

# **Nuclear obscuration in AGN: an X-ray perspective**

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***Kavli Institute for Astronomy and Astrophysics, Beijing, China***

1993



1993



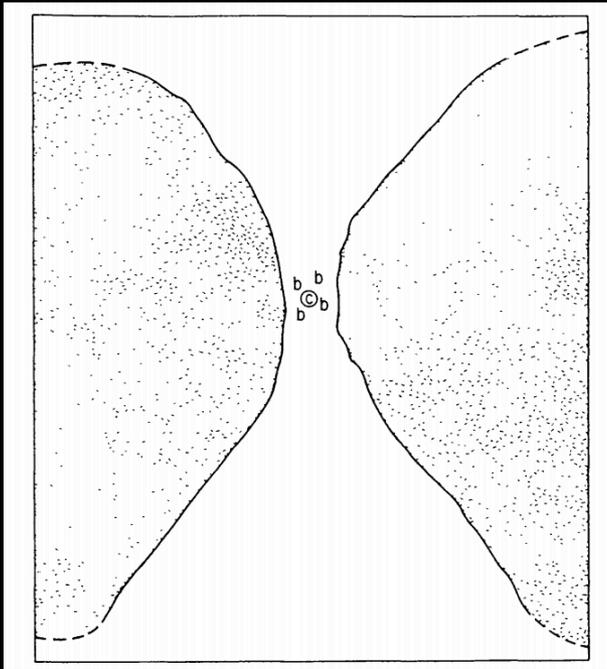
2018



Cake by P. Boorman (see

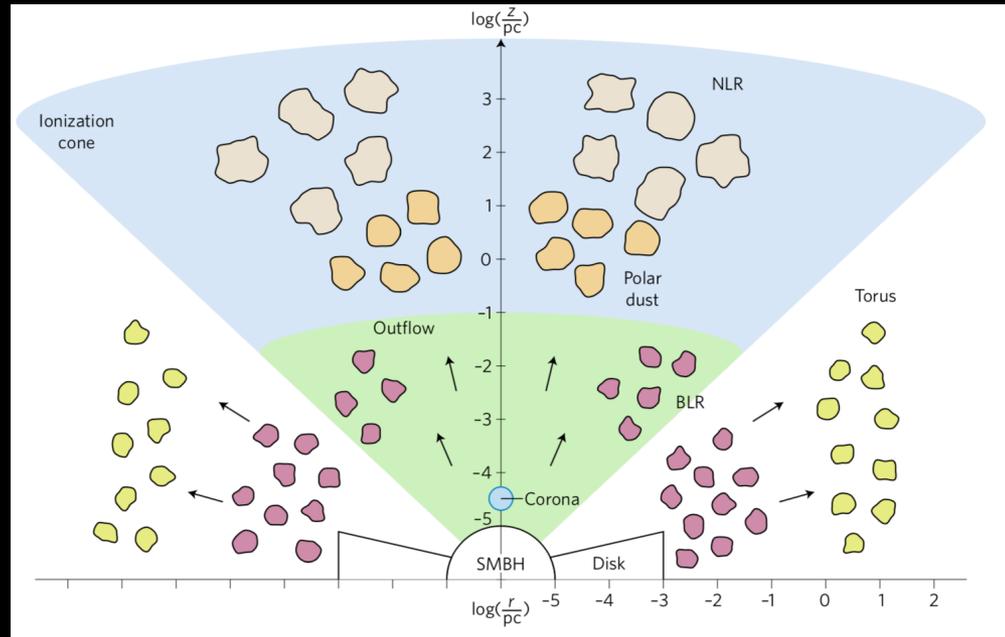
<https://www.southampton.ac.uk/~pgb2g11/baking.html>)

1993



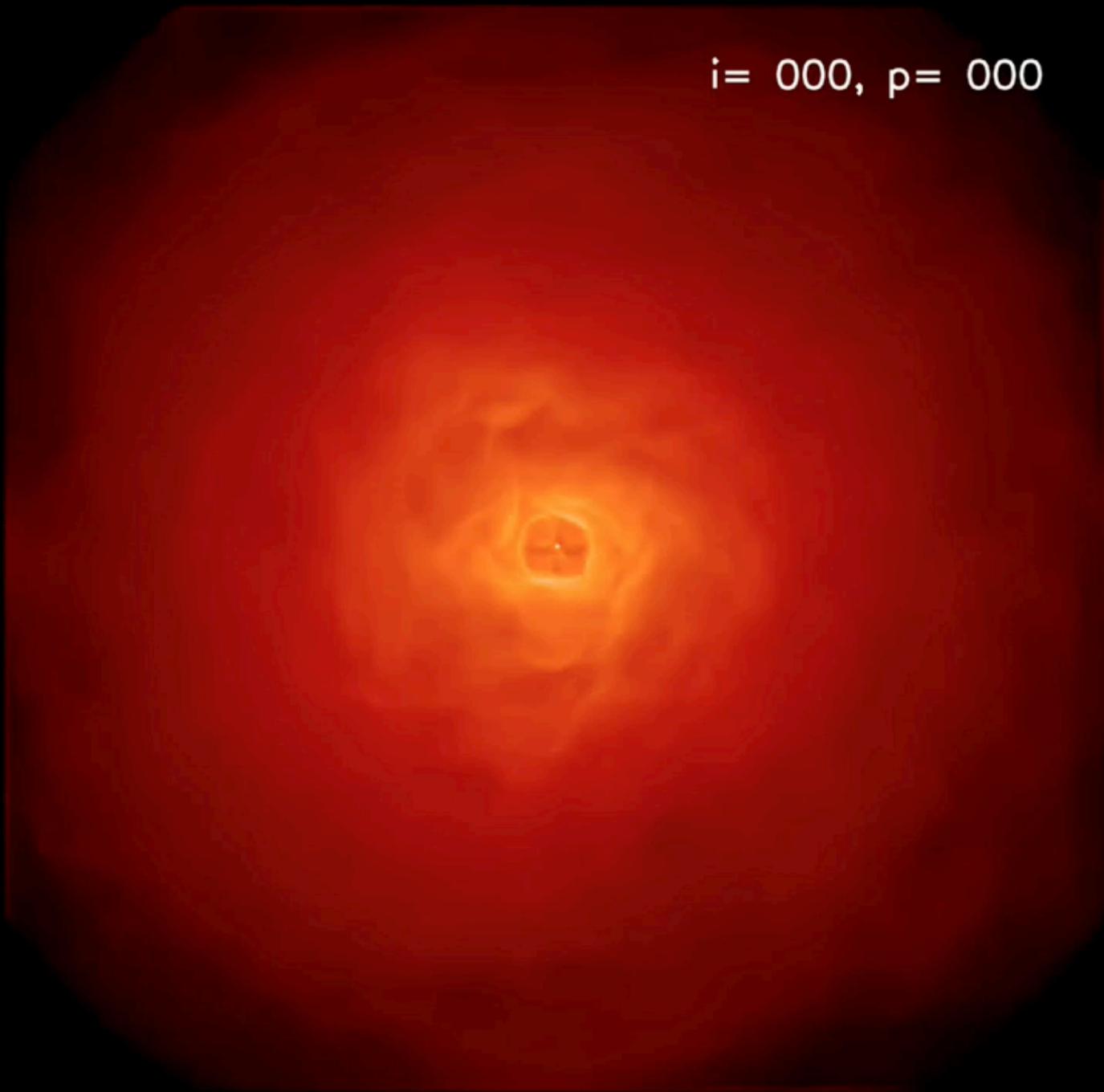
Antonucci (1993)

