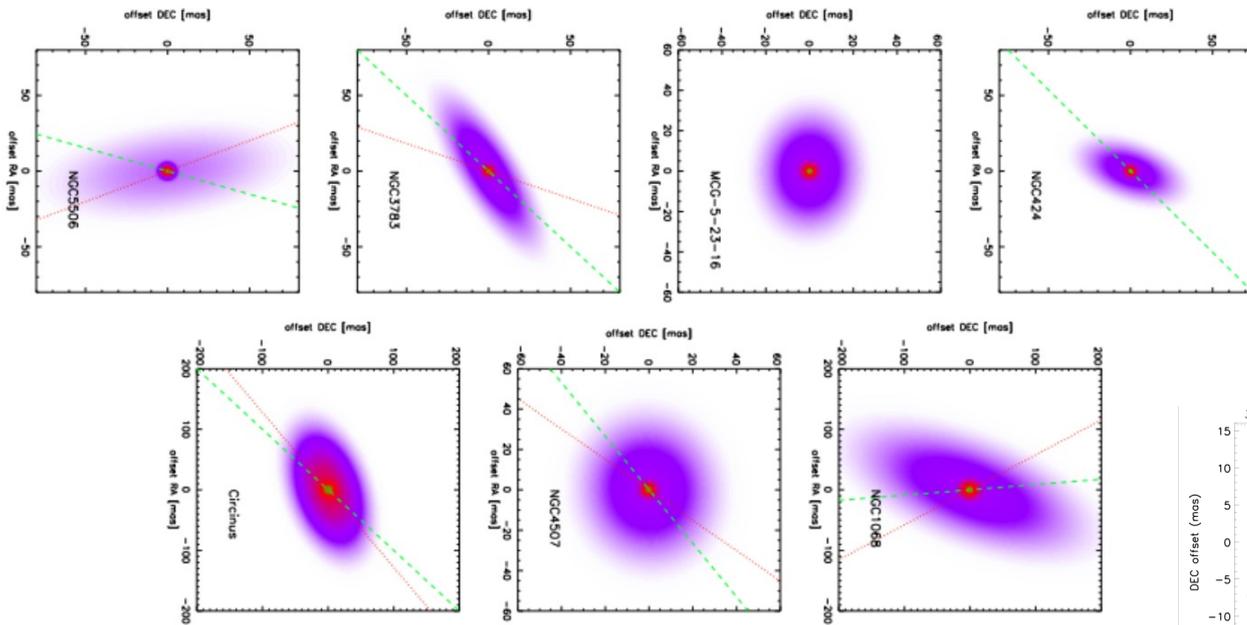
First successful MATISSE observations of NGC **1068**

The dust and cloud distribution of the torus of NGC **1068**

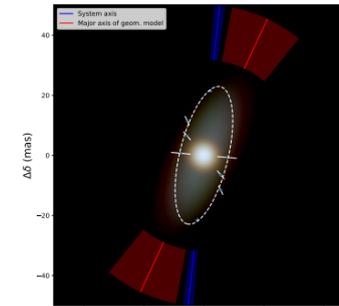
Investigating the nature and geometry of NGC **1068** through NuSTAR observations and future X-ray polarimetry

Models of thick turbulent gas disks with magnetocentrifugal winds in AGN and their application to Circinus and NGC **1068**

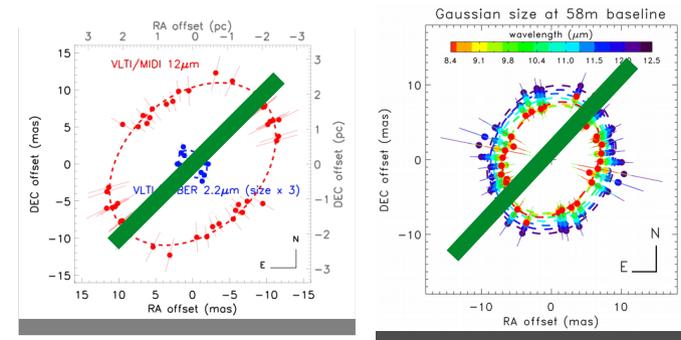
VLT/MIDI: polar elongation on pc-scale



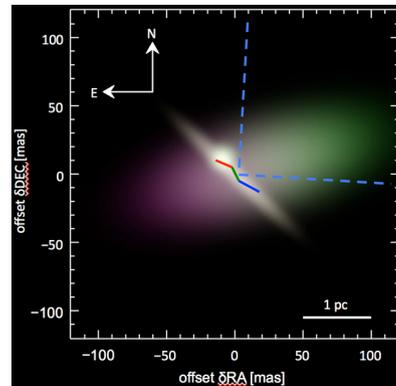
López Gonzaga et al. (2014)



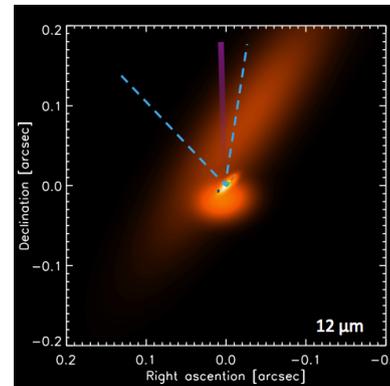
Leftley et al. (2018)



Hönig et al. (2013) Hönig et al. (2012)



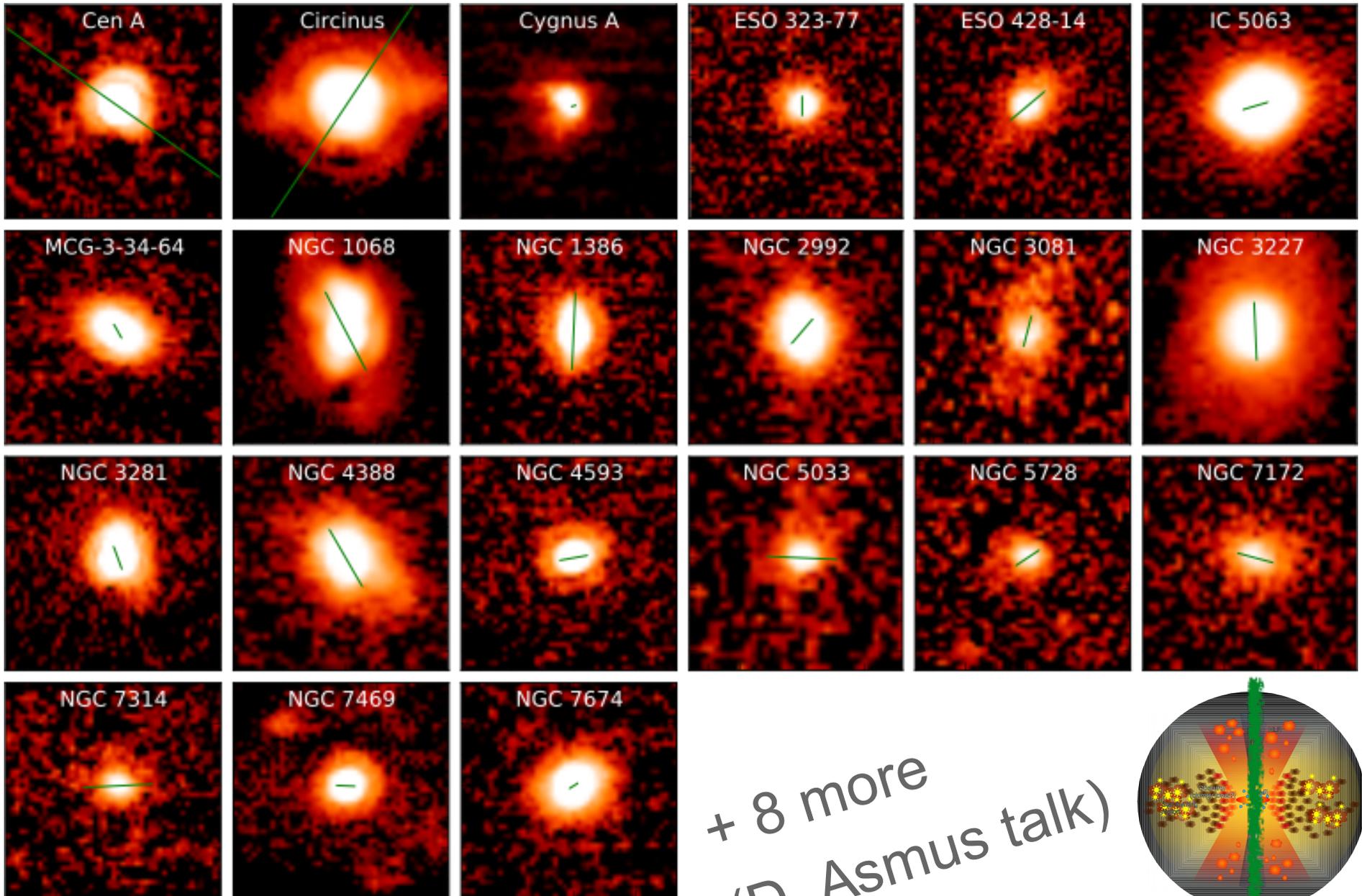
Circinus;
Tristram et al. (2014)



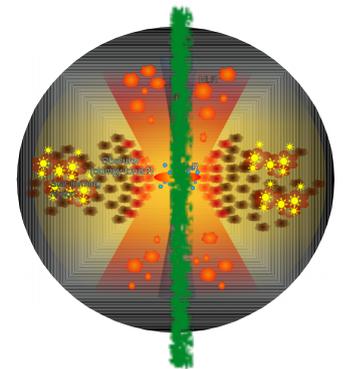
NGC 1068;
López Gonzaga et al. (2014)

MIR polar elongation on 10s-100s pc scale

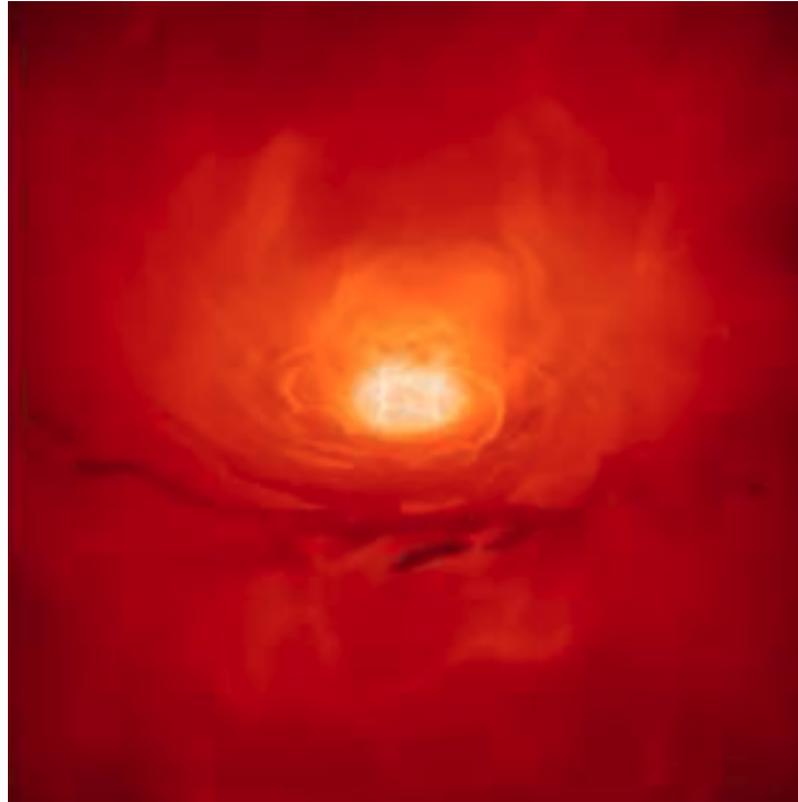
Asmus, Honig, Gandhi (2016)



+ 8 more
(D. Asmus talk)

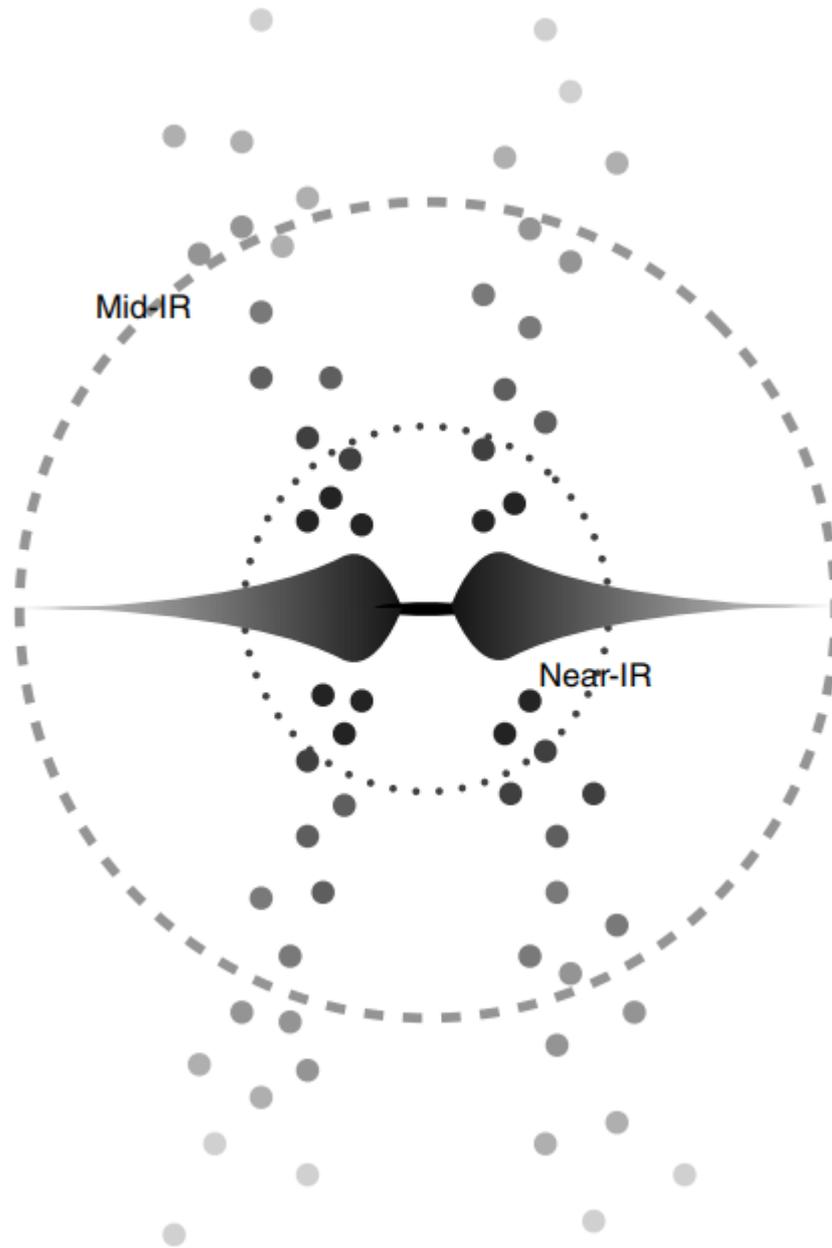


Polar elongation – inclined torus ?



Schartmann et al. (2014)

Polar elongation – dusty wind ?

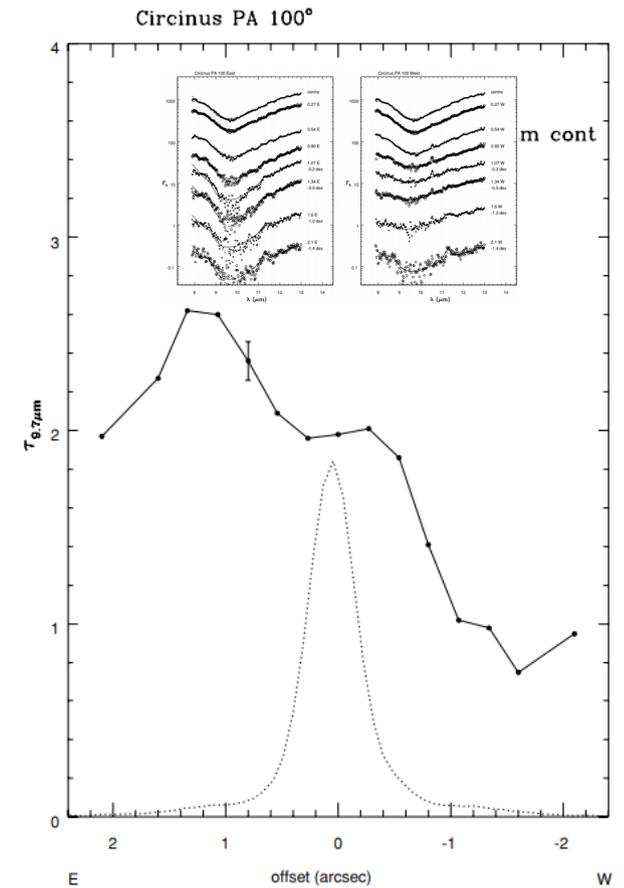
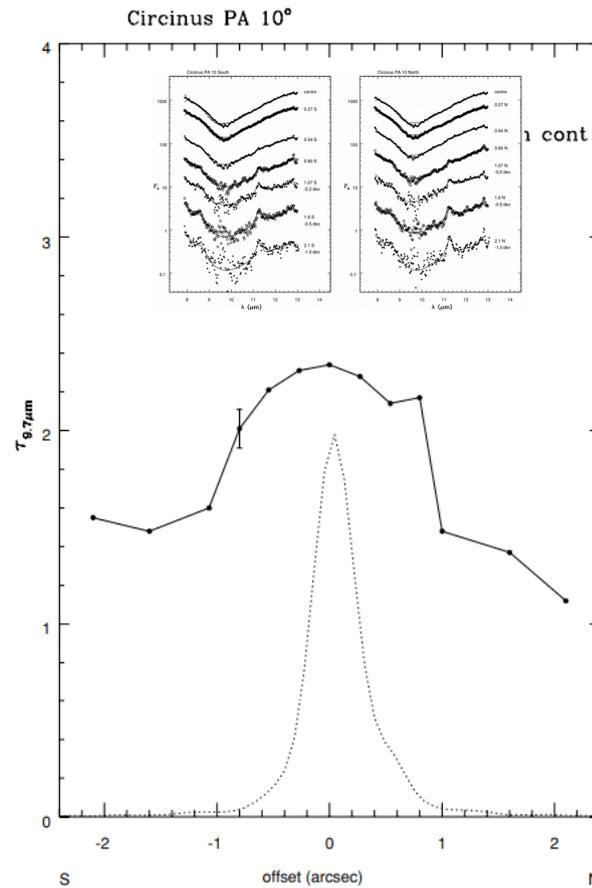
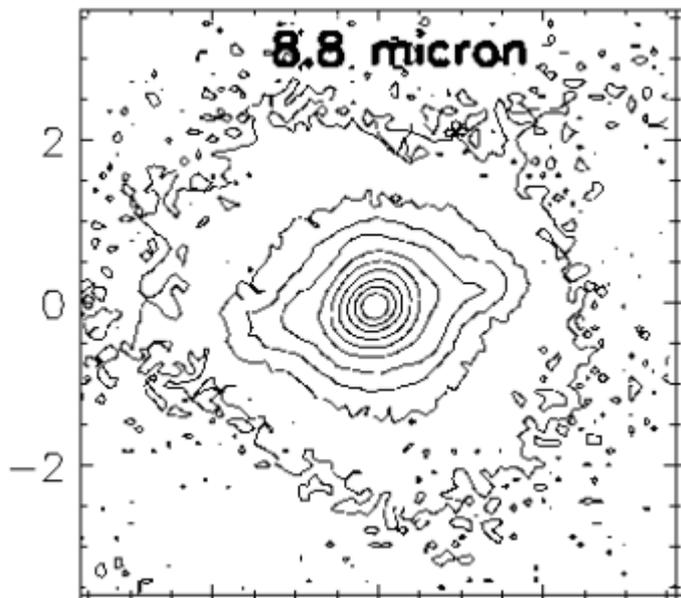