2018



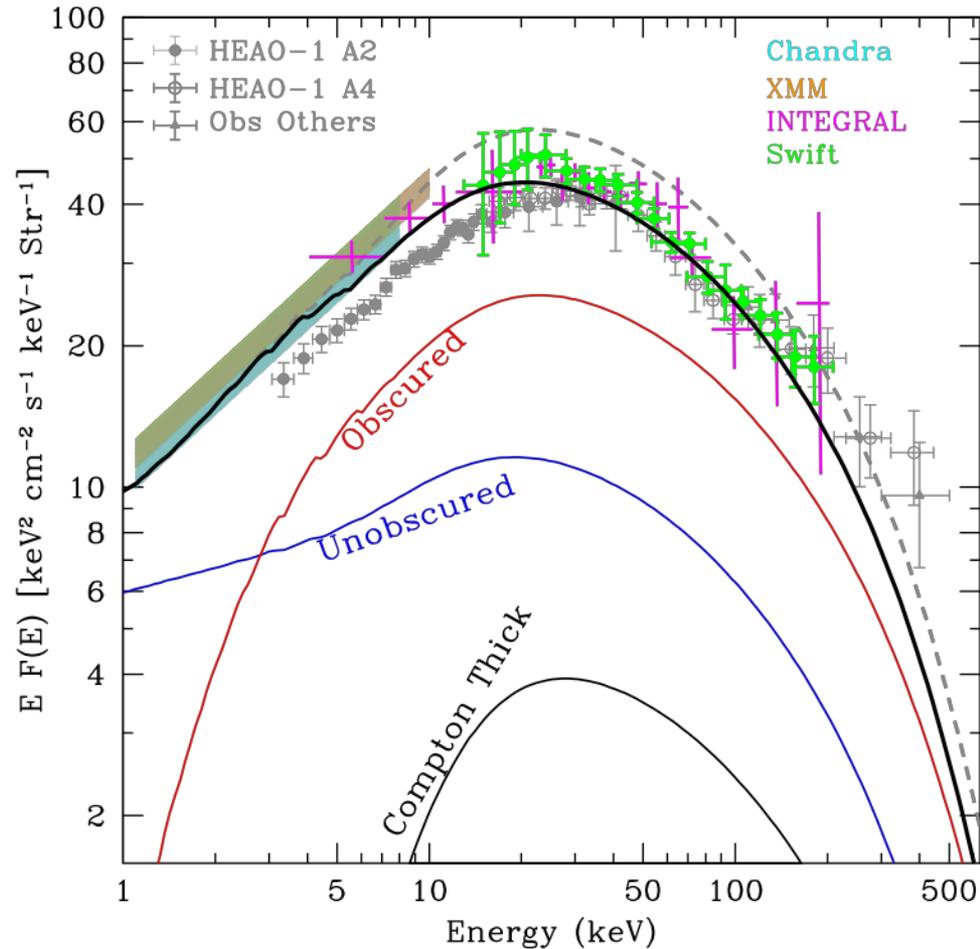
Ramos Almeida & Ricci (2017)

$i = 000, p = 000$



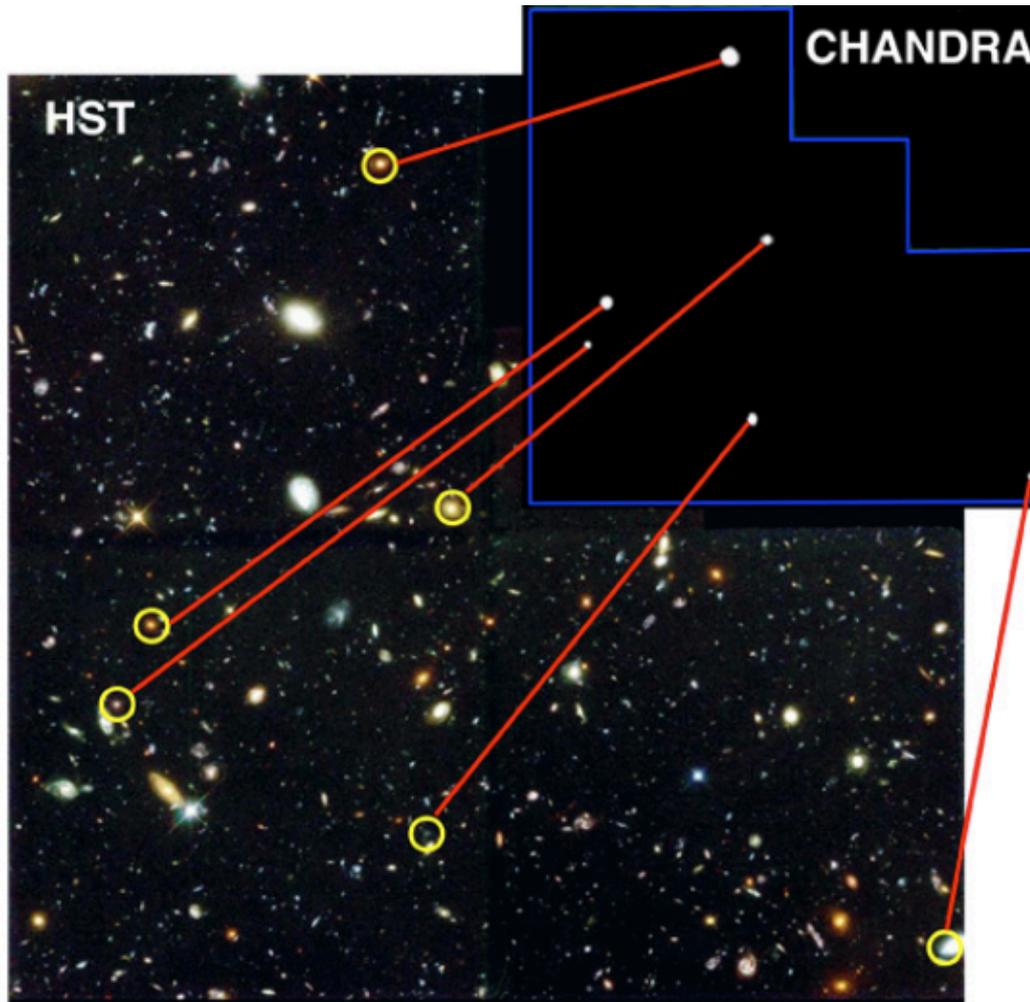
*Wada et al.  
(2016); See  
talks by K.  
Wada, D.  
Williamson,  
C.-H. Chan, A.  
Dorodnitsyn,  
D. Angles-  
Alcazar*

# Obscured accretion



Treister et al. (2009); see also Tasnim Ananna et al. (2018), Ueda et al. (2014), Akylas et al. (2012), Draper & Ballantyne (2010), Gilli et al. (2007)

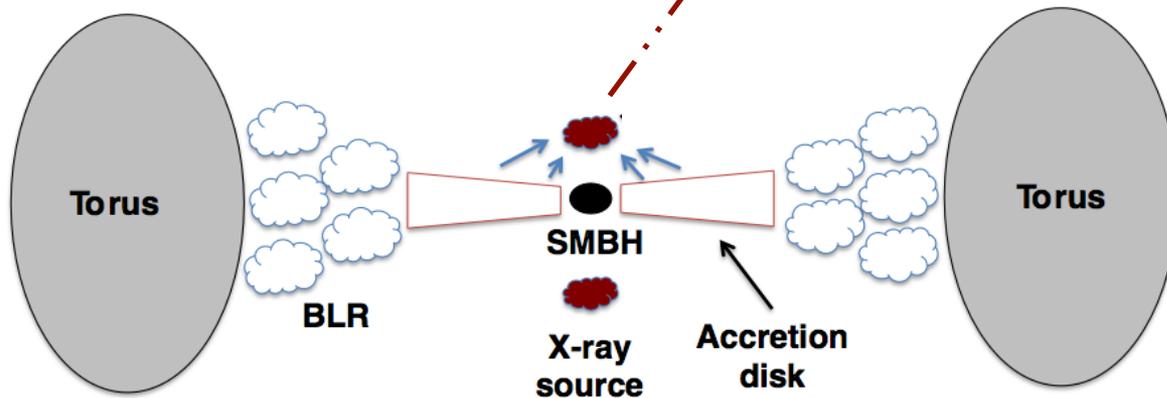
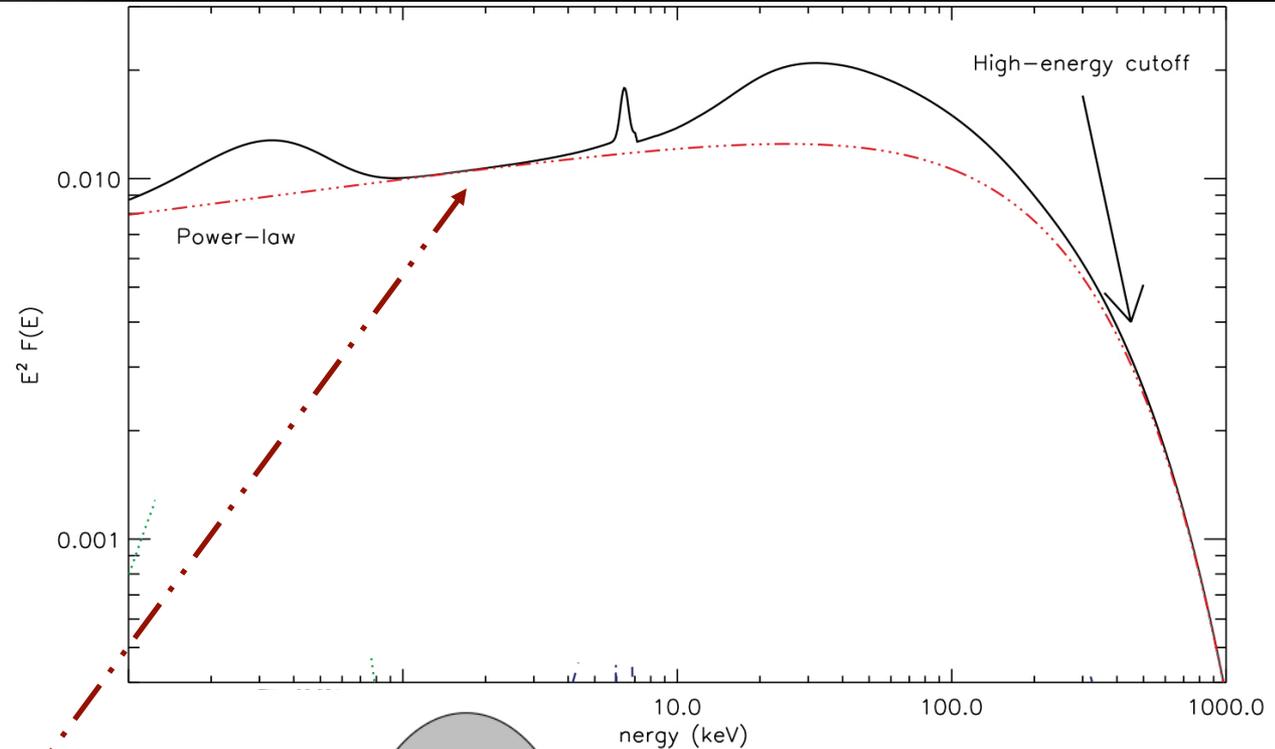
# X-ray emission of AGN



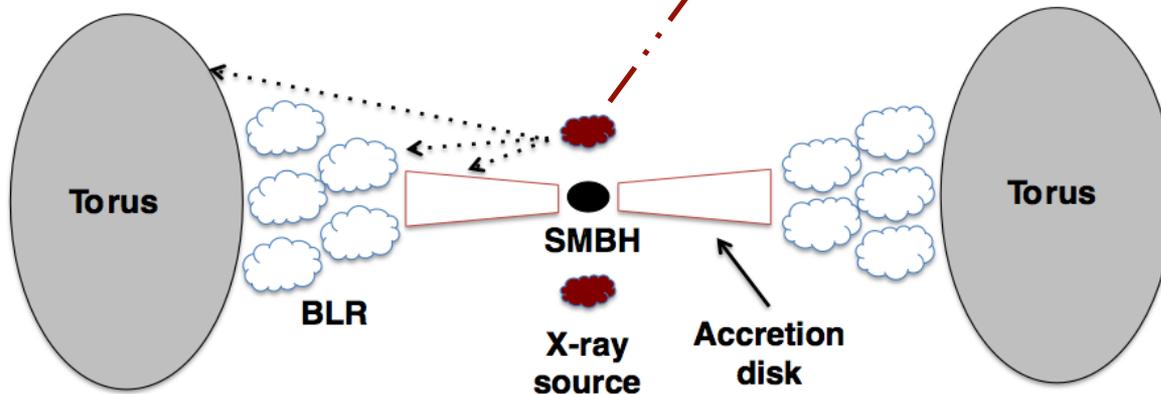
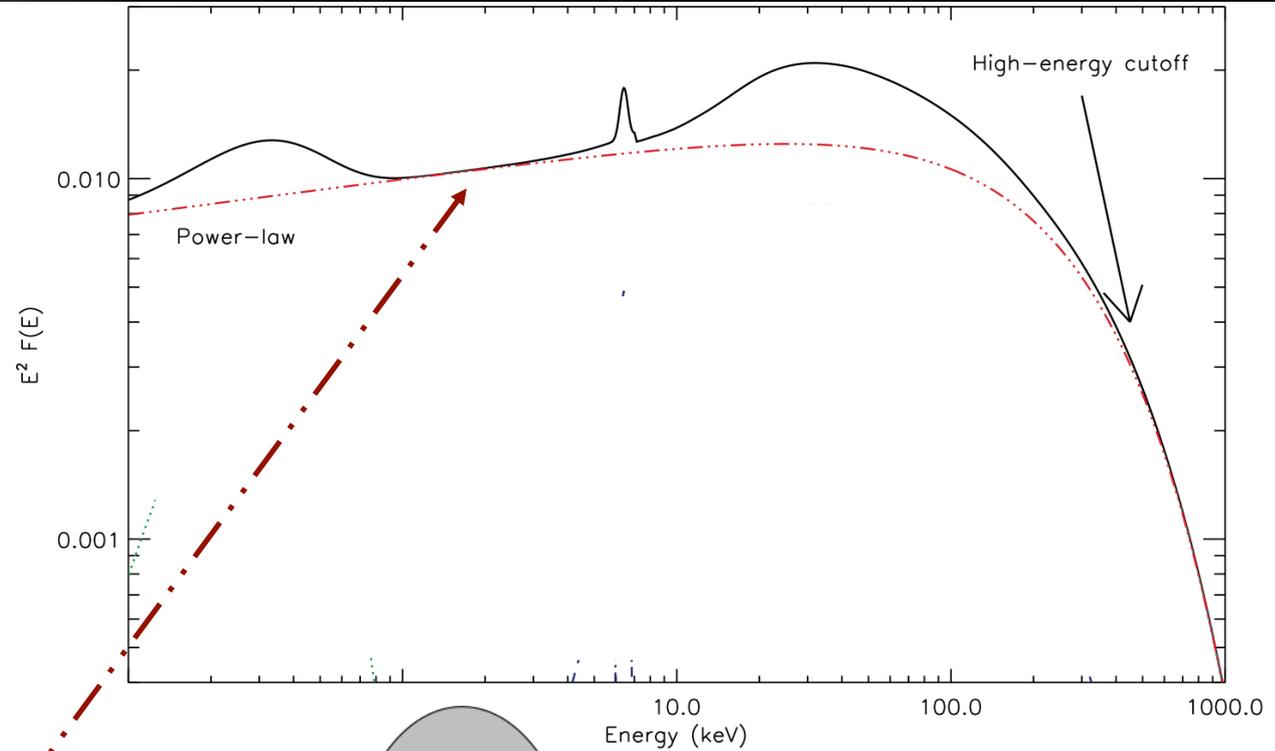
Hubble  
deep-field  
north