Honig et al. (2012)

Circinus extended emission and host galaxy extinction

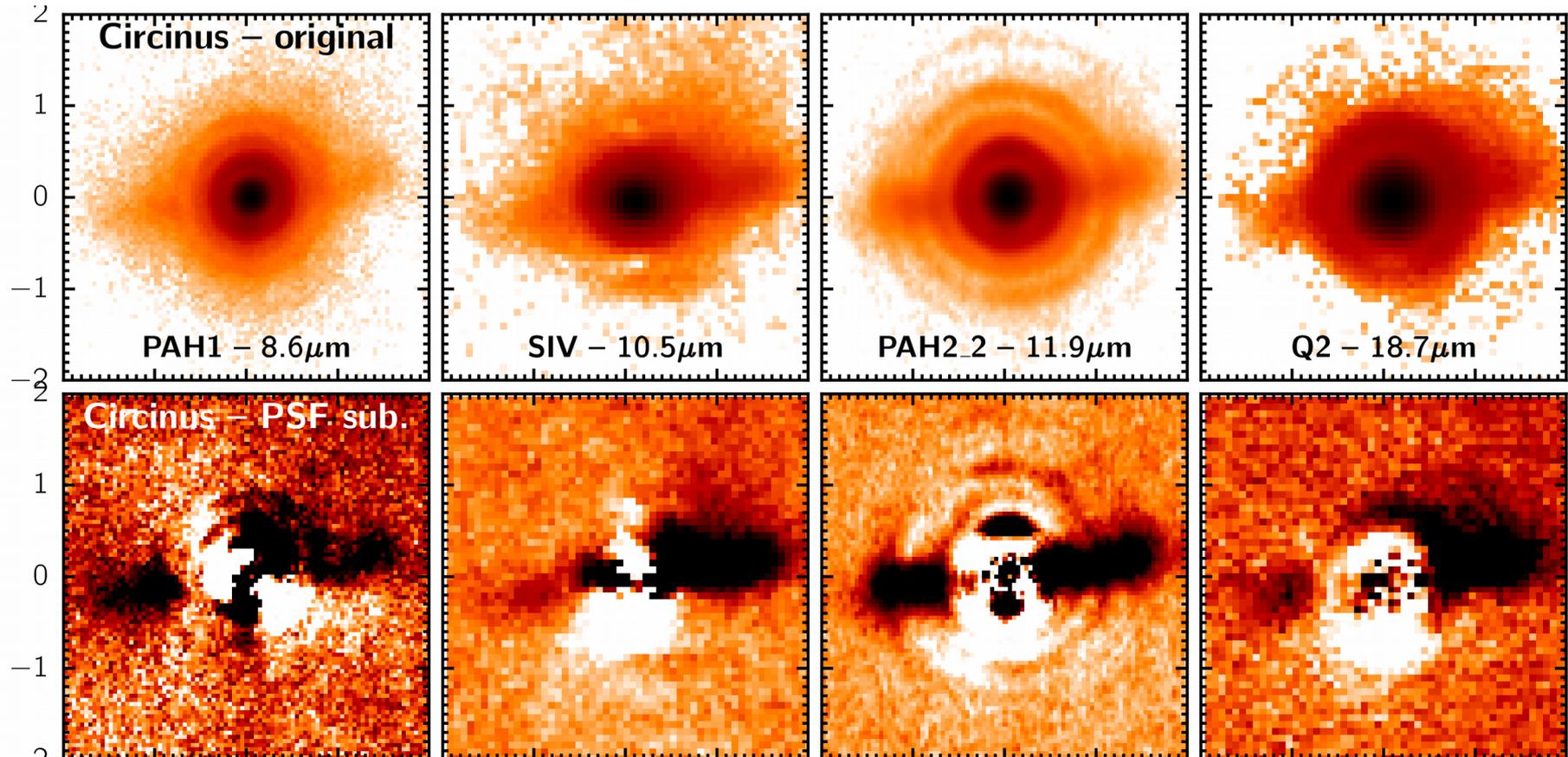
Roche et al. 2006

Packham et al. 2005



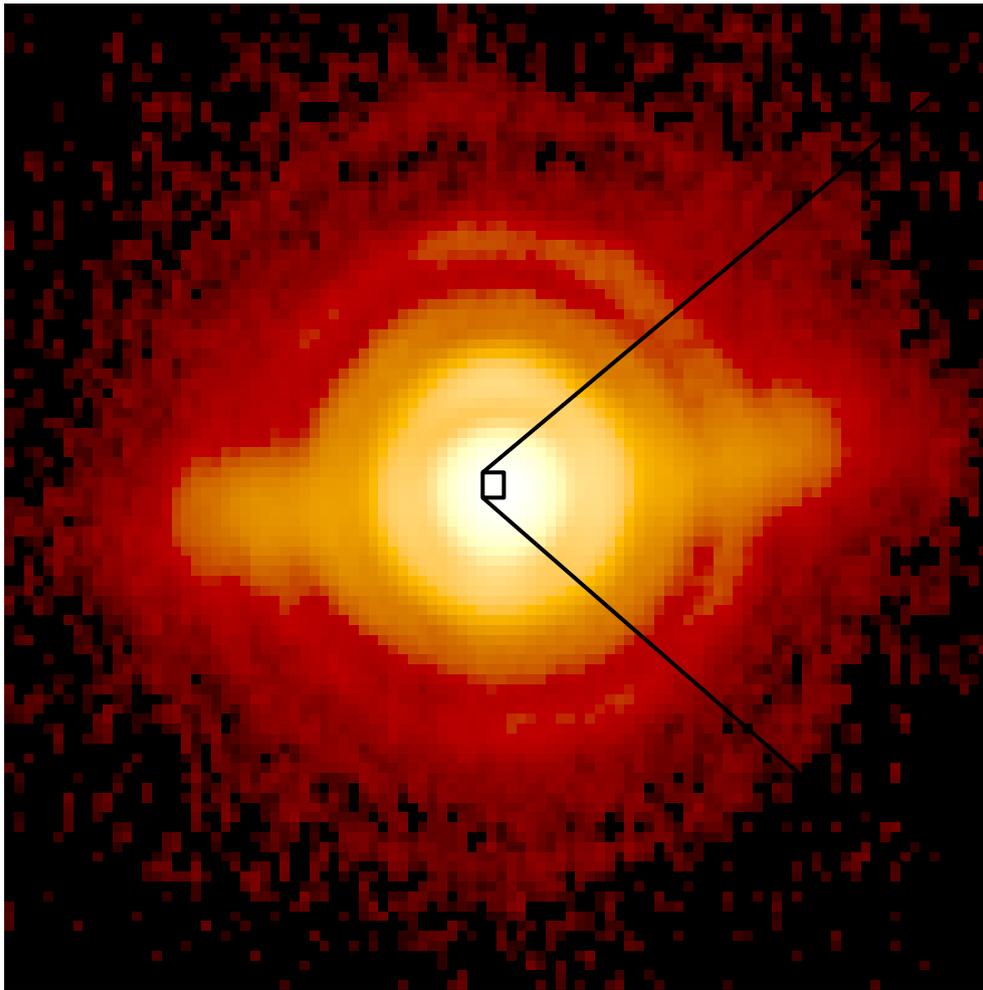
(Gemini South / T-Recs)

Circinus with VLT/VISIR

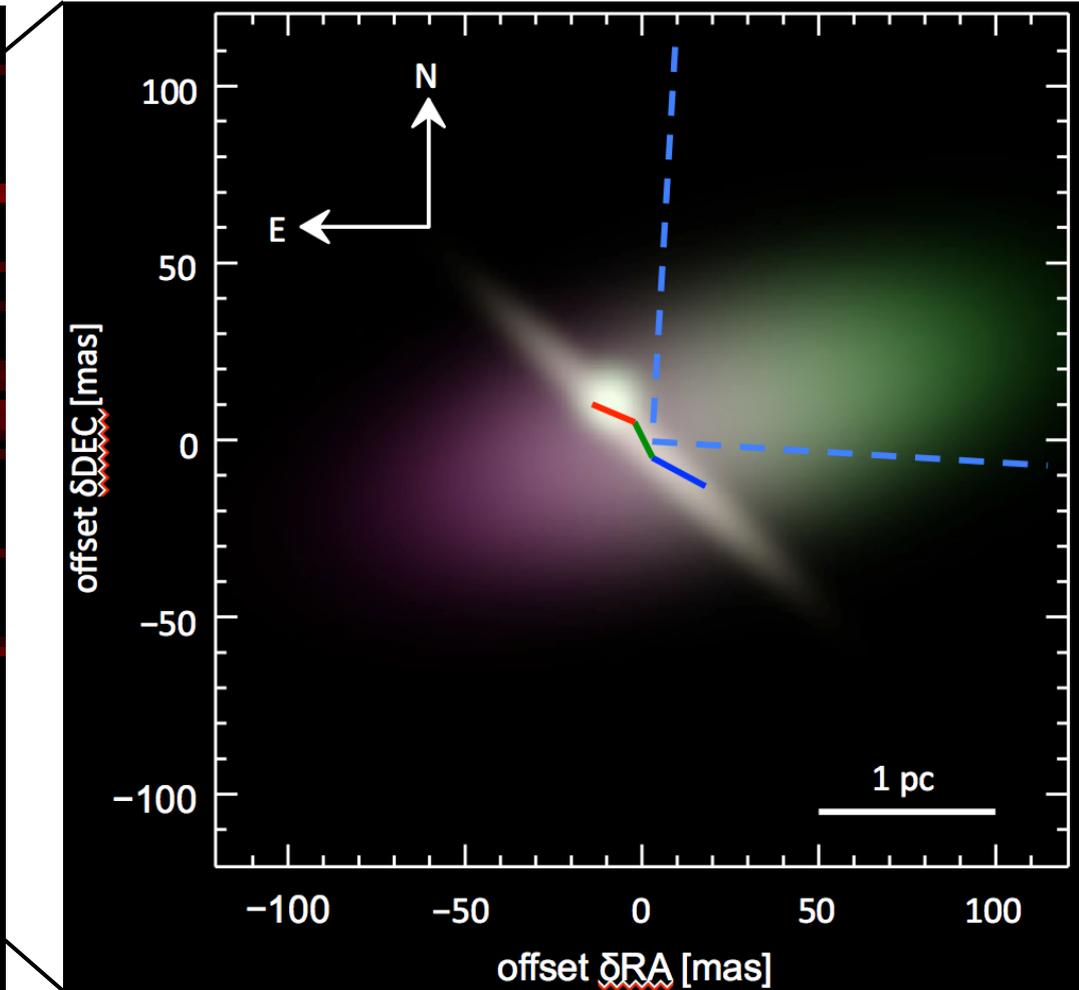


Stalevski, Asmus & Tristram (2017)

Circinus MIR emission

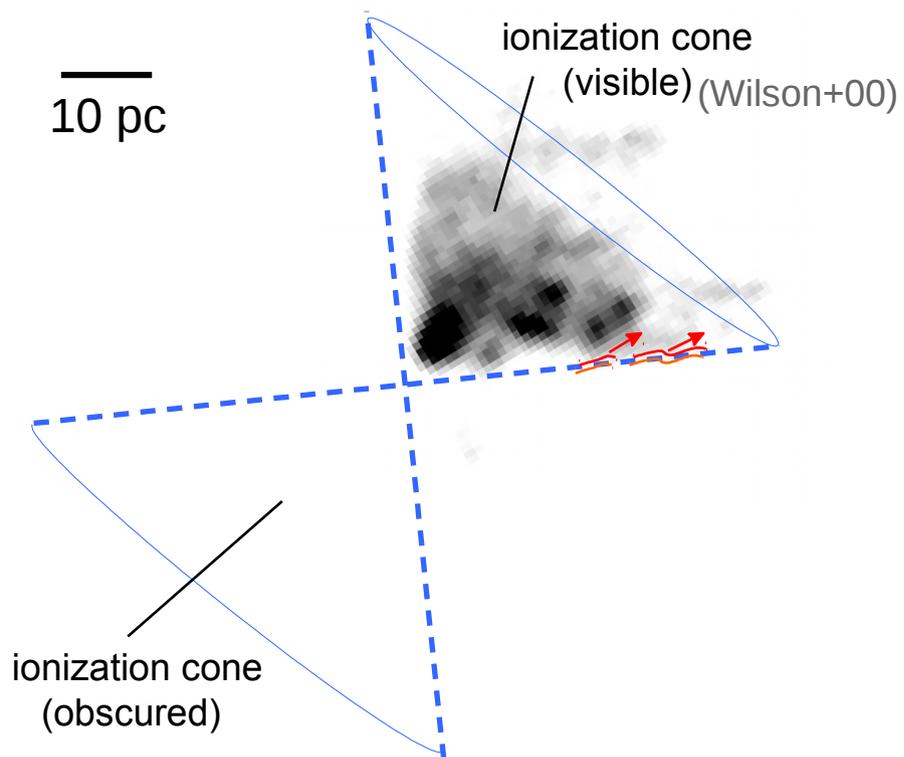
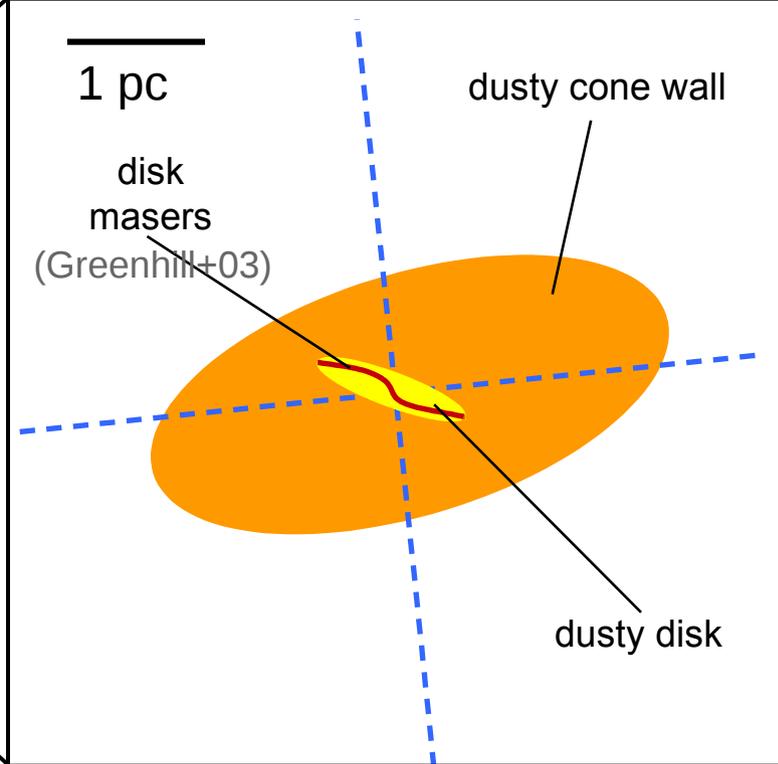
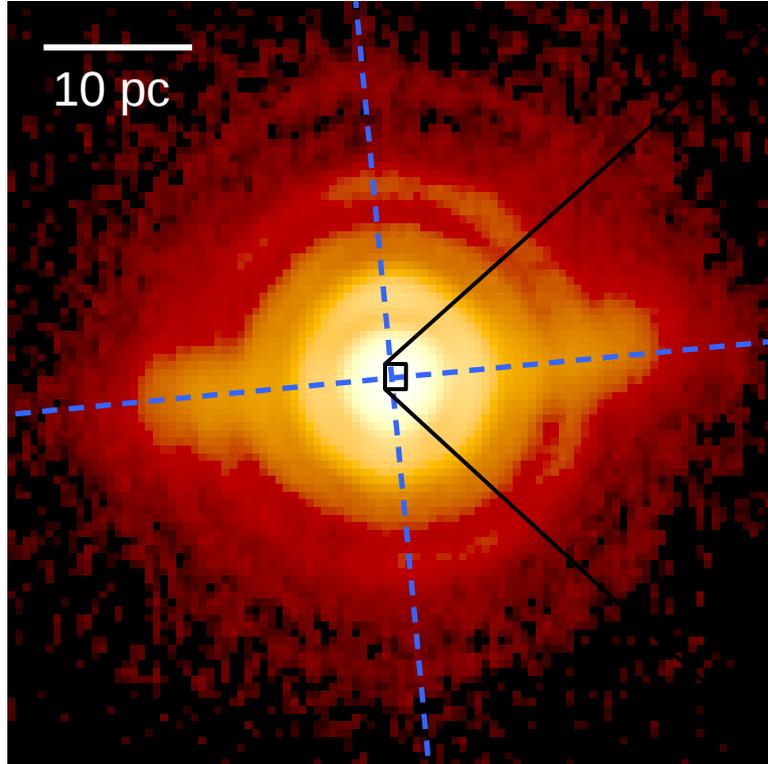