*Credit: Optical: NASA/HST, X-ray: NASA/PSU*

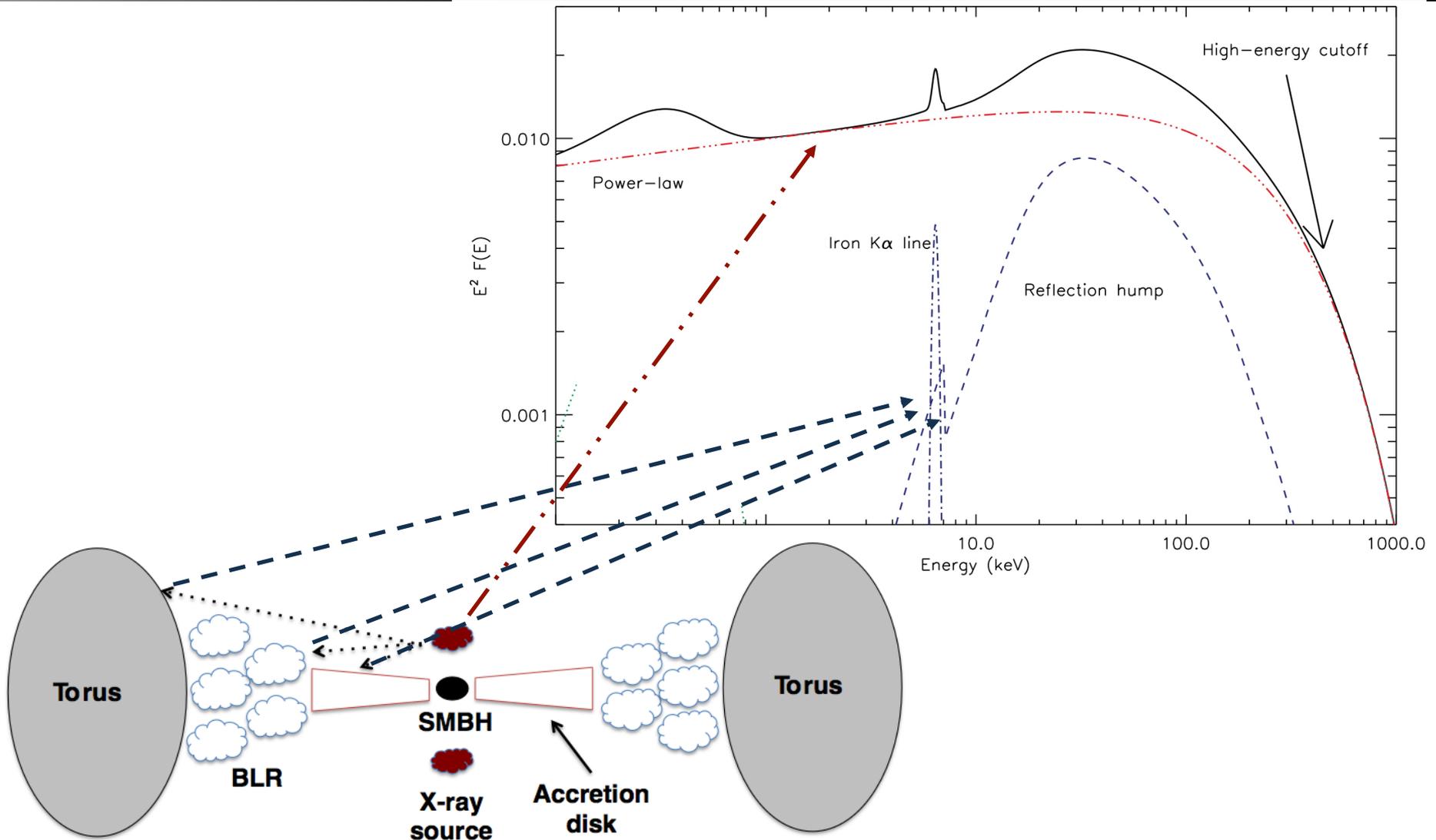
# X-ray spectra of AGN



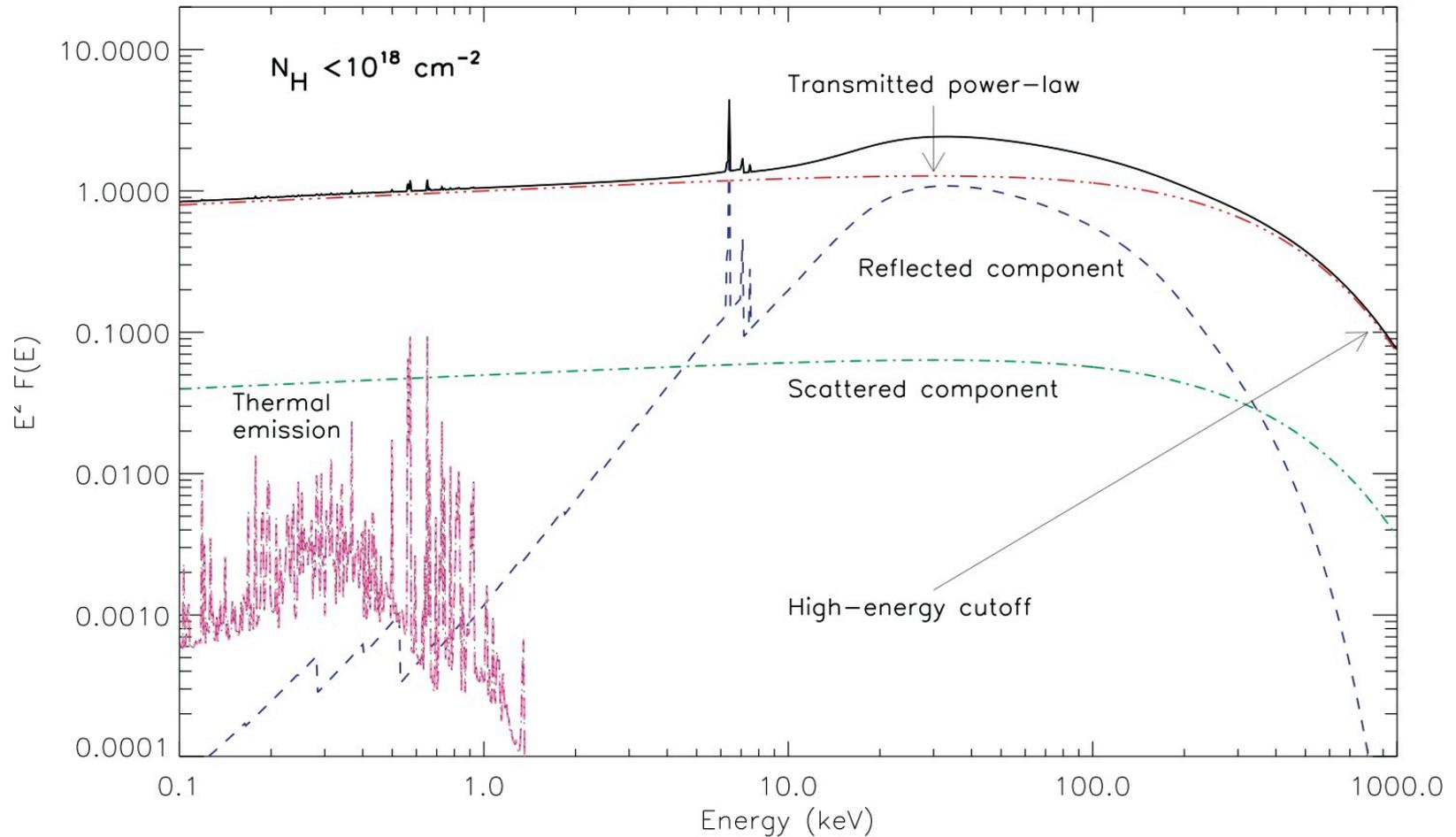
# X-ray spectra of AGN



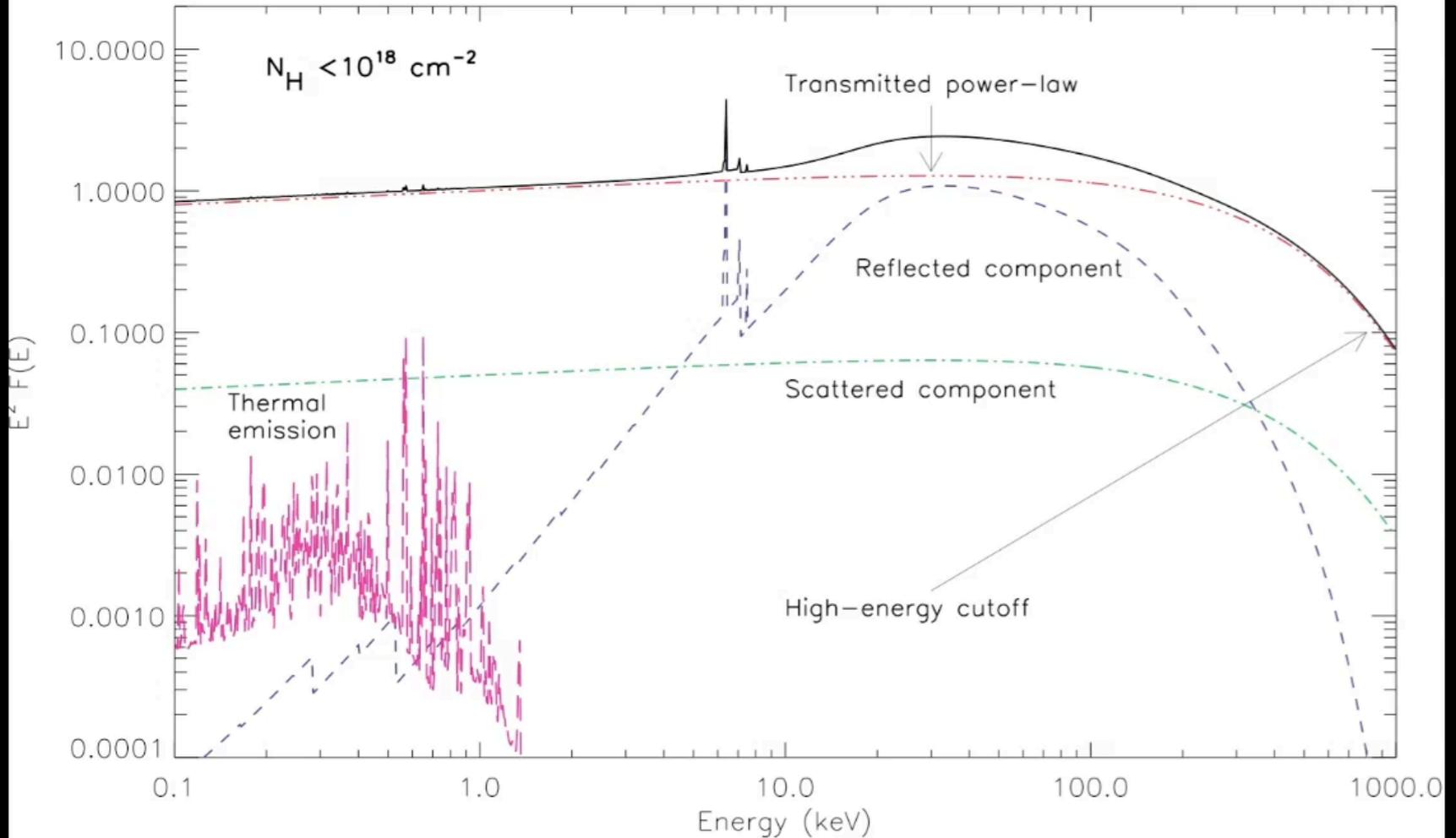
# X-ray spectra of AGN



# Absorption in the X-rays

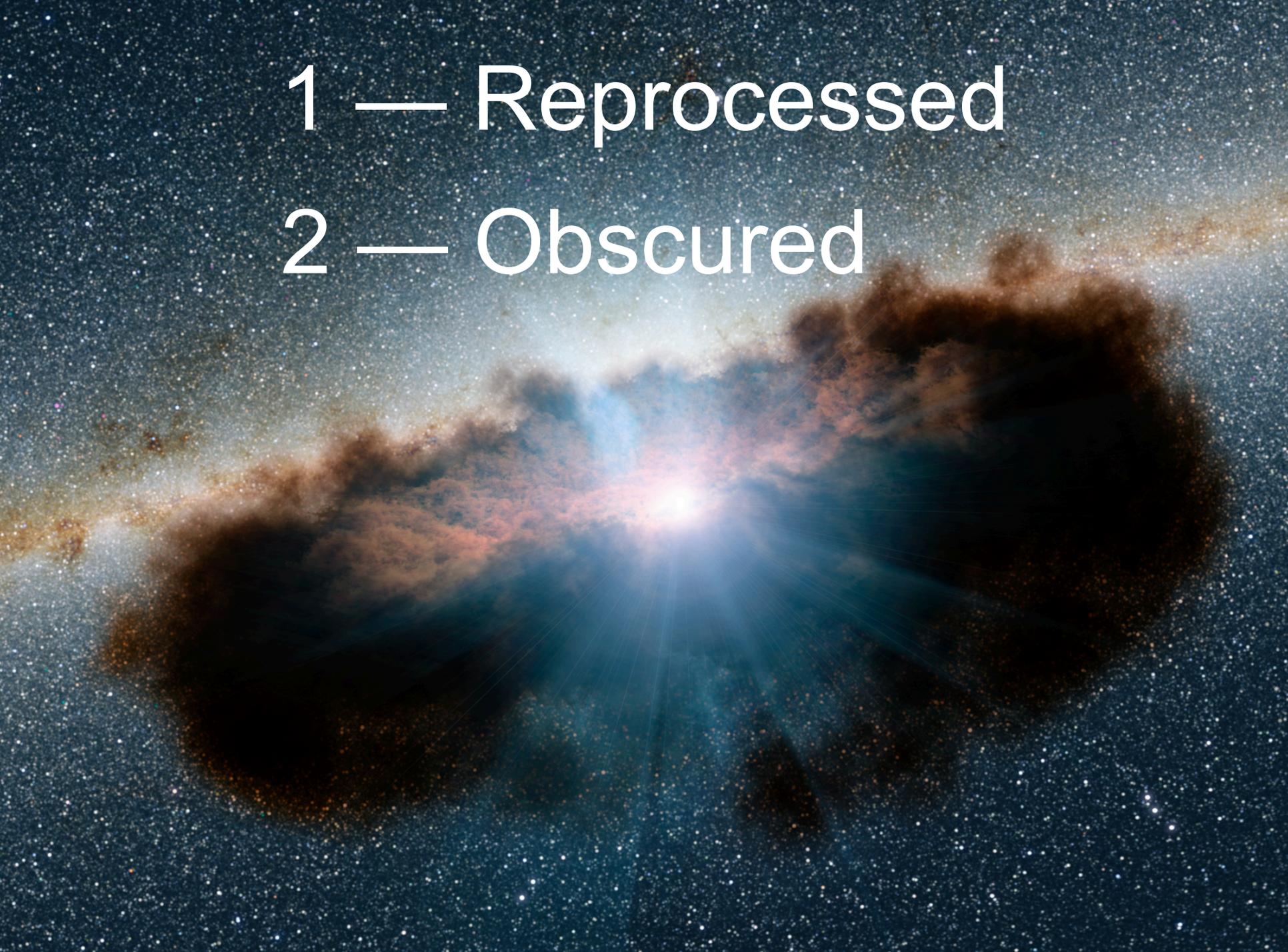


# Absorption in the X-rays