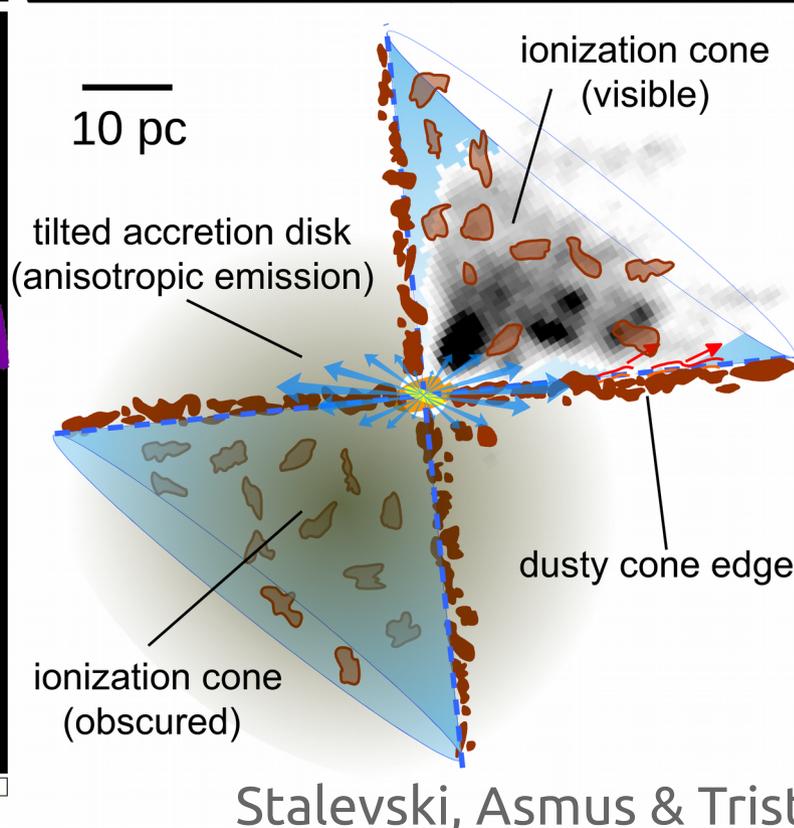
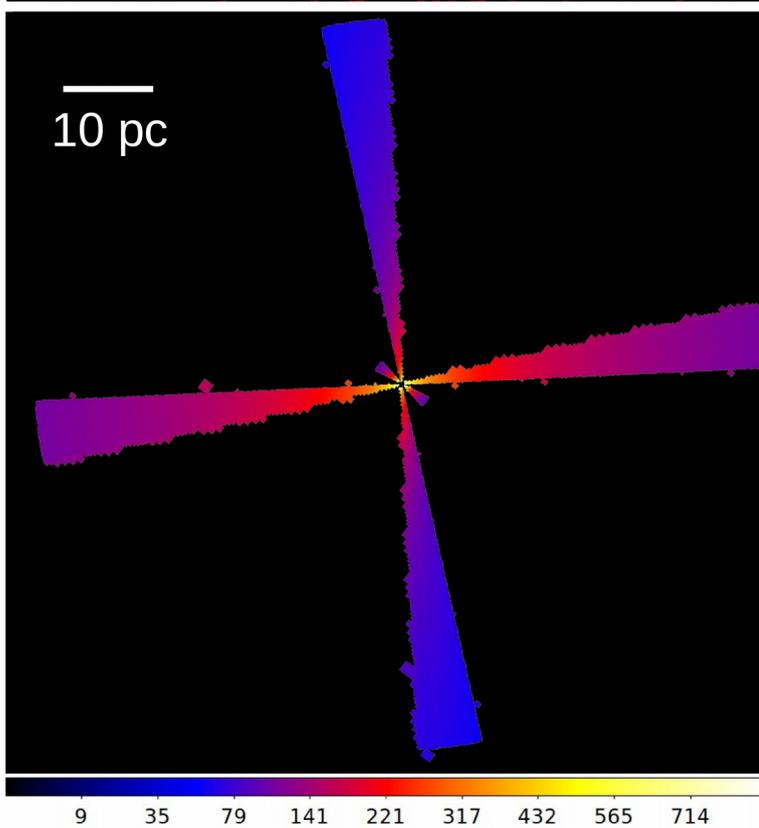
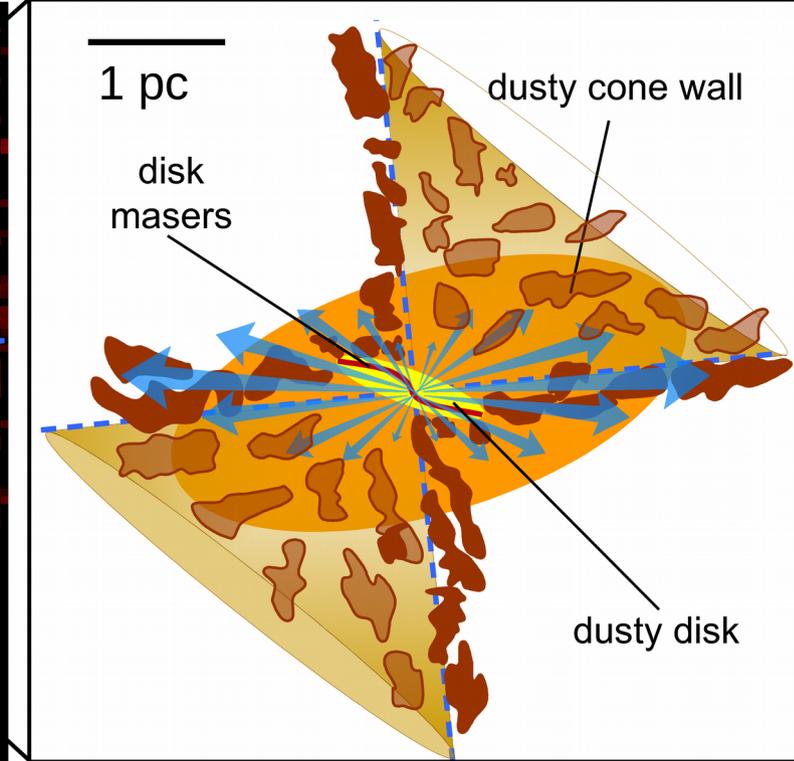
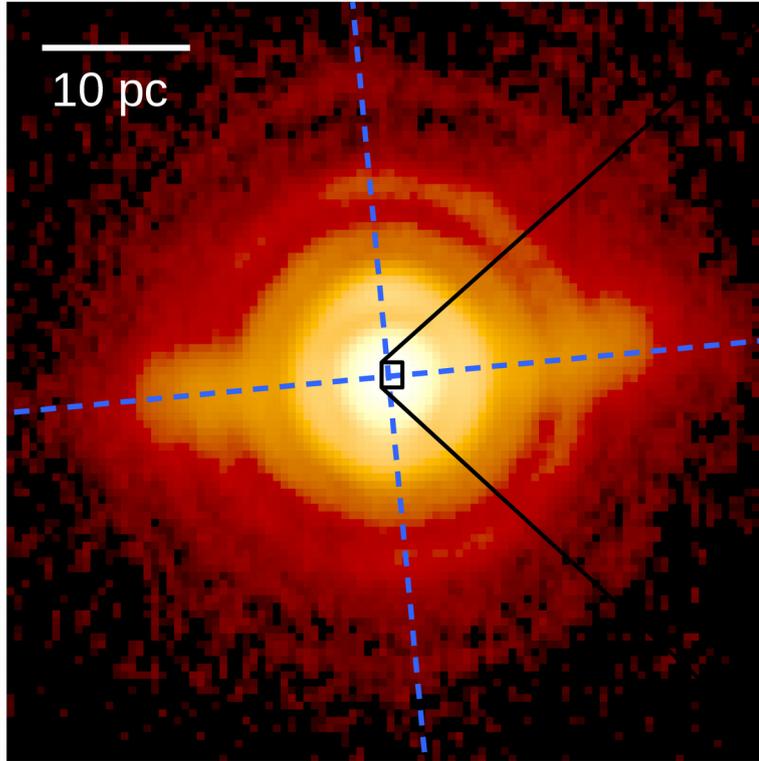
Stalevski, Asmus & Tristram (2017)



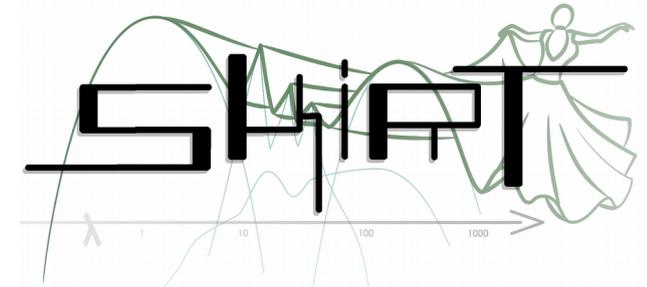
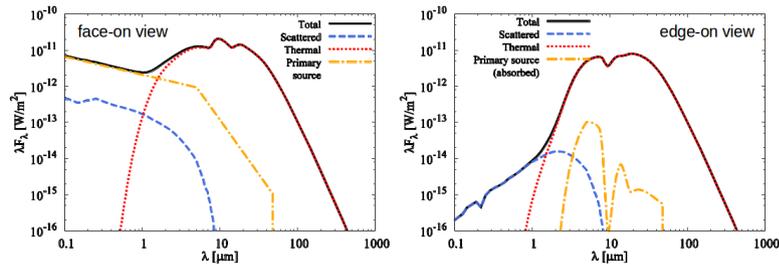
Tristram et al. (2014)

A prototype that does not look like a prototype!





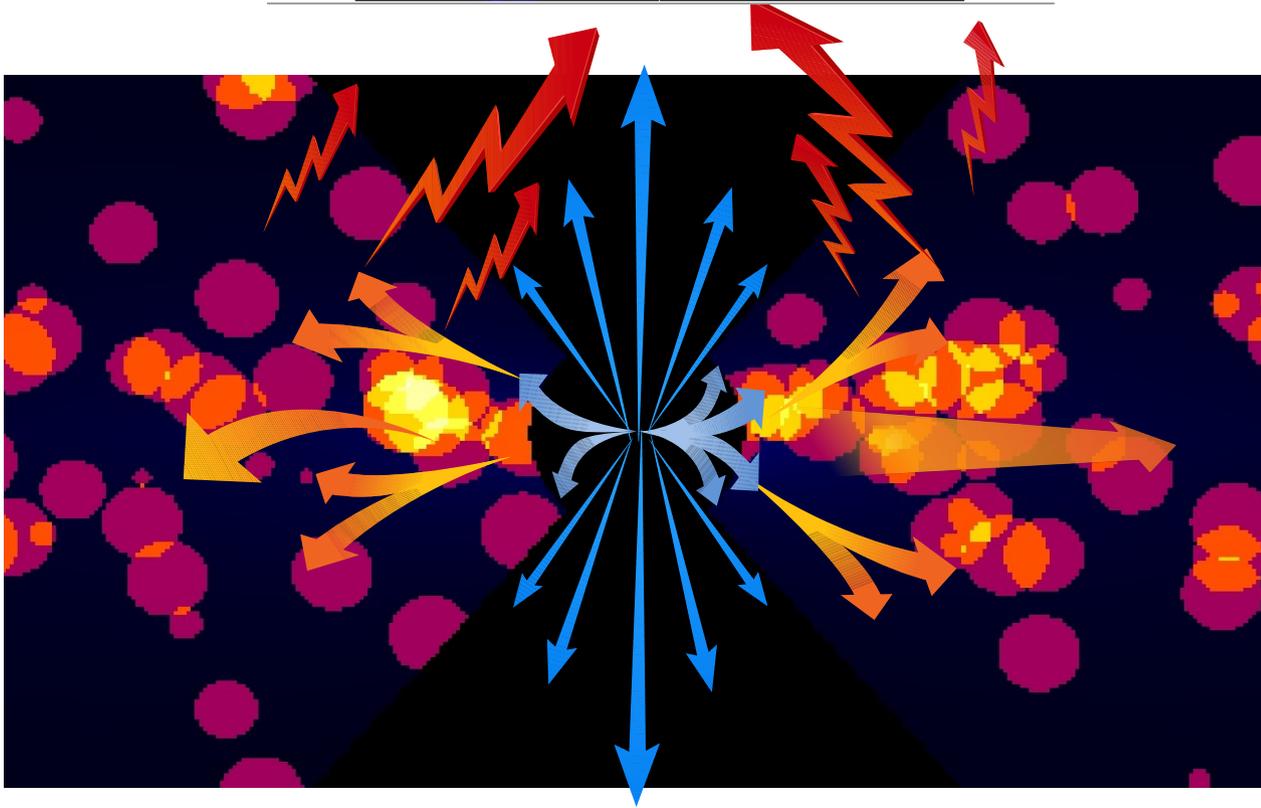
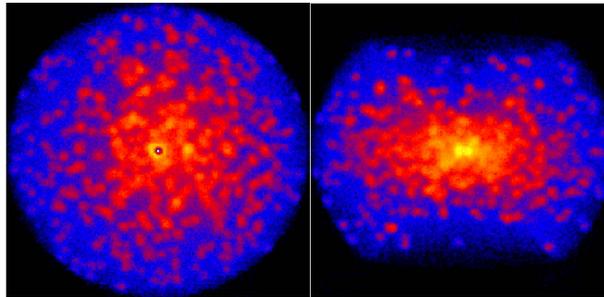
Monte Carlo radiative transfer



<http://www.skirt.ugent.be>

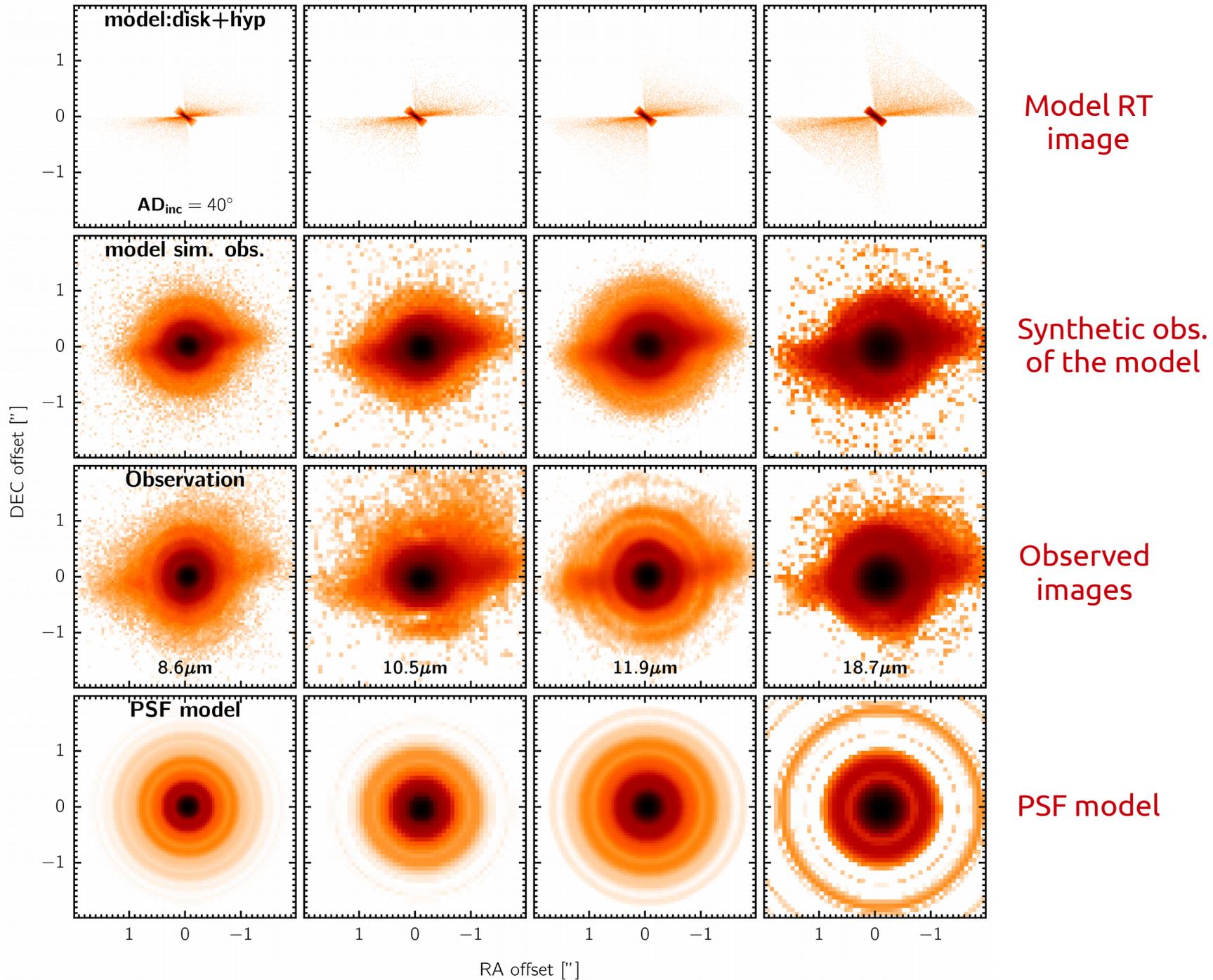
Baes+ 2011
Baes & Camps 2015
Camps & Baes 2015

Stalevski+ 2012, 2016



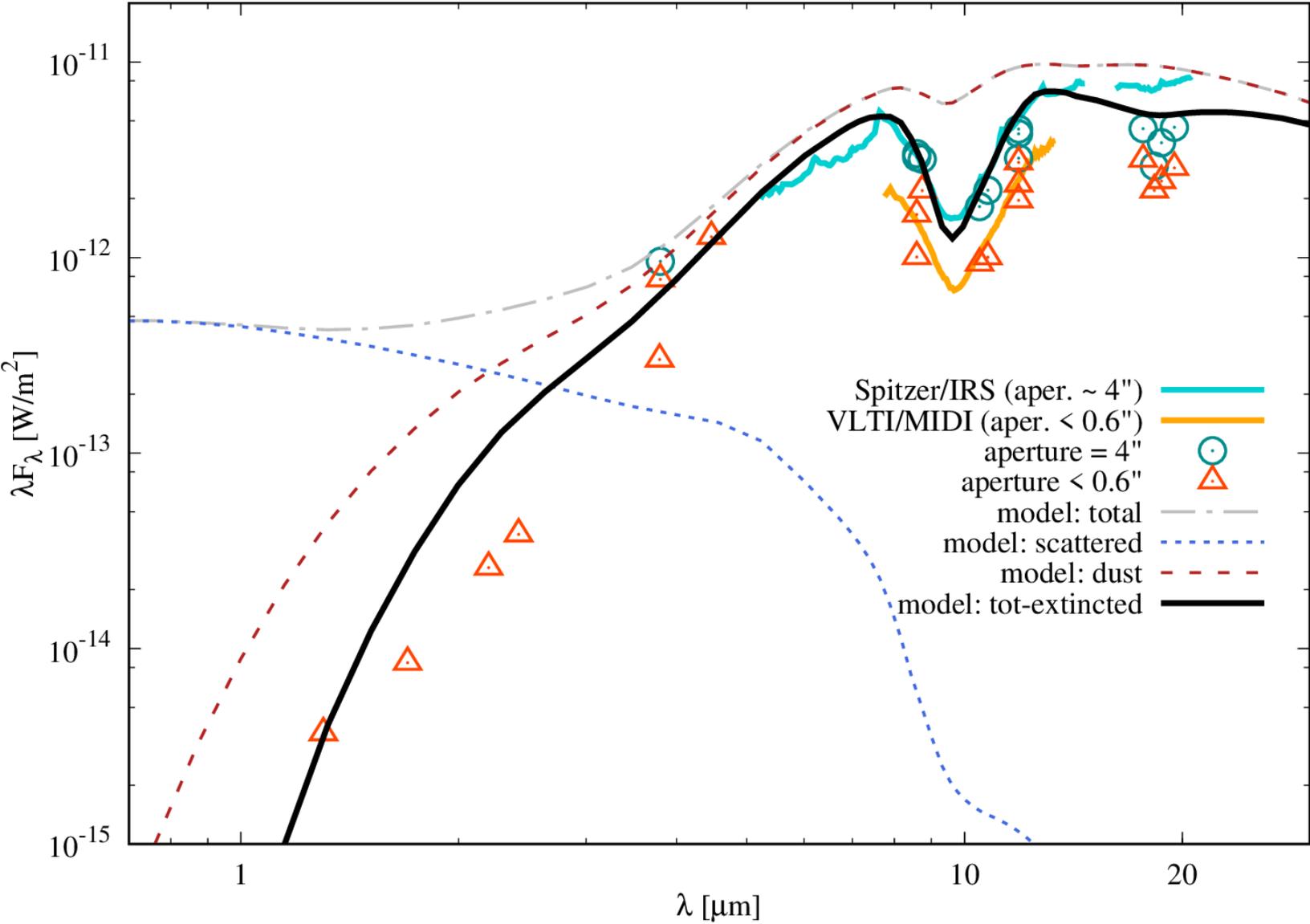
Good match with the MIR morphology

Stalevski, Asmus & Tristram (2017)