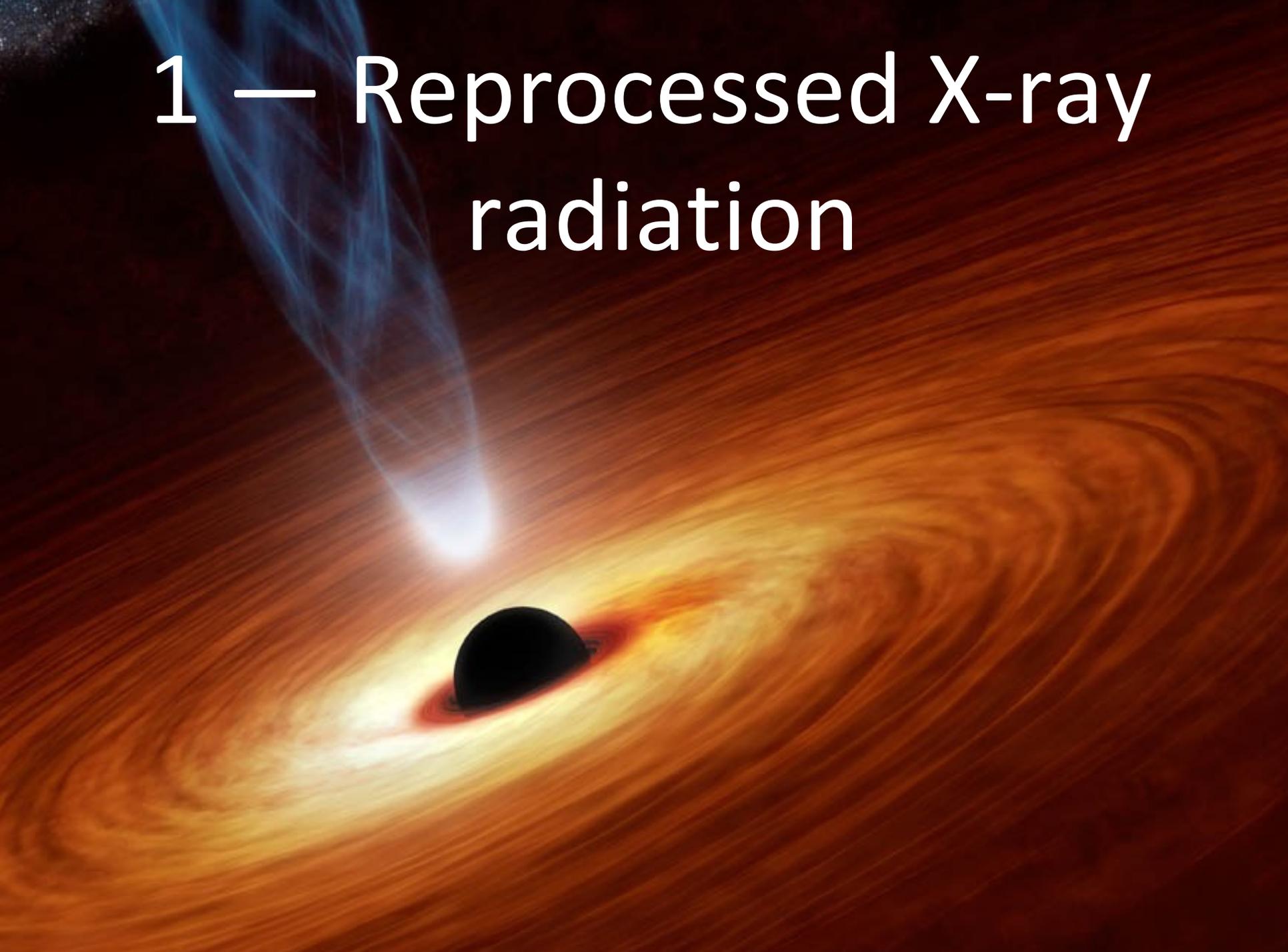


1 — Reprocessed

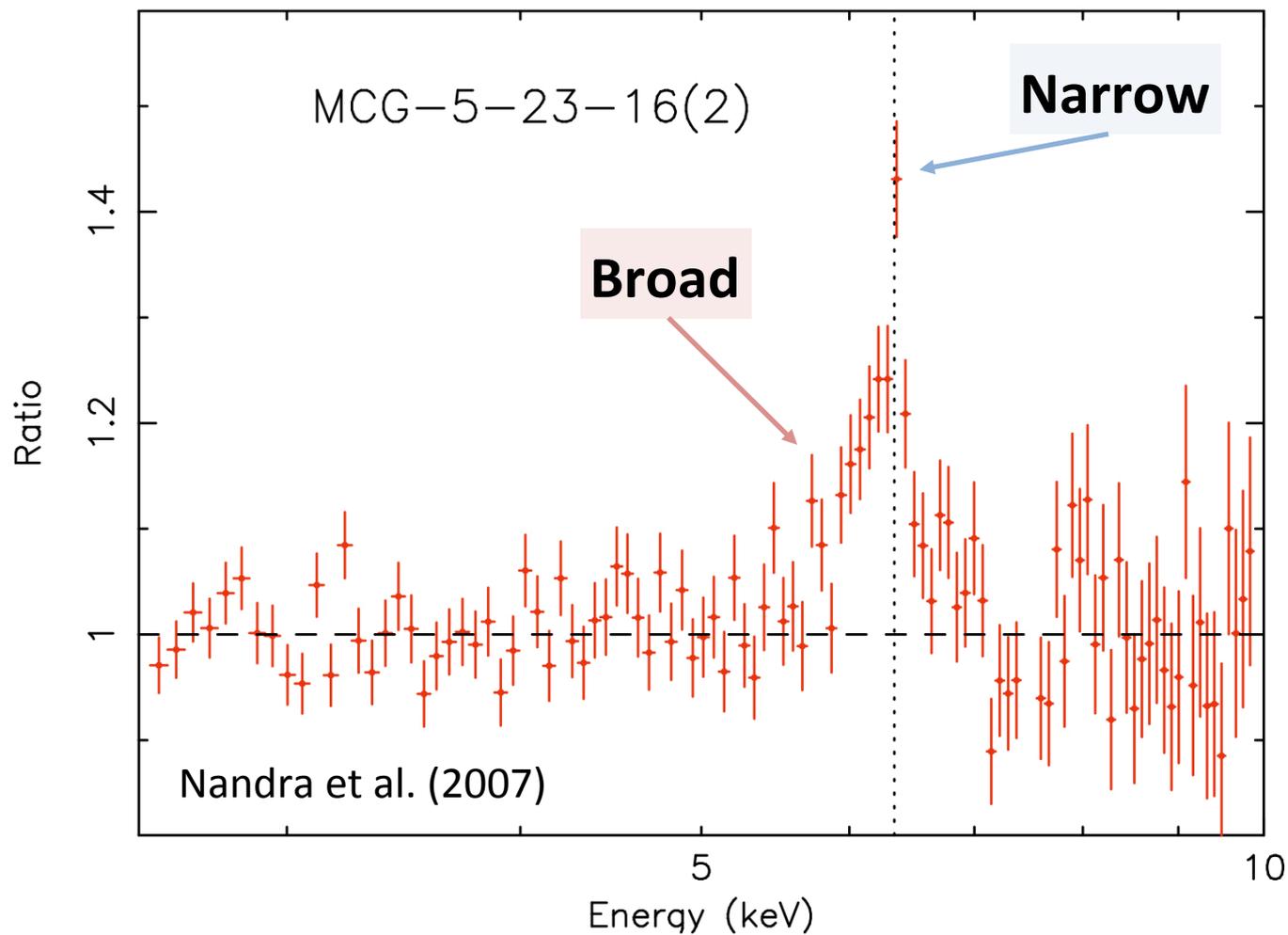
2 — Obscured



# 1 — Reprocessed X-ray radiation



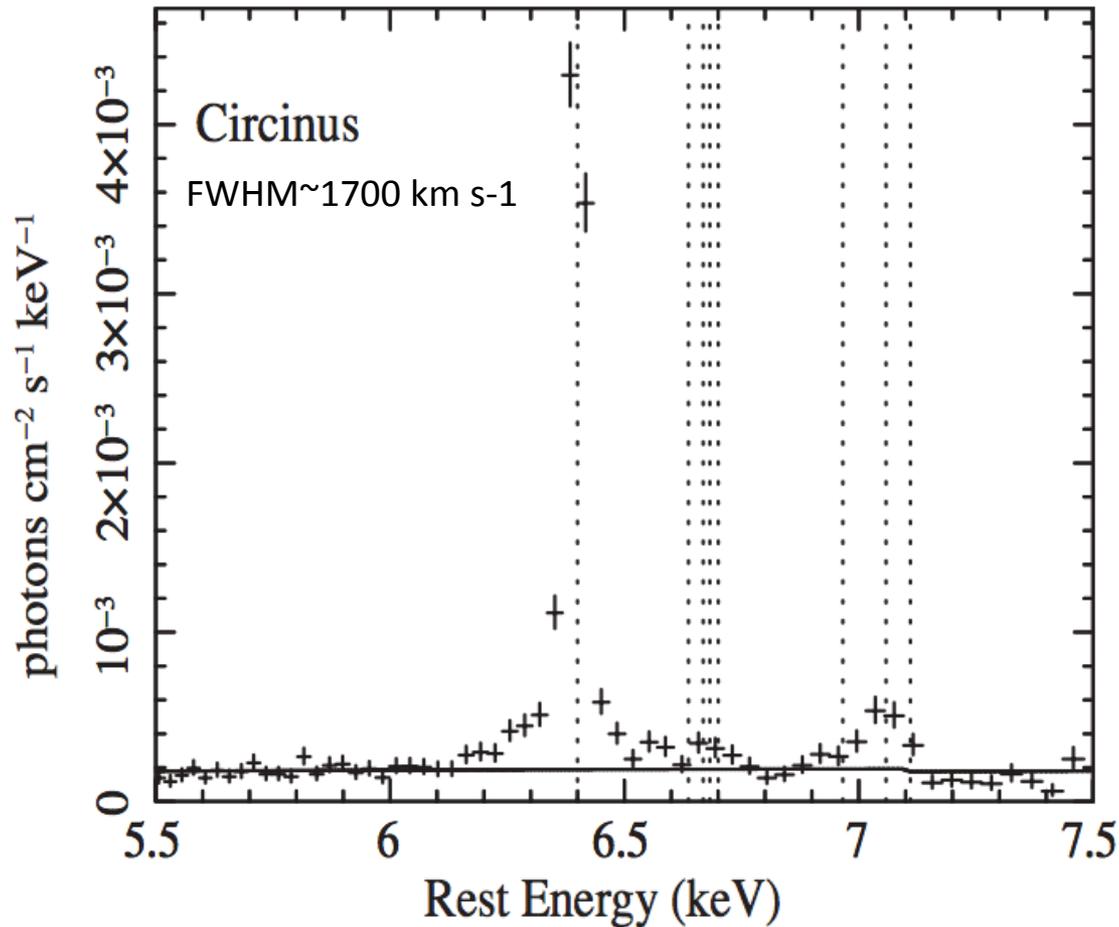