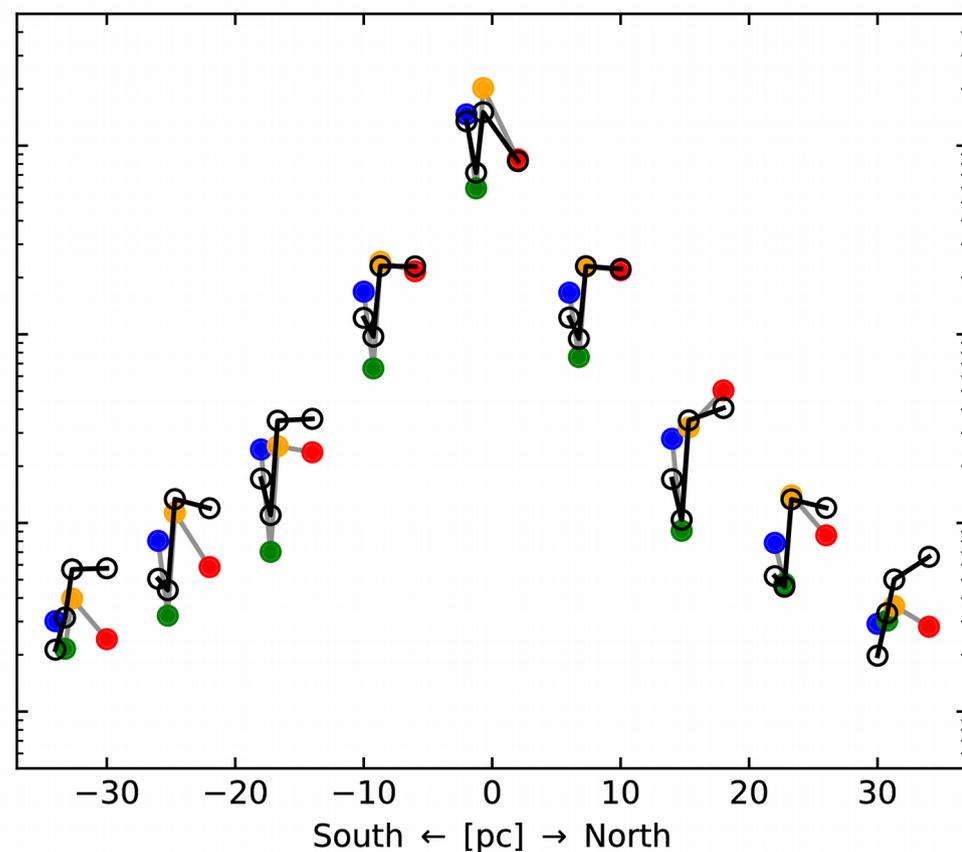
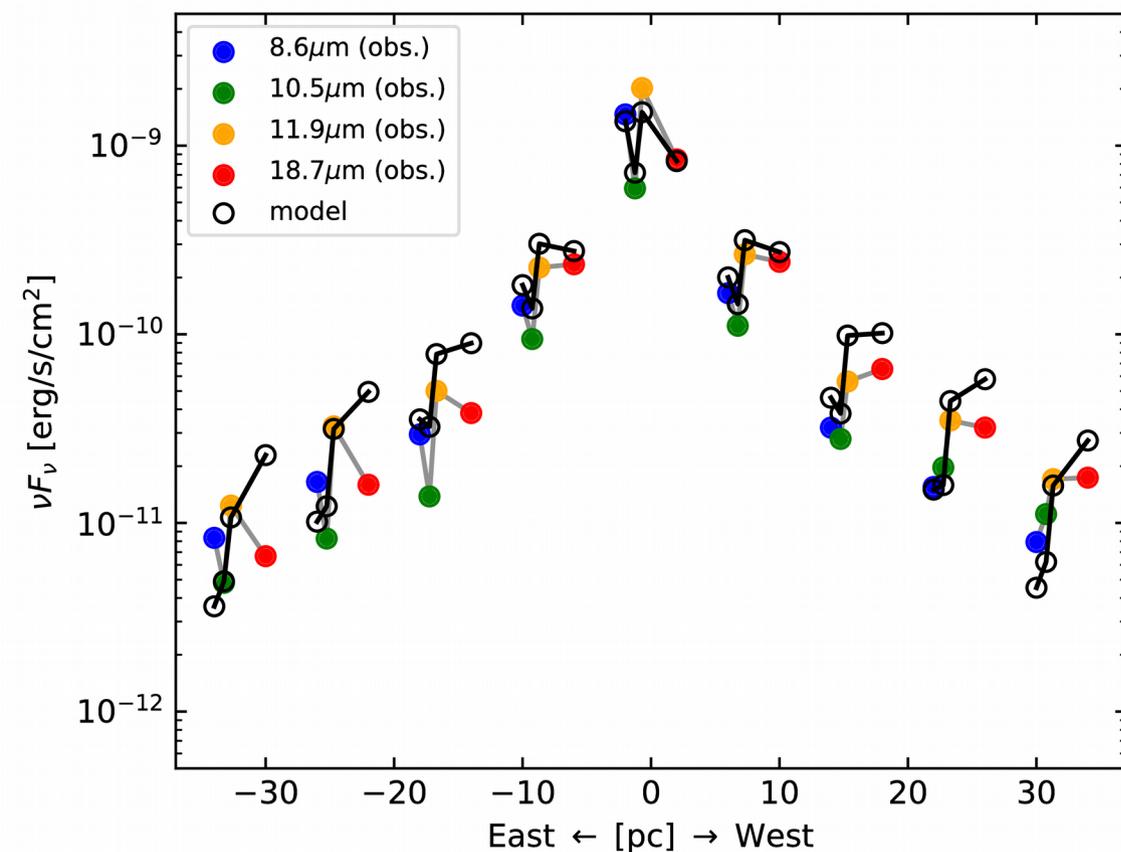
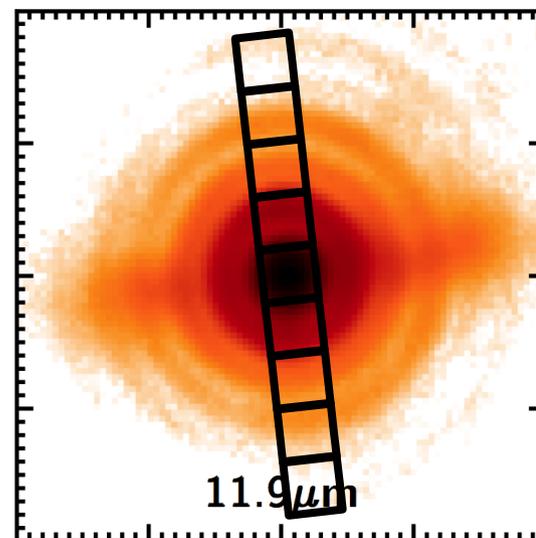
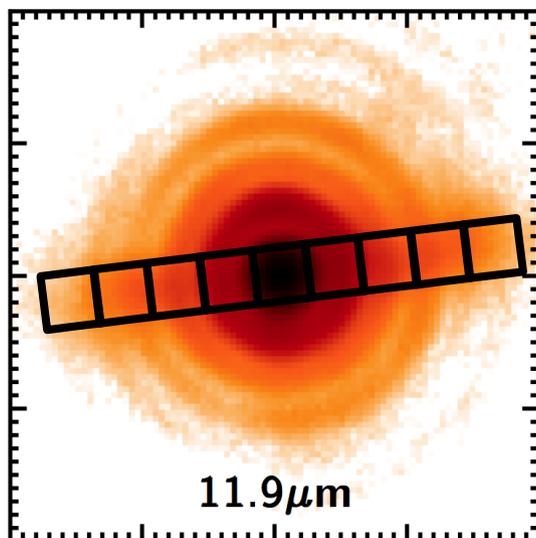
Good match with the SED

Stalevski, Asmus & Tristram (2017)



Good match with the resolved photometry

Stalevski, Asmus & Tristram (2017)

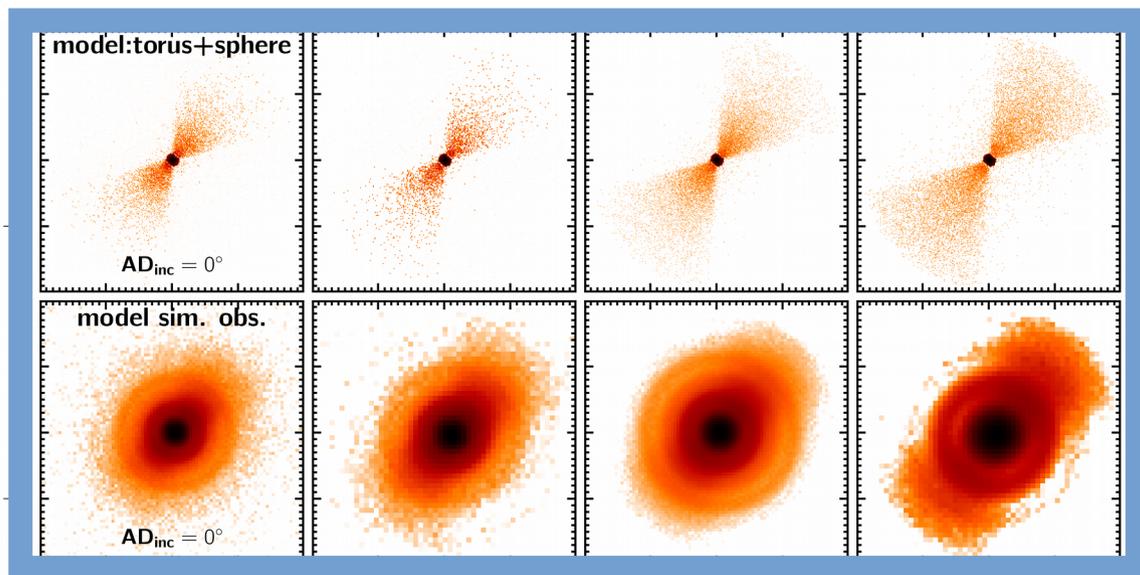


torus+sphere shell: does not work

Monte Carlo radiative transfer with



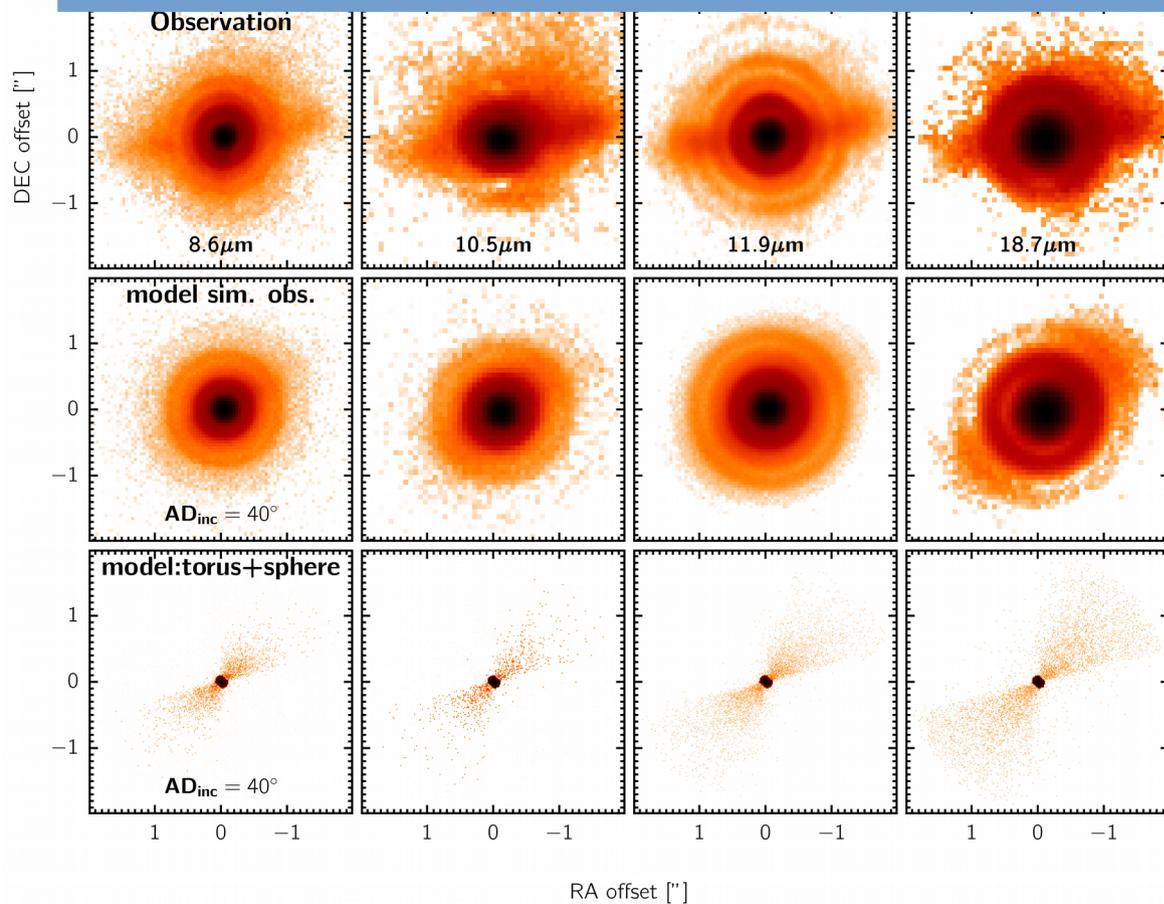
<http://www.skirt.ugent.be>



Model RT image

Aligned AD

Synthetic obs. of the model

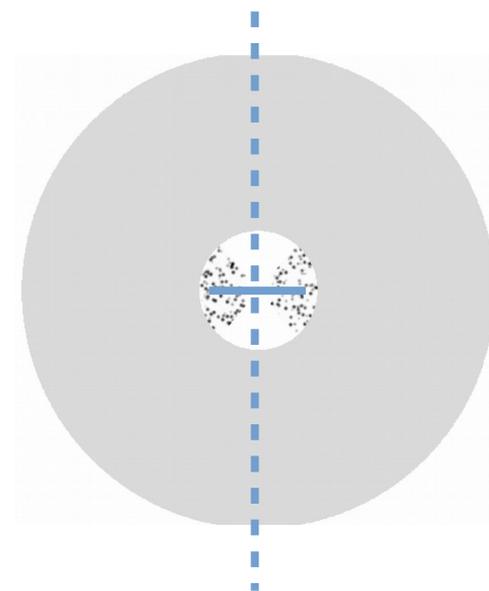


Observed images

Synthetic obs. of the model

Model RT image

Tilted AD

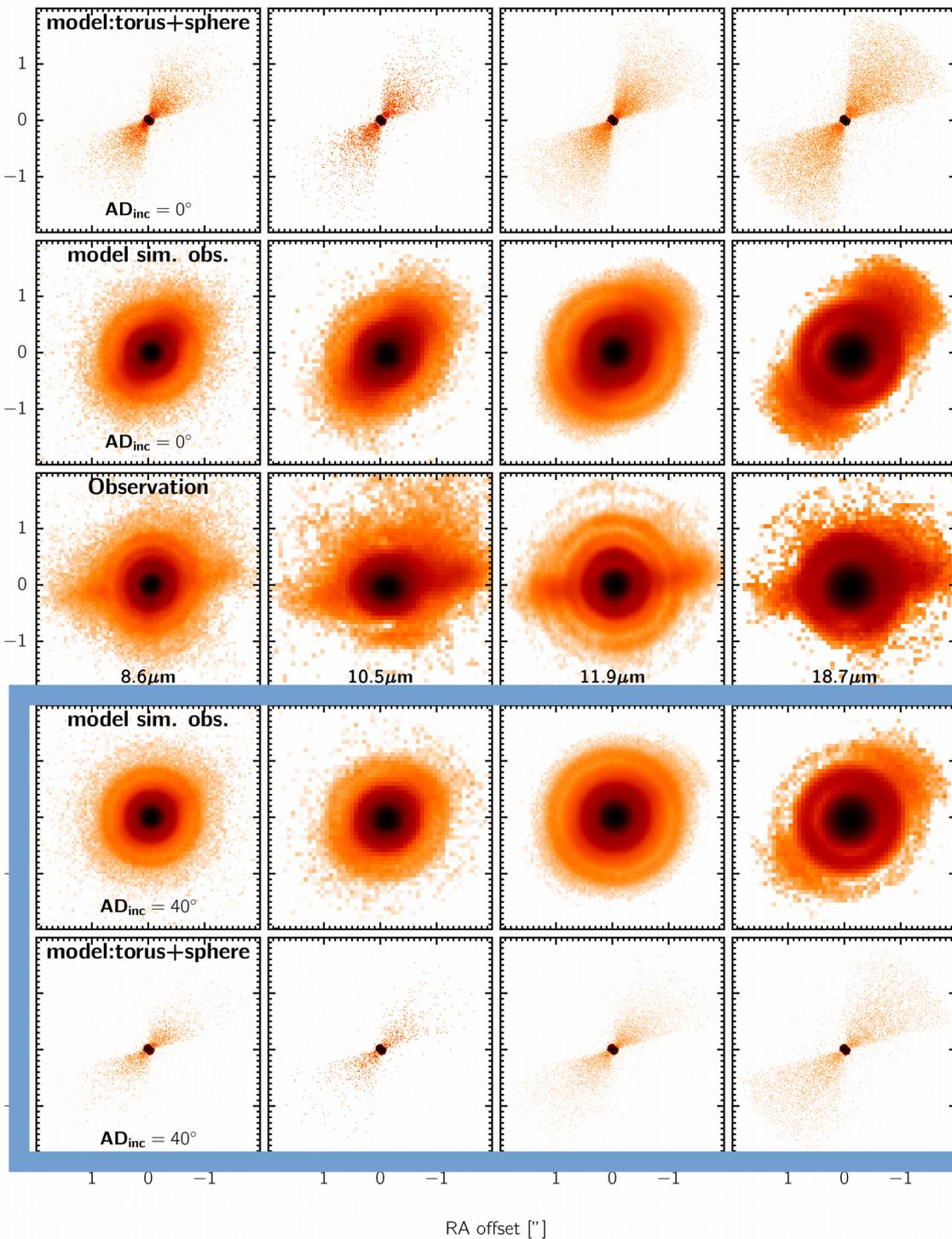


torus+sphere shell: does not work

Monte Carlo radiative transfer with



<http://www.skirt.ugent.be>



Model RT image

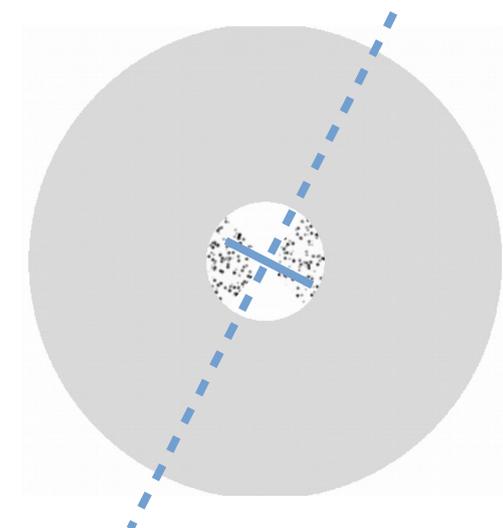
Synthetic obs. of the model

Observed images

Synthetic obs. of the model

Model RT image

Aligned AD



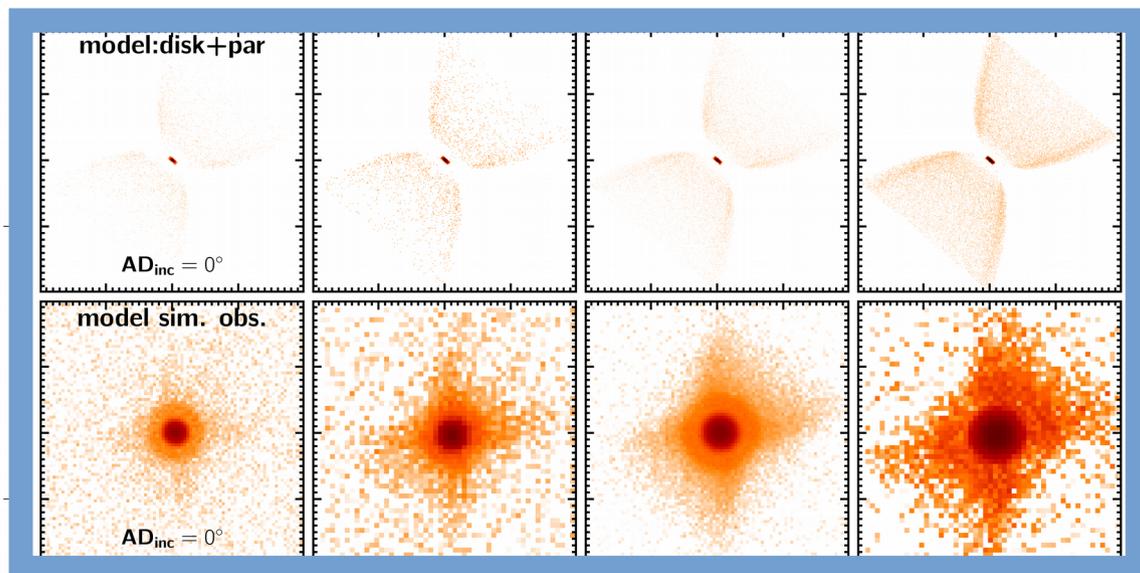
Tilted AD

disk+paraboloid: does not work

Monte Carlo radiative transfer with



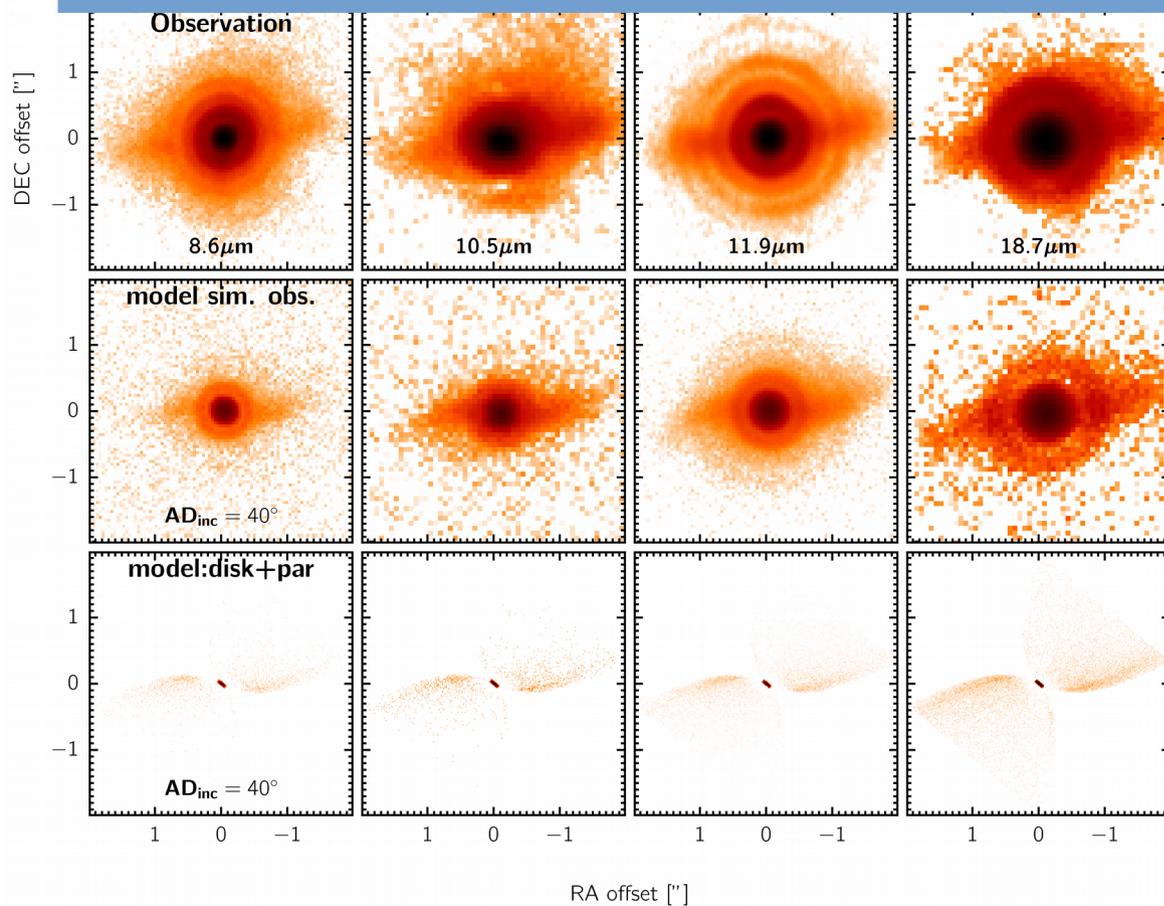
<http://www.skirt.ugent.be>



Model RT image

Aligned AD

Synthetic obs. of the model

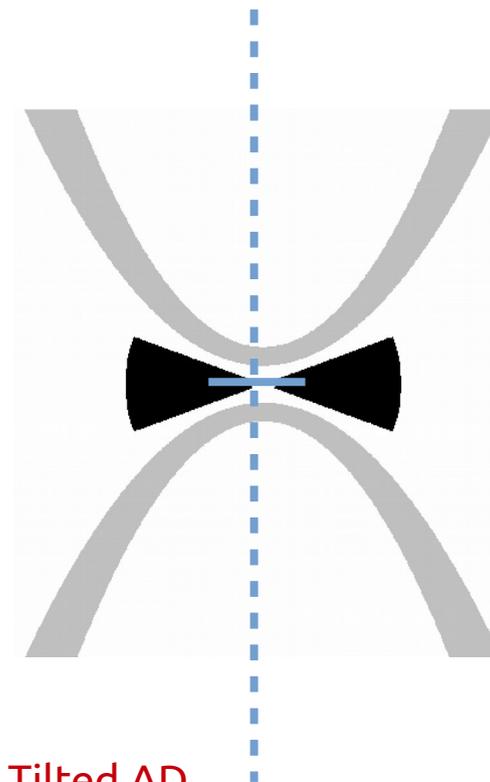


Observed images

Synthetic obs. of the model

Model RT image

Tilted AD

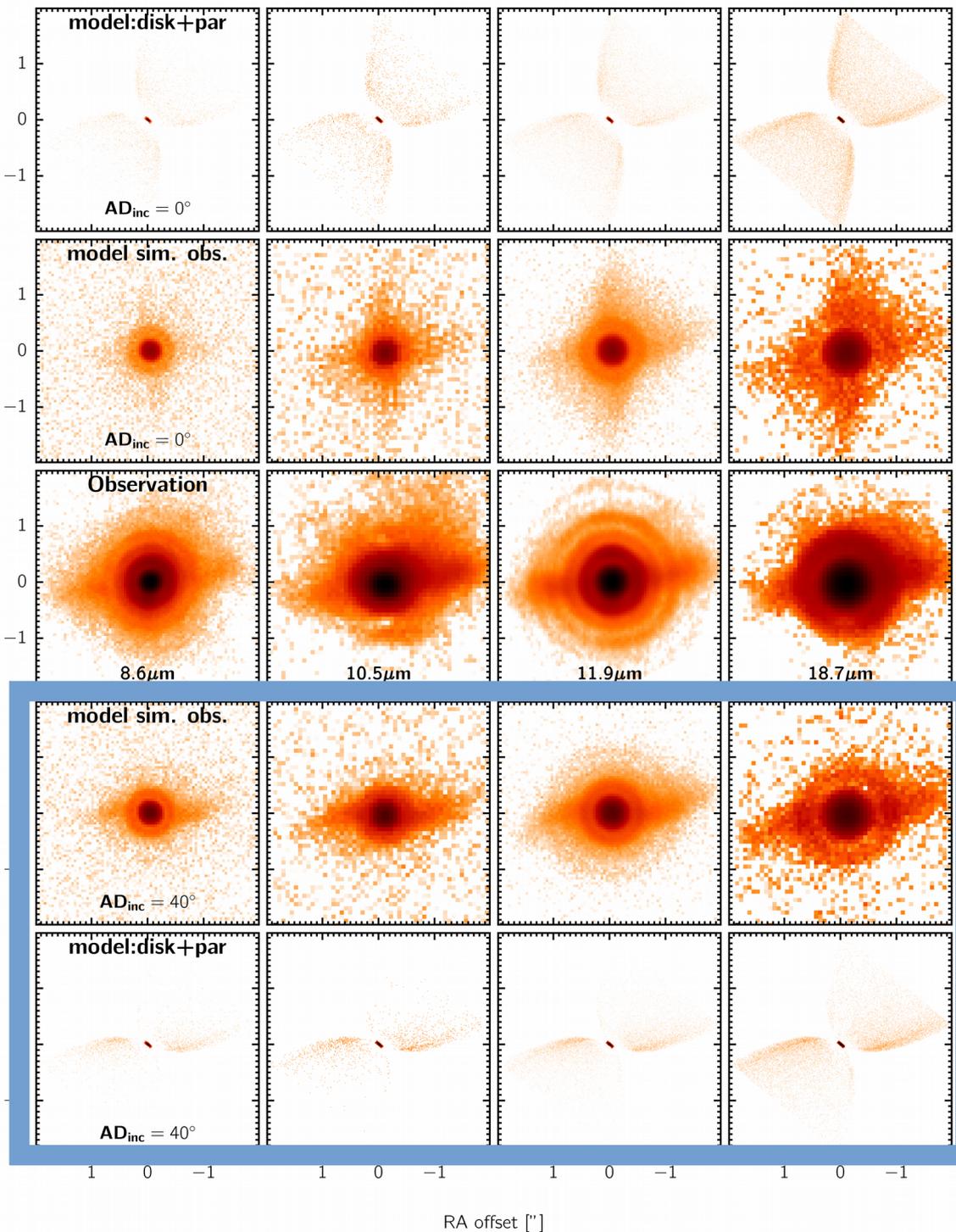


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Model RT image

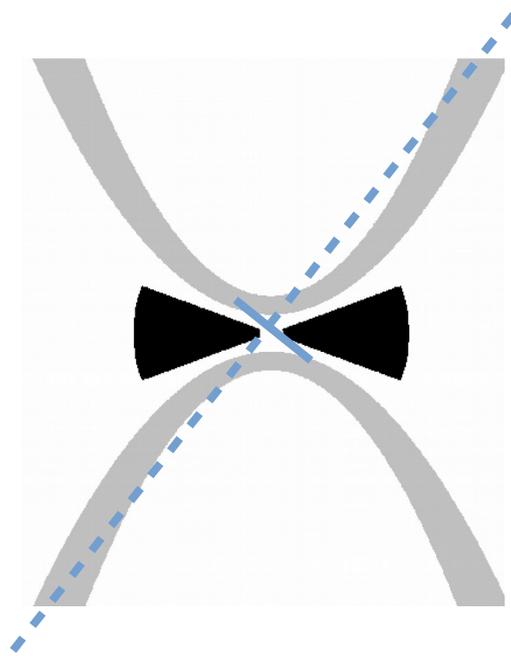
Synthetic obs. of the model

Observed images

Synthetic obs. of the model

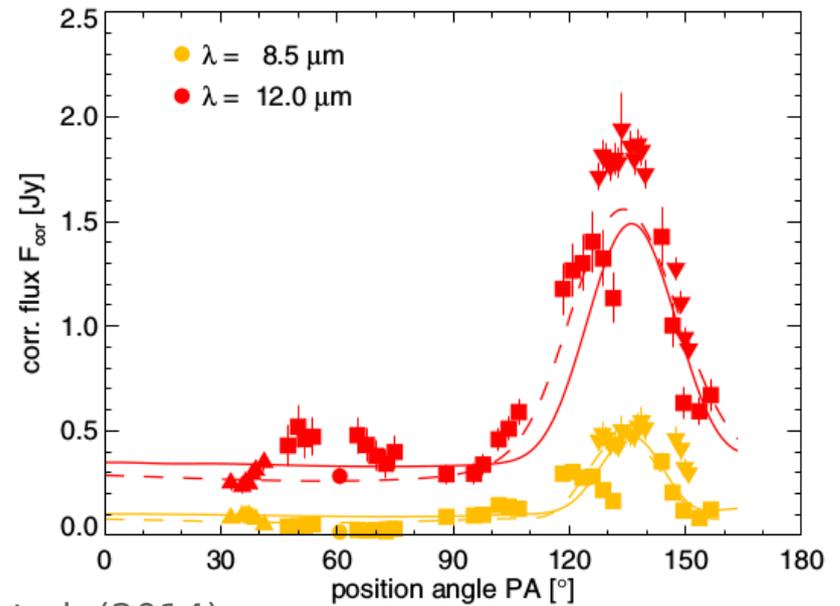
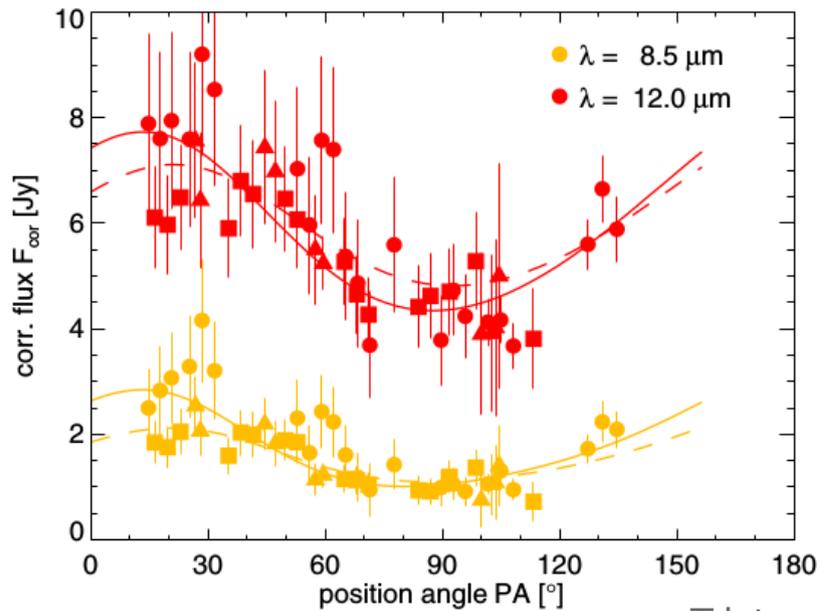
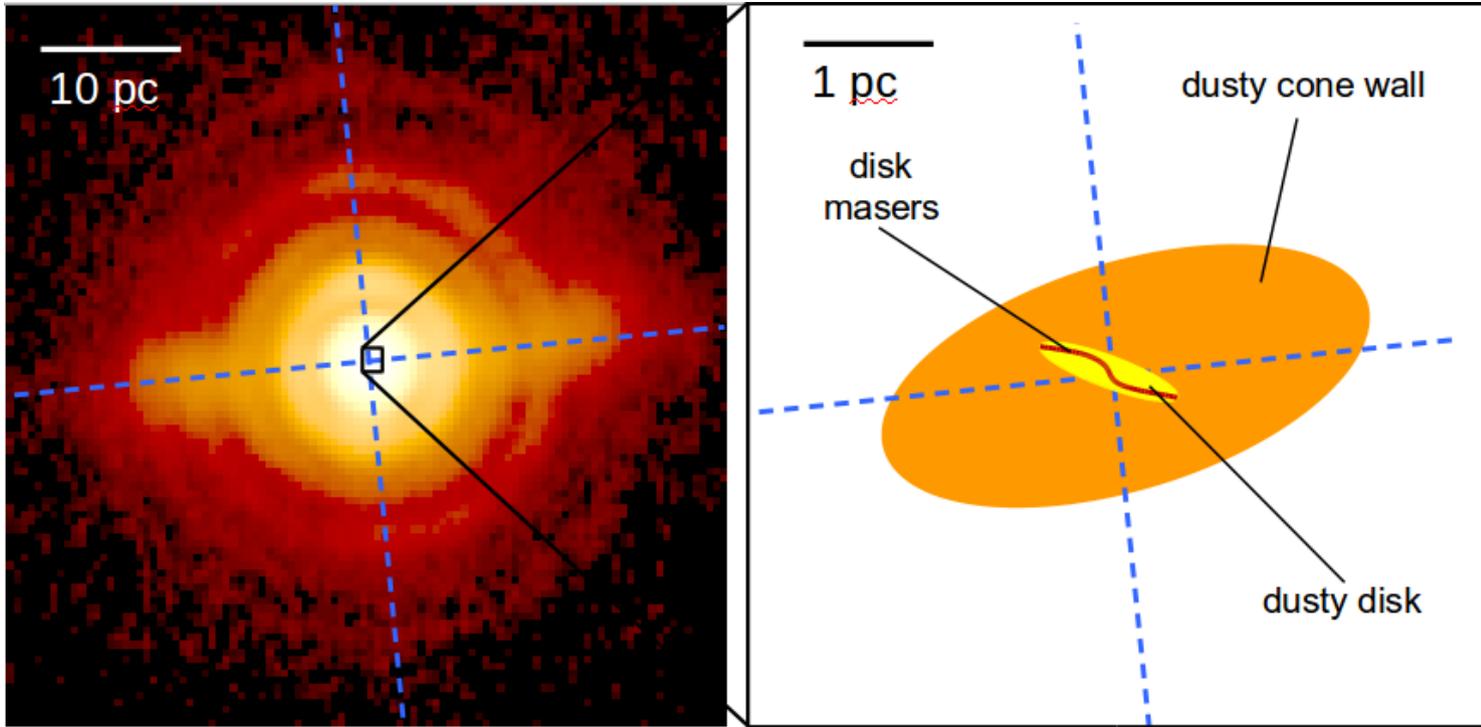
Model RT image