# The Fe $K\alpha$ line



# The Fe K $\alpha$ line



*Chandra/HETG*

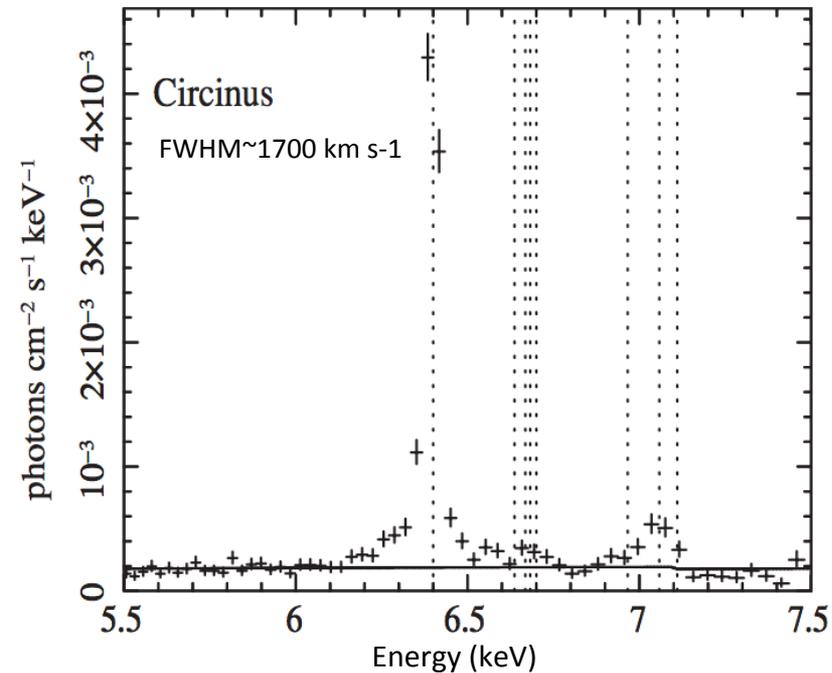


Shu et al. (2011), see also Shu et al. (2010)