Aligned AD



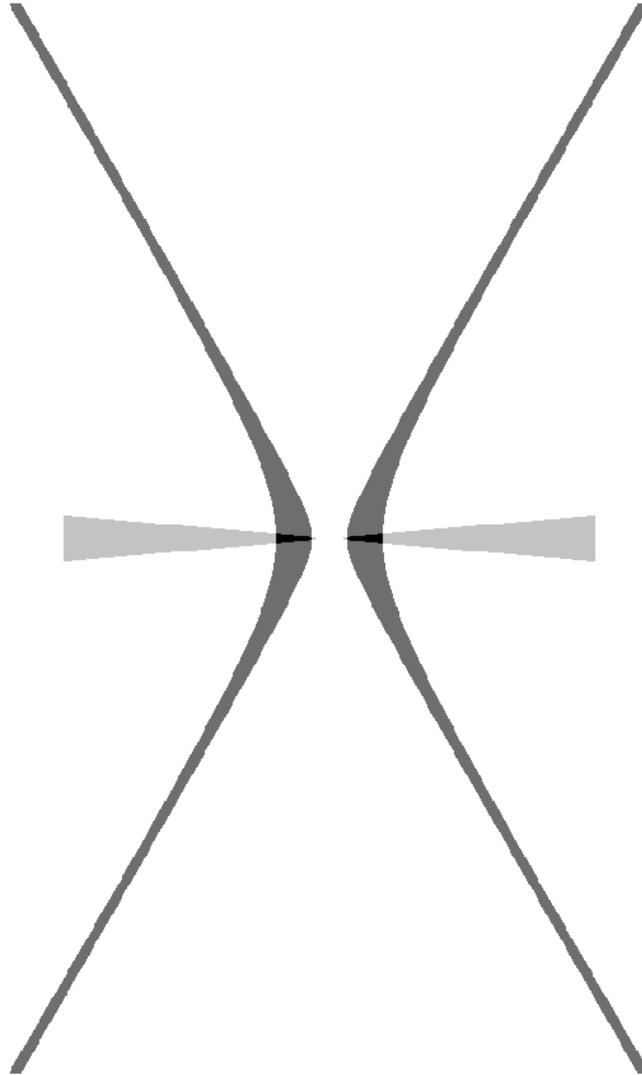
Tilted AD

Zooming in with VLT/MIDI



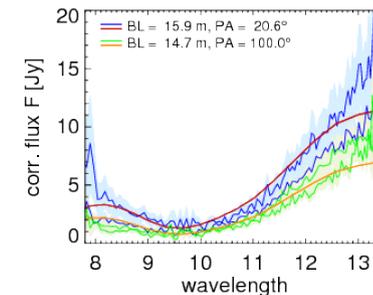
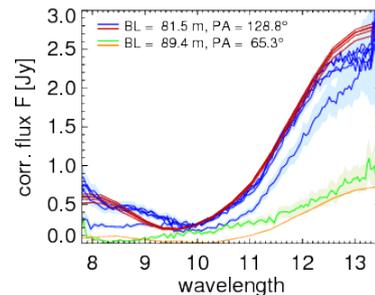
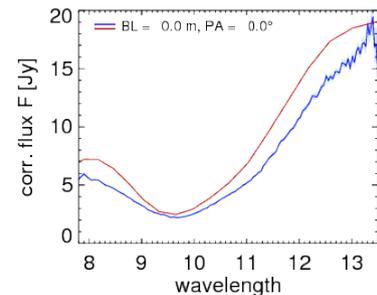
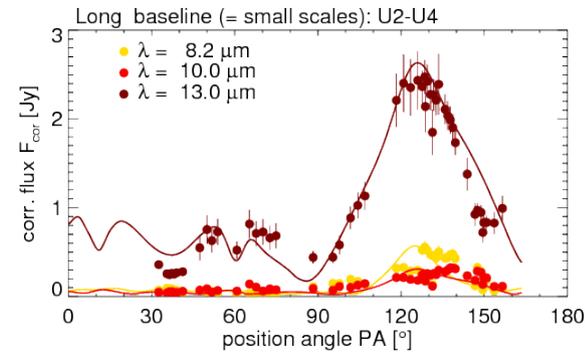
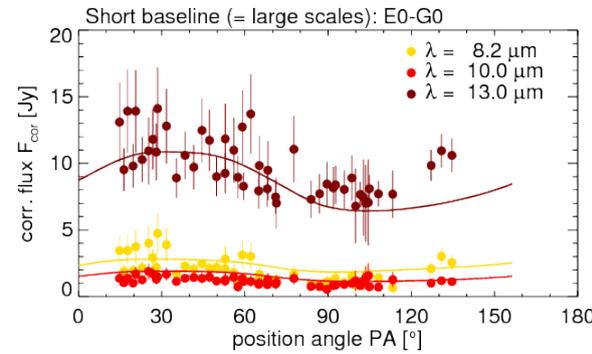
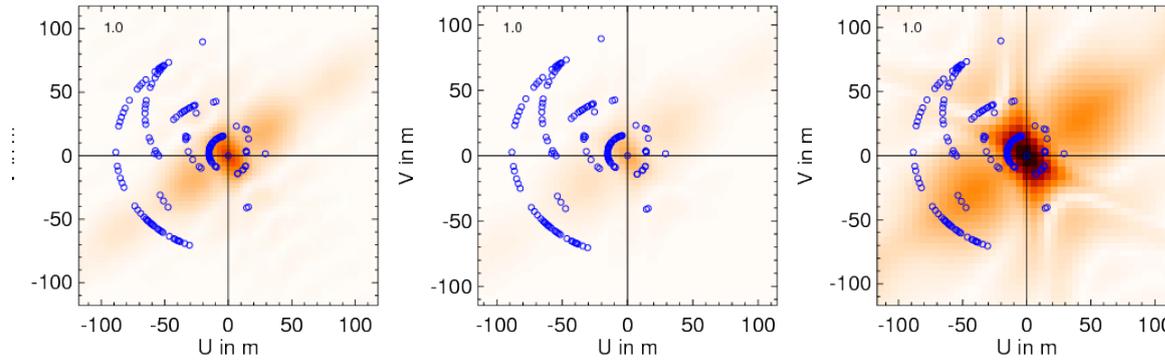
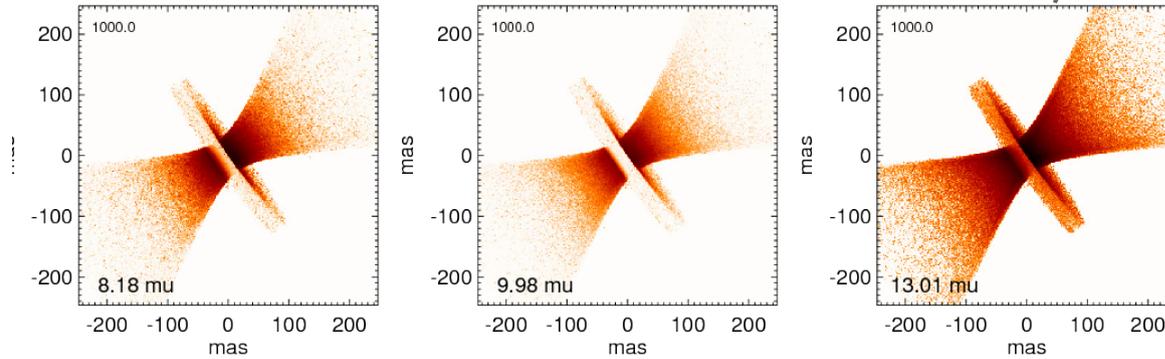
disk + hyperboloid polar wind

Stalevski, Tristram & Asmus, subm.



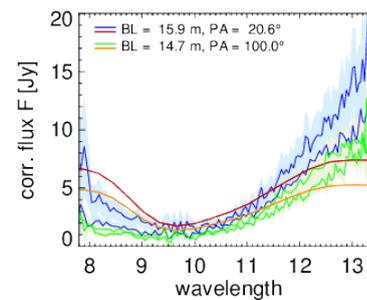
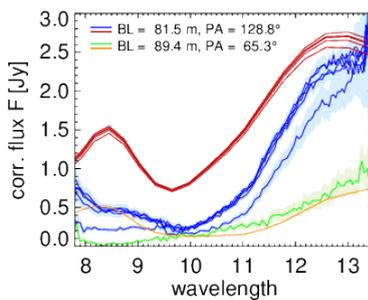
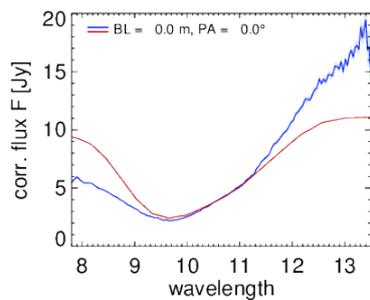
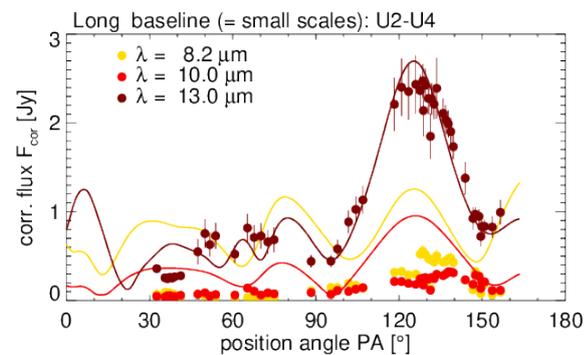
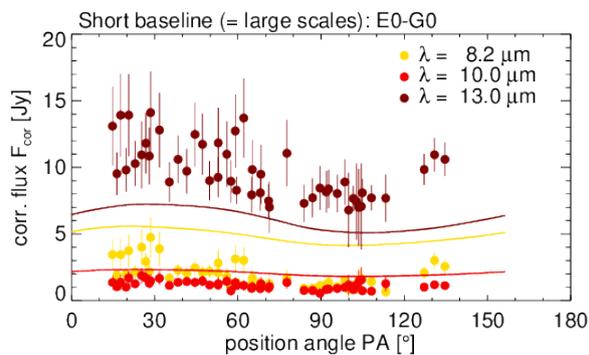
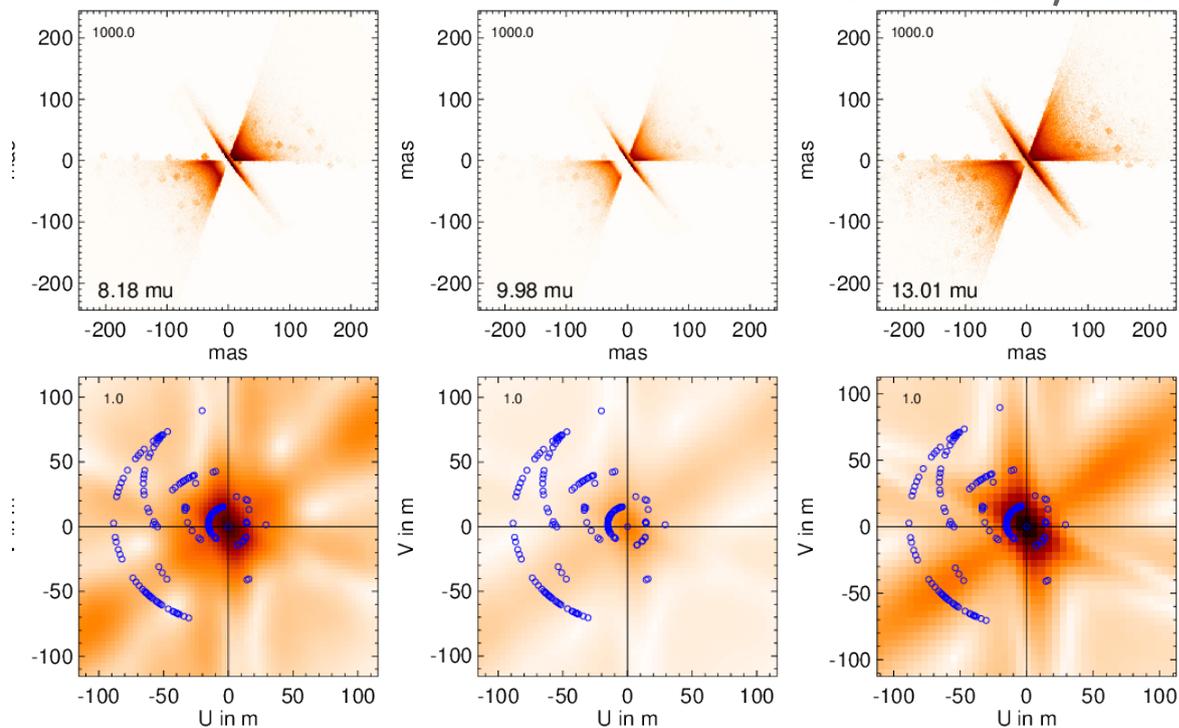
disk+hyp wind: good match with VLTI/MIDI

Stalevski, Tristram & Asmus, subm.



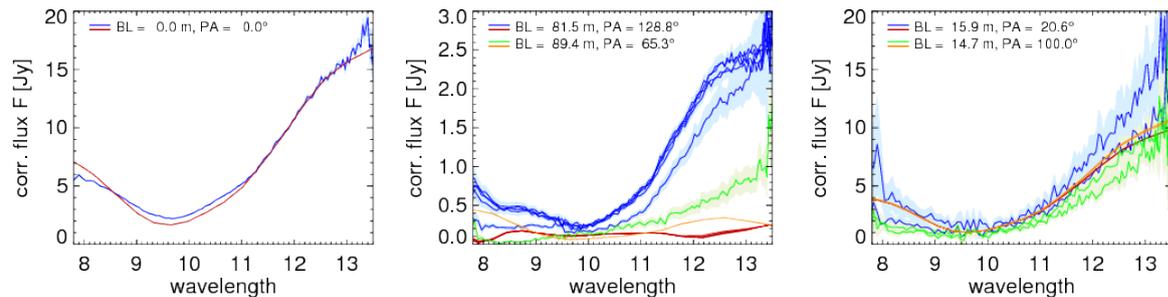
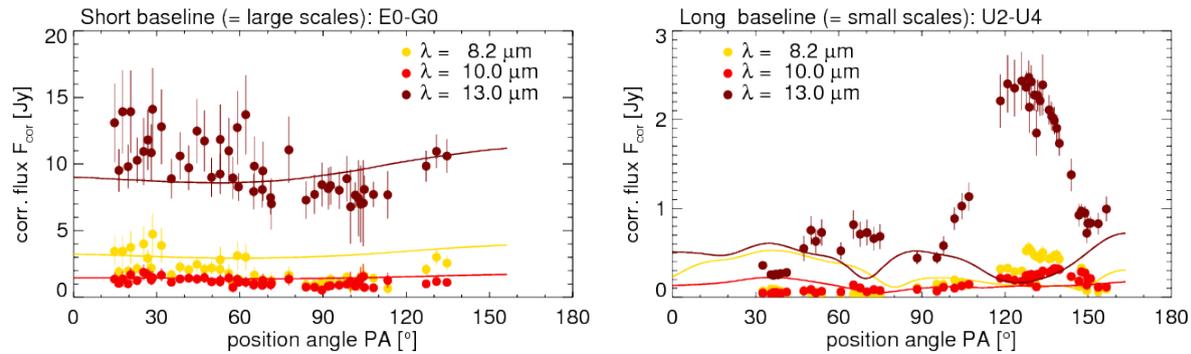
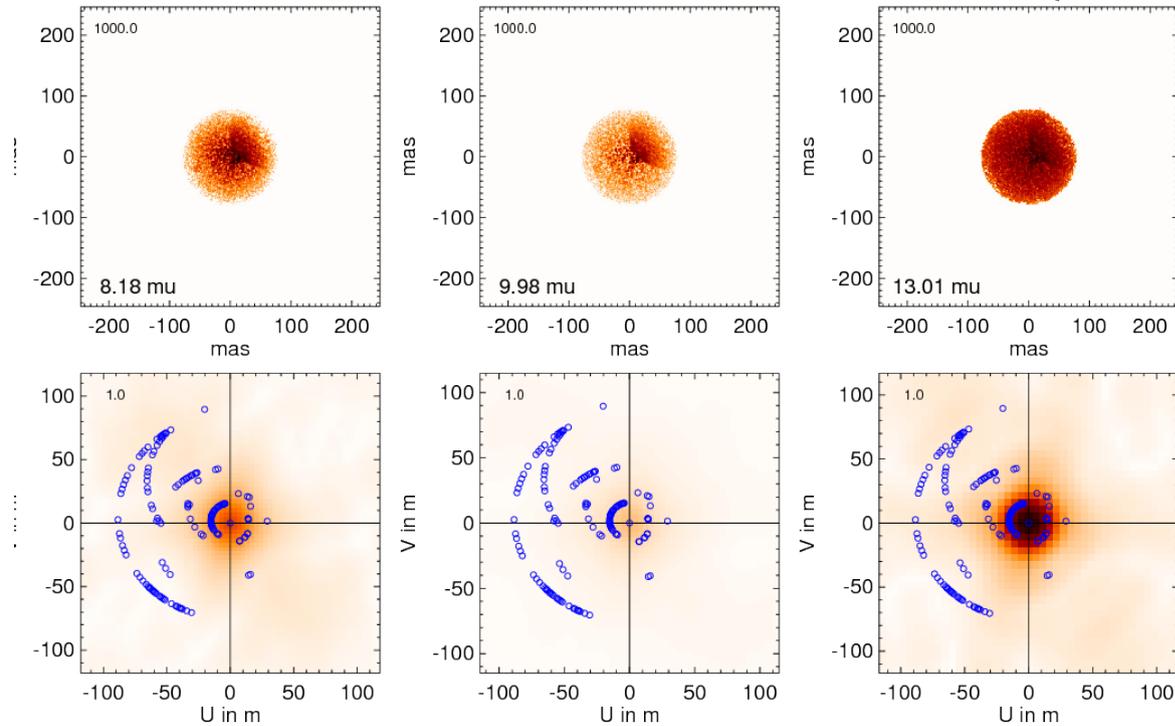
disk+cone wind: not so good match with VLTI/MIDI

Stalevski, Tristram & Asmus, subm.



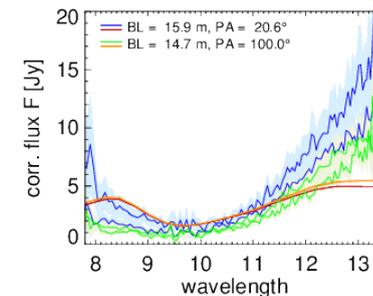
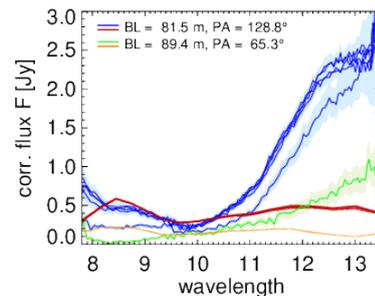
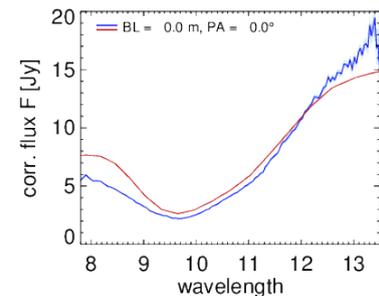
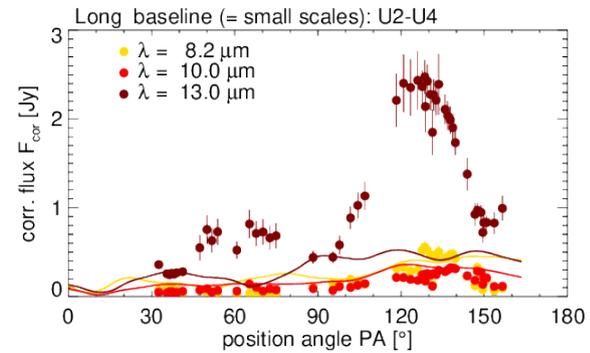
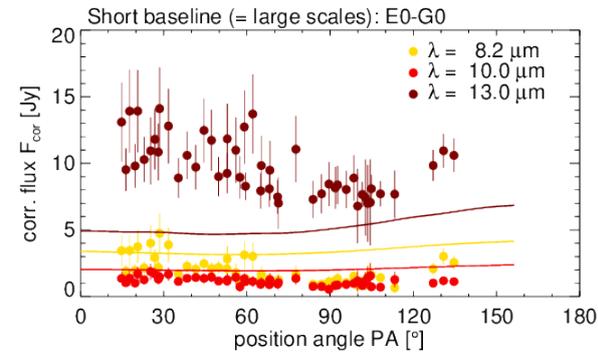
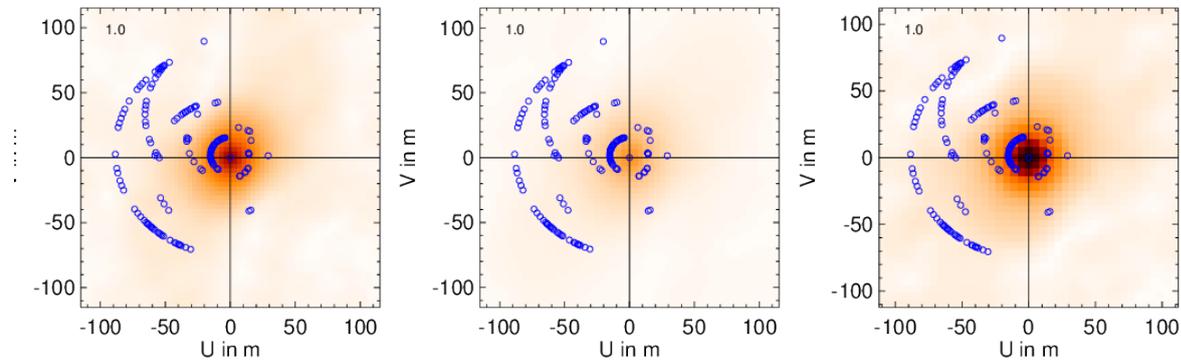
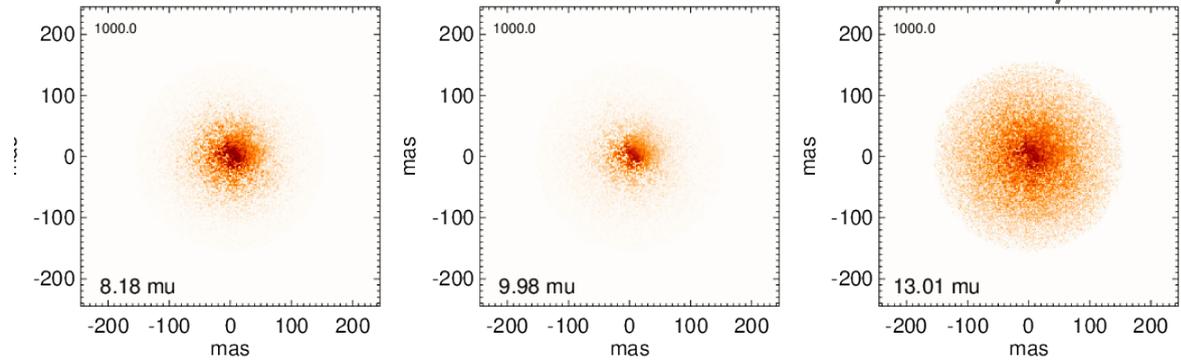
clumpy torus: doesn't work with VLTI/MIDI

Stalevski, Tristram & Asmus, subm.



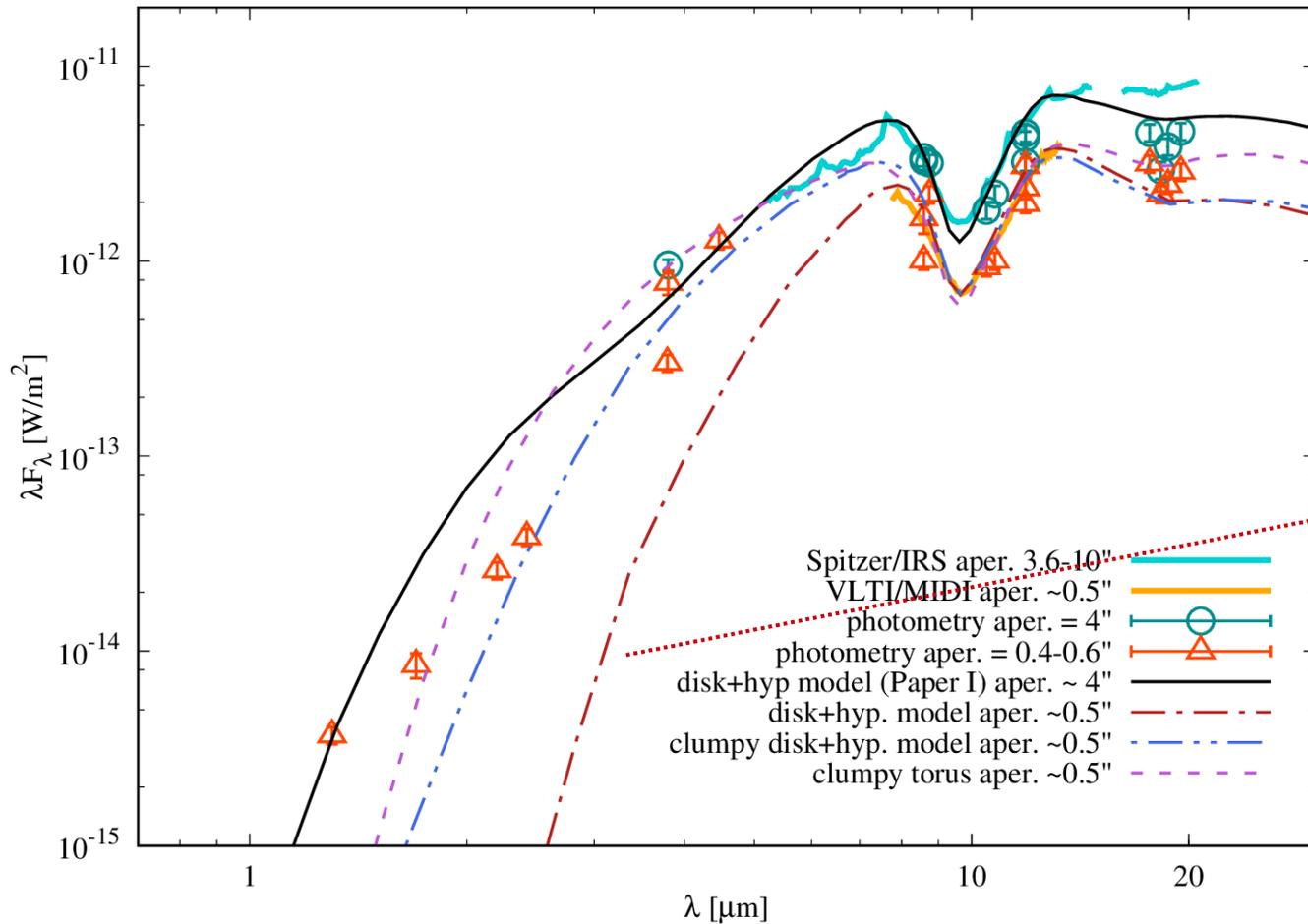
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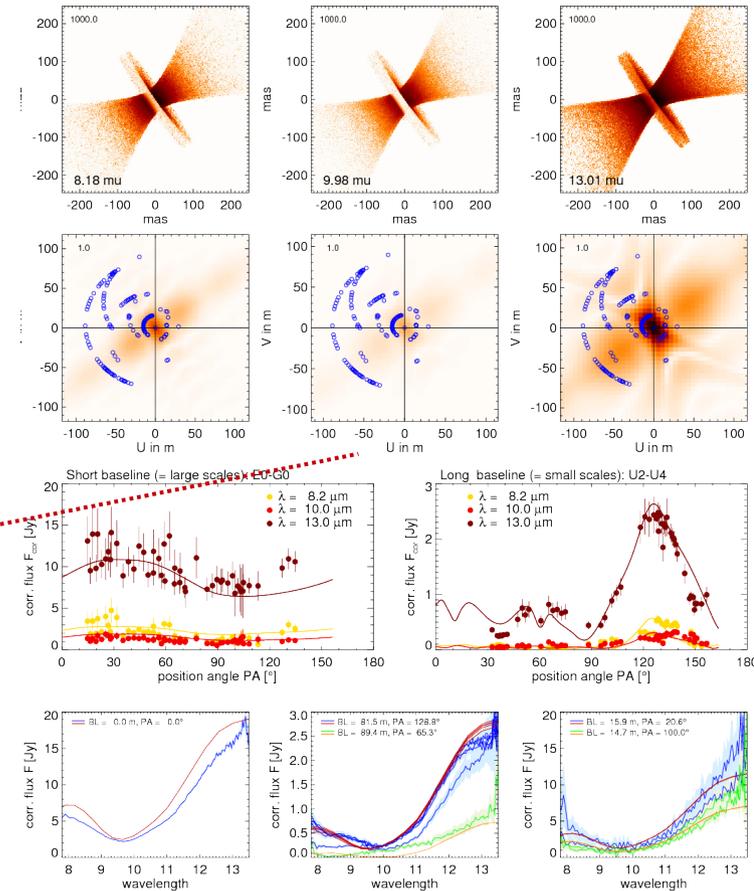


disk+wind

Stalevski, Tristram & Asmus, subm.



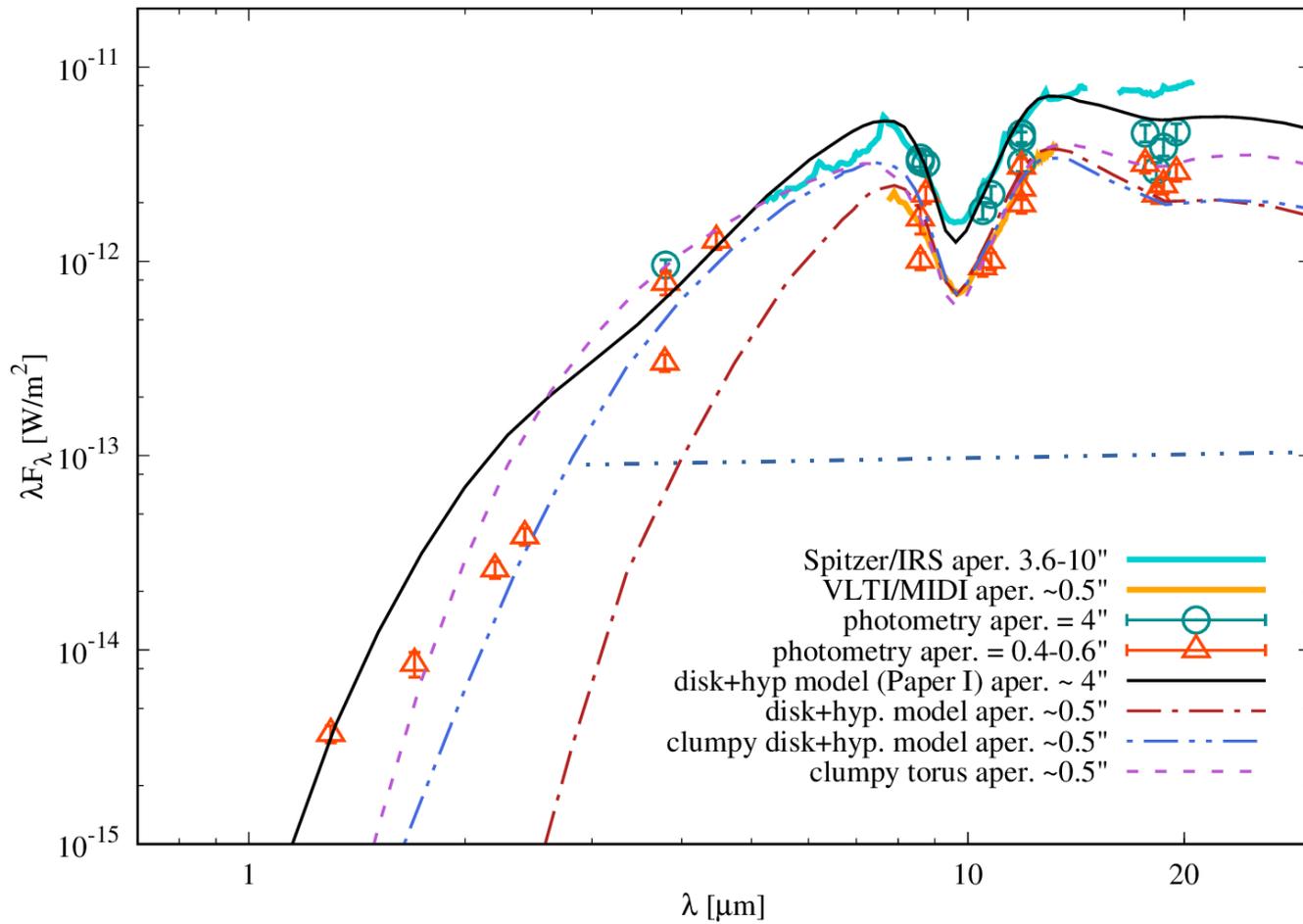
MIR good, NIR missing



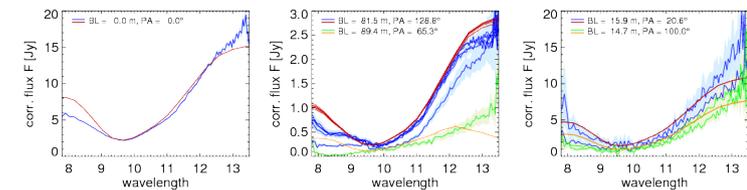
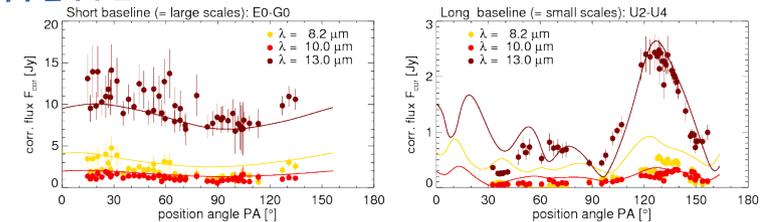
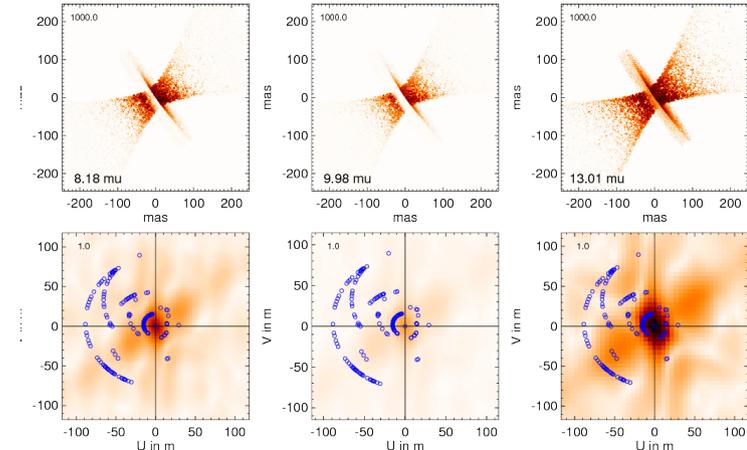
VLT/MIDI: good

disk+clumpy wind

Stalevski, Tristram & Asmus, subm.



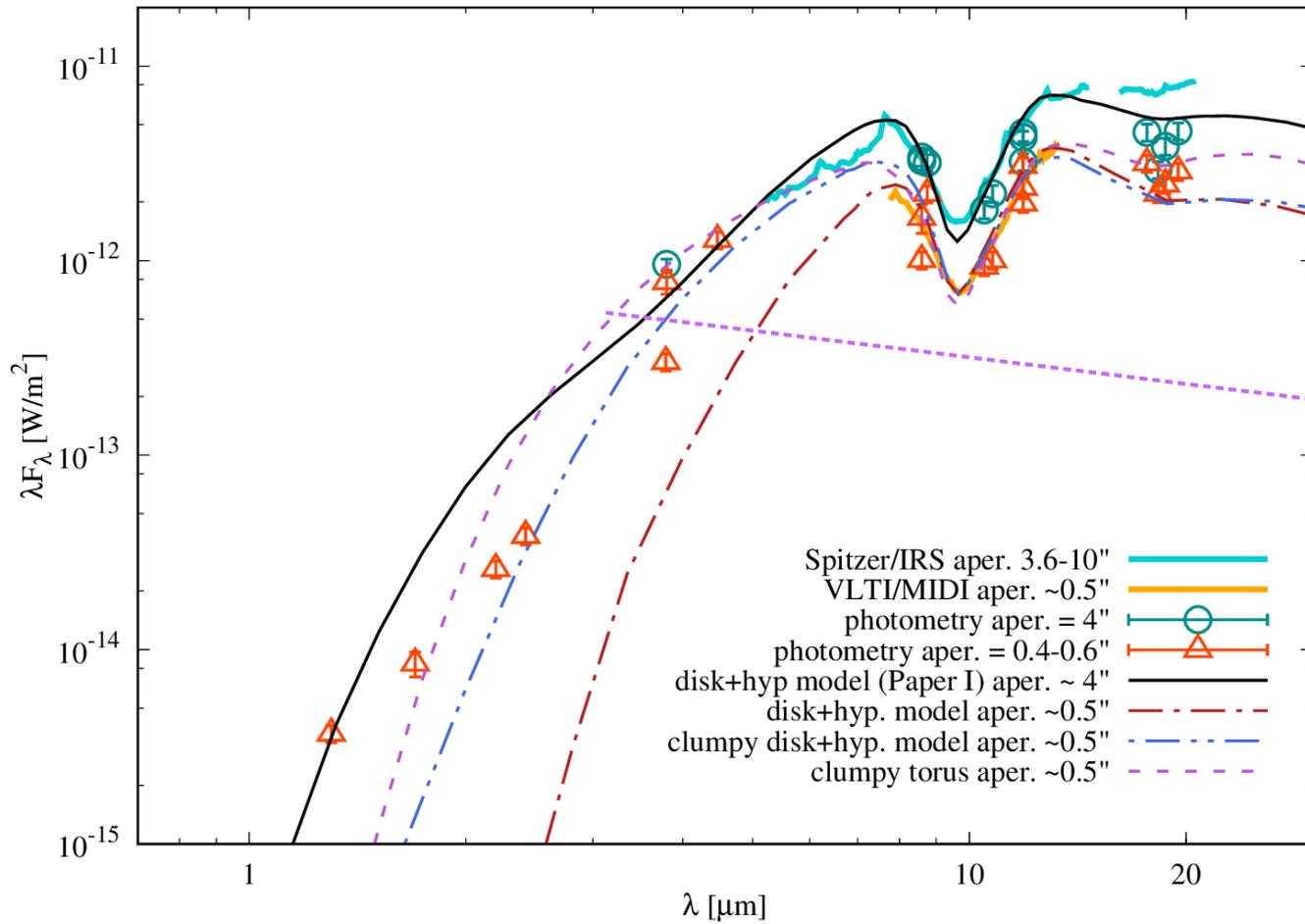
NIR-MIR: good



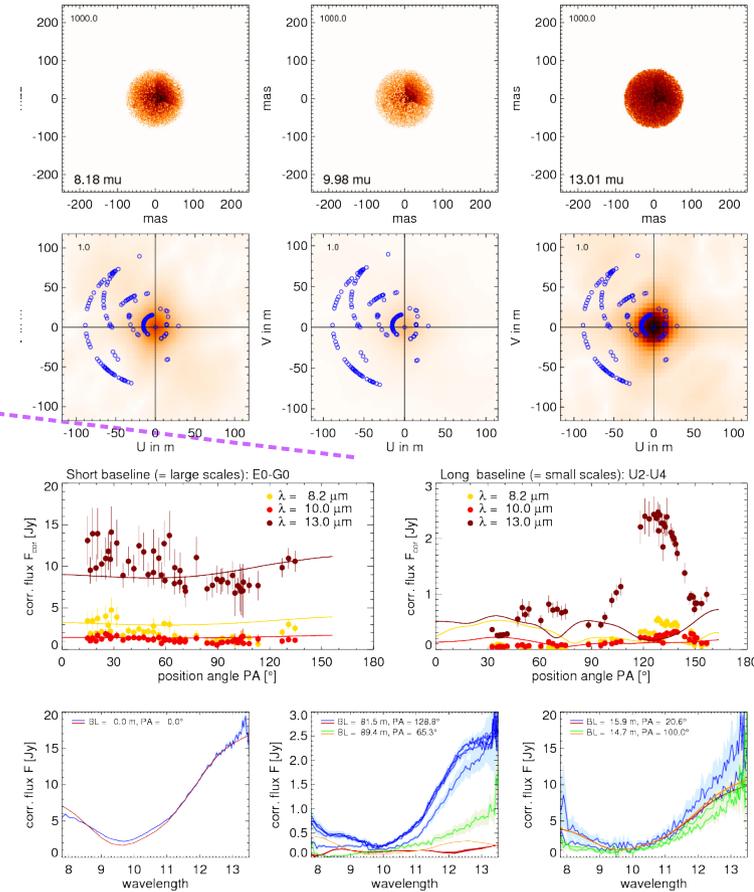
VLT/MIDI: good

clumpy torus

Stalevski, Tristram & Asmus, subm.



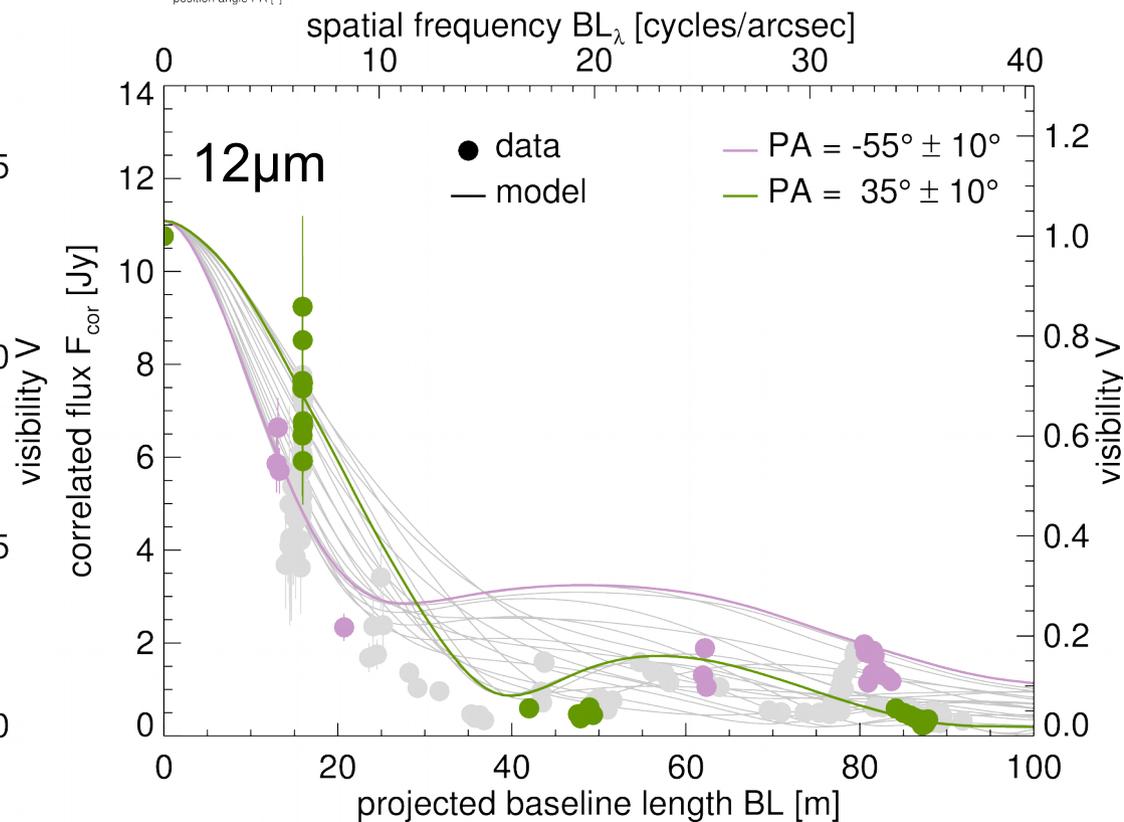
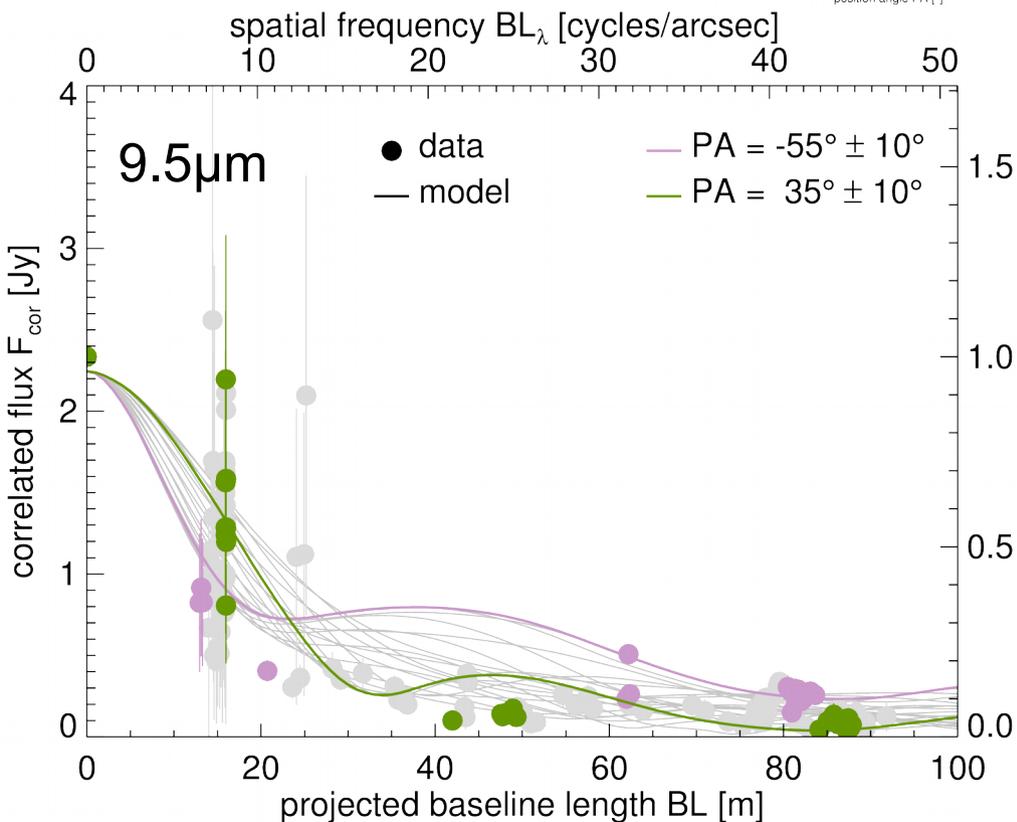
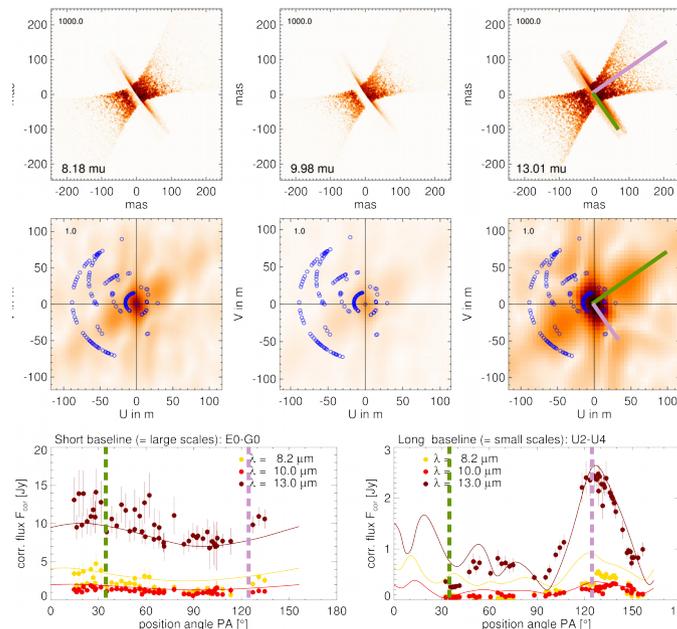
NIR-MIR: good



VLT/MIDI: ugly

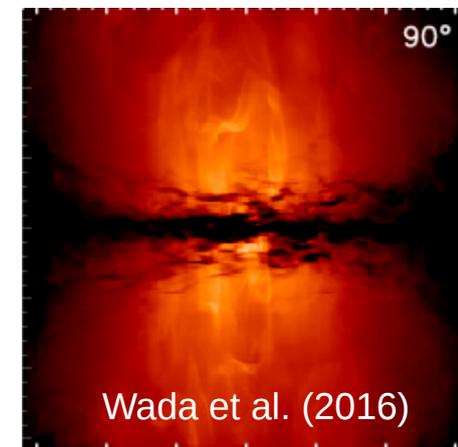
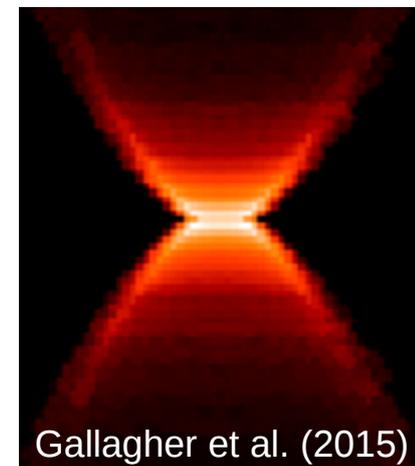
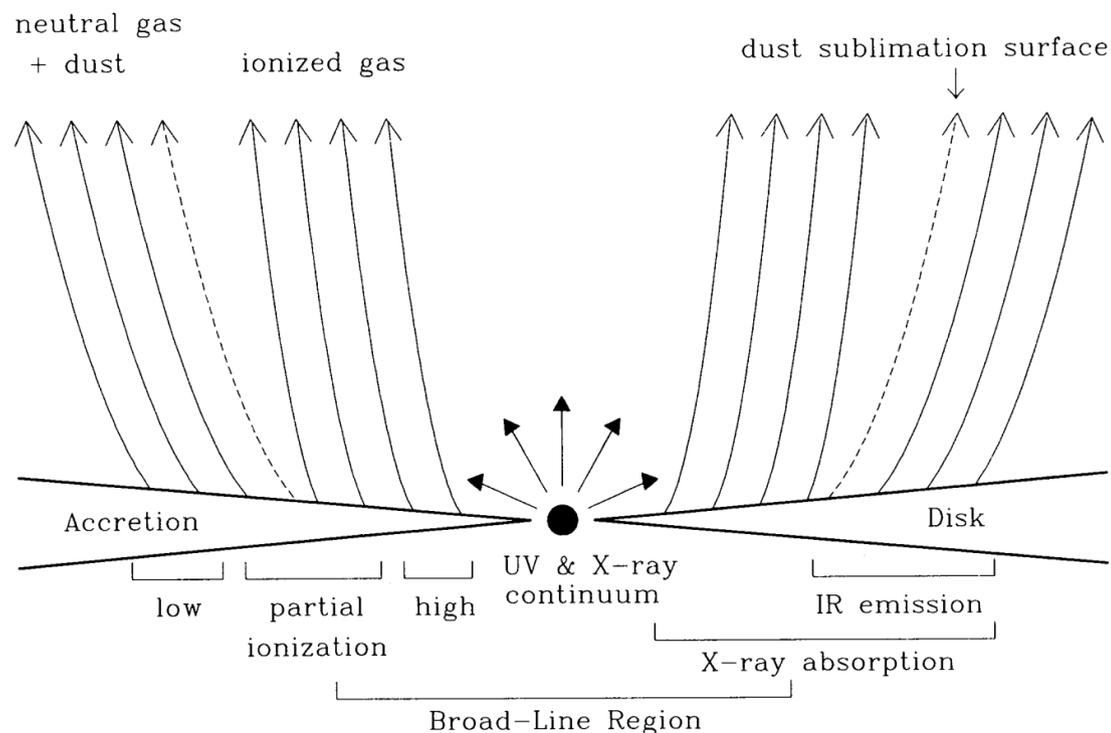
disk+ (clumpy) wind: good match with VLTI/MIDI

Stalevski, Tristram & Asmus, subm.



Radiation pressure driven dusty winds

Radiation pressure drives away the gas and dust from the accretion disk to the polar regions, leaving behind what may constitute the obscuring dusty disk/torus



[Konigl & Kartje (1994); Keating+ (2012); Roth et al. (2012); Gallagher+ (2015)]

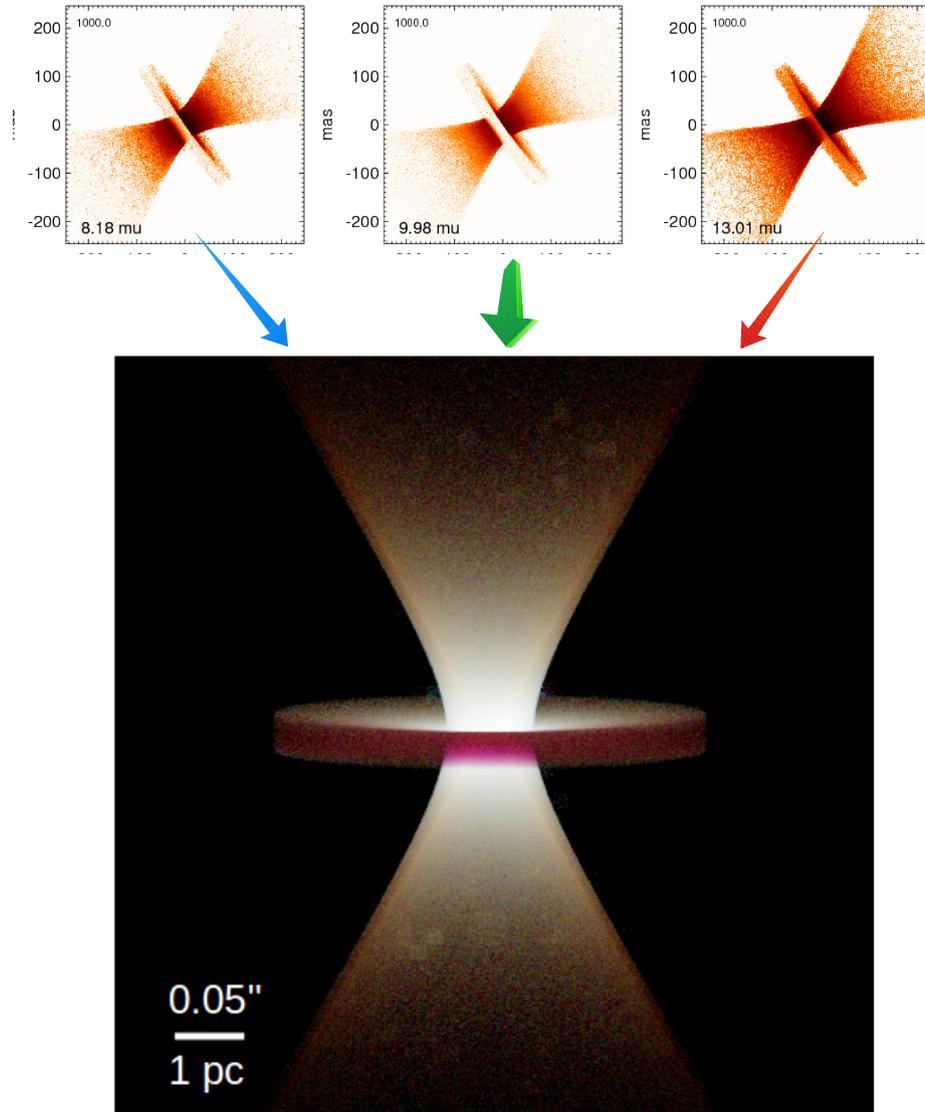
[Dorodnitsyn+ (2011, 2012); Dorodnitsyn & Kallman (2012); Dorodnitsyn+ (2016)]

[Wada (2012), Wada+ (2016), Chan & Krolik (2016, 2017), Vollmer et al. (2018)]

Conclusions

Stalevski, Asmus & Tristram (2017)
+
Stalevski, Tristram & Asmus, subm.

Dust in Circinus AGN: disk + wind



A prototype for polar dust AGN population