# The Fe $K\alpha$ line

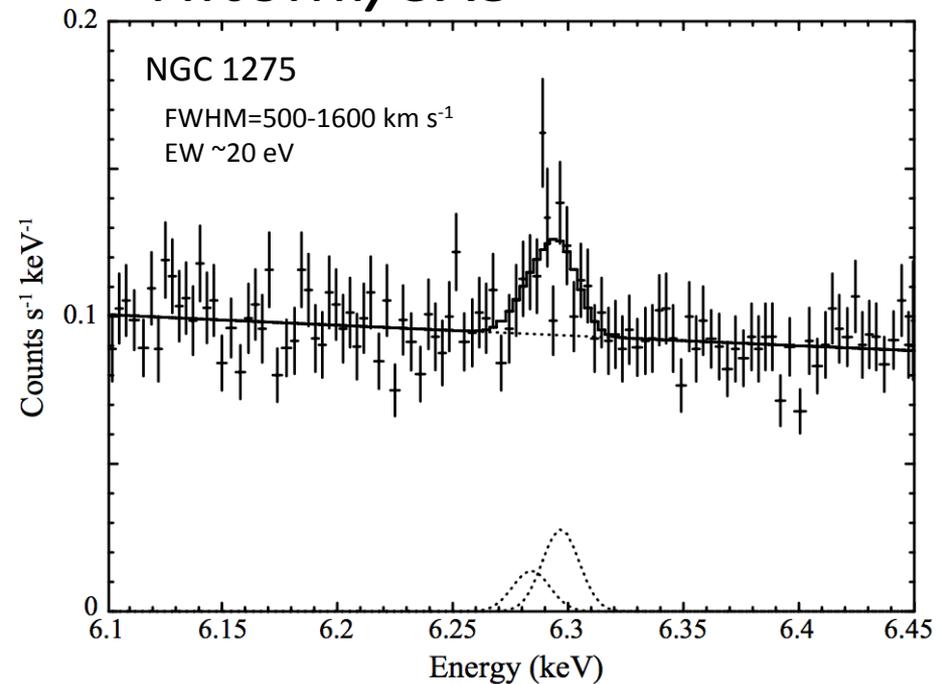


## Chandra/HETG



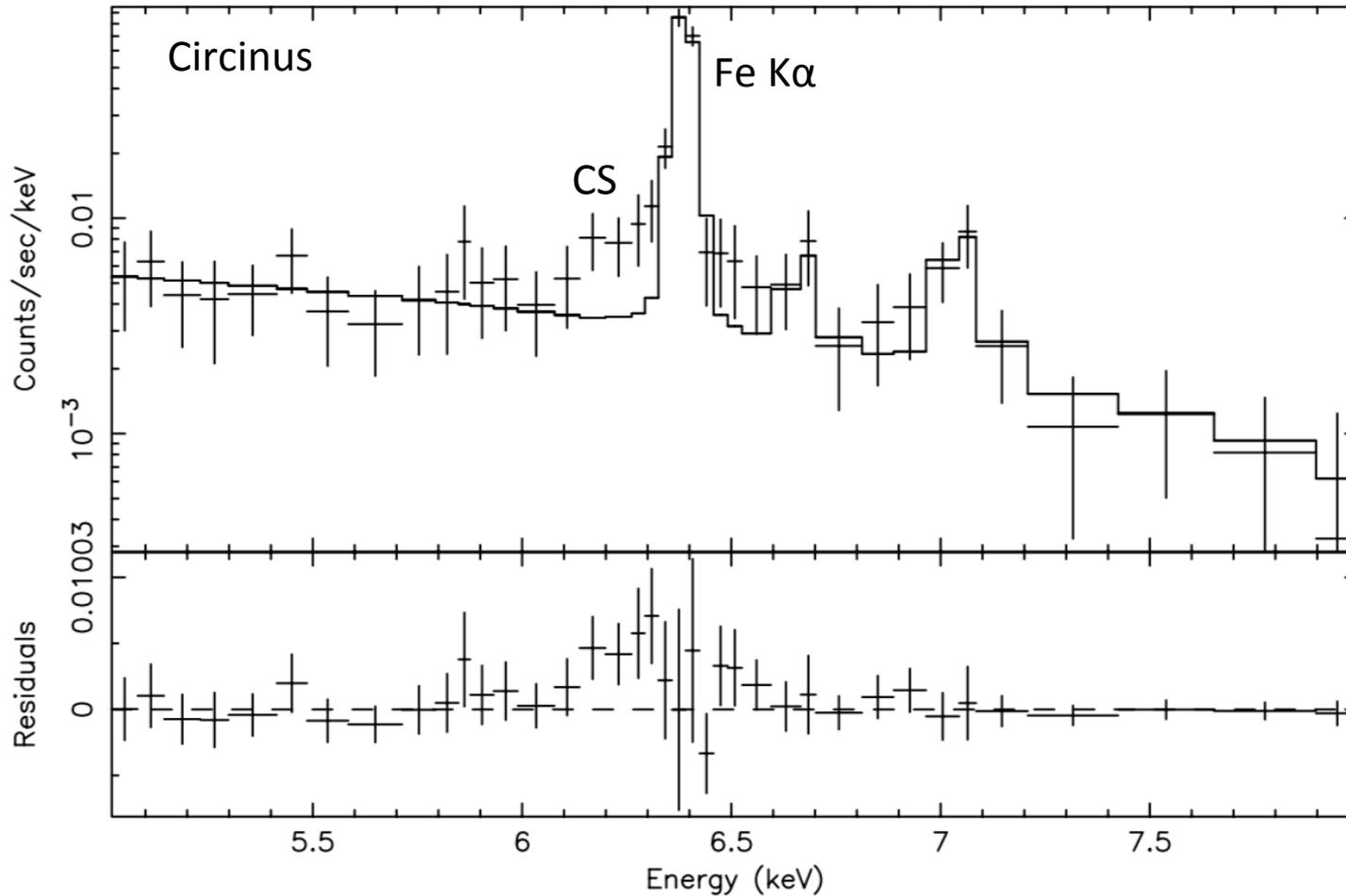
Shu et al. (2011), see also Shu et al. (2010)

## Hitomi/SXS



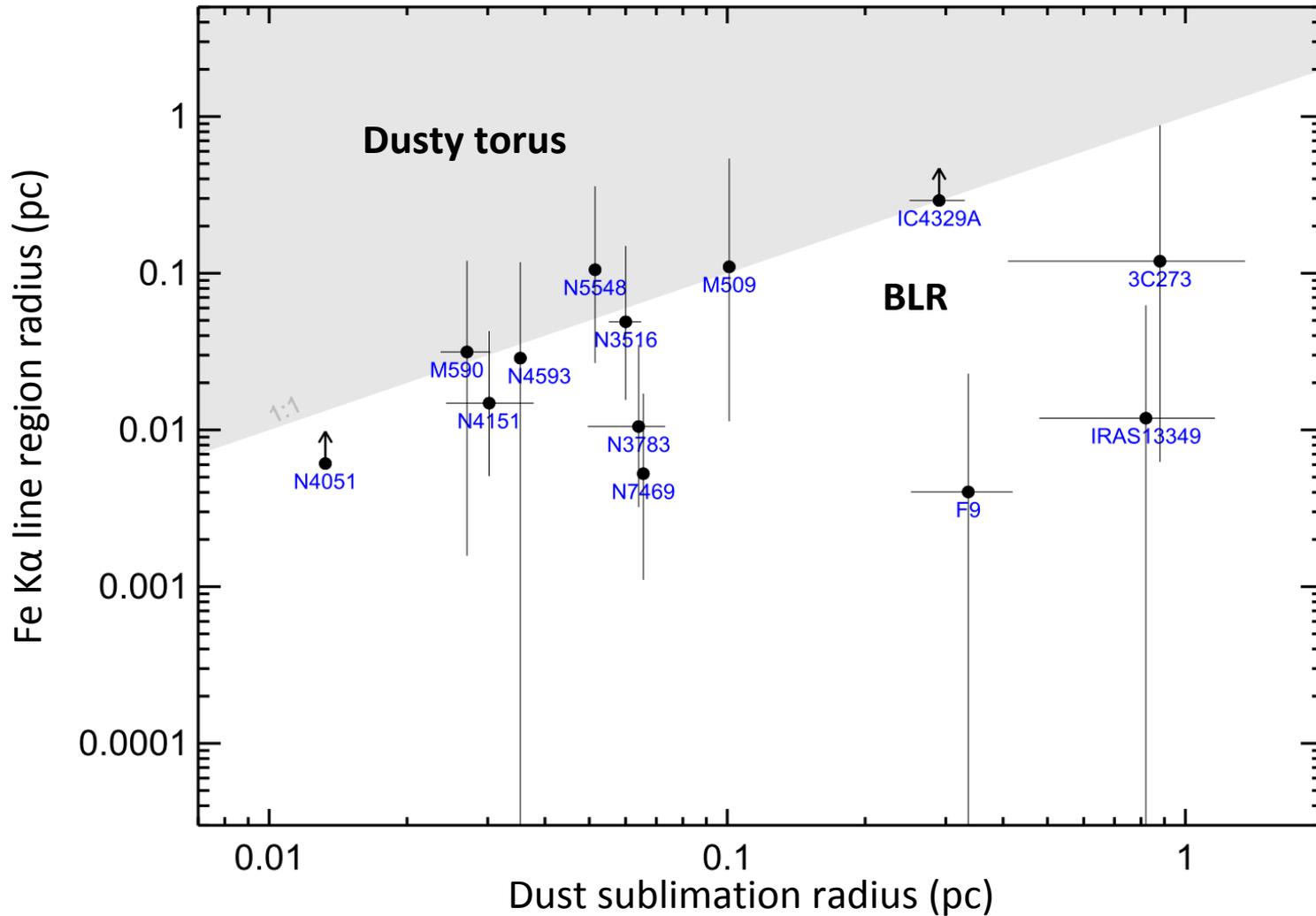
Hitomi collaboration (2018)

# The Compton shoulder



Bianchi et al. (2002); See C. Andonie's talk

# The Fe K $\alpha$ line



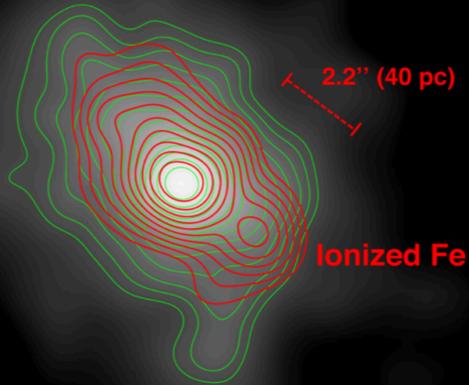
Gandhi et al. (2015)

# Extended Fe K $\alpha$ lines



NGC 4945: Marinucci et al. (2017)

Neutral Fe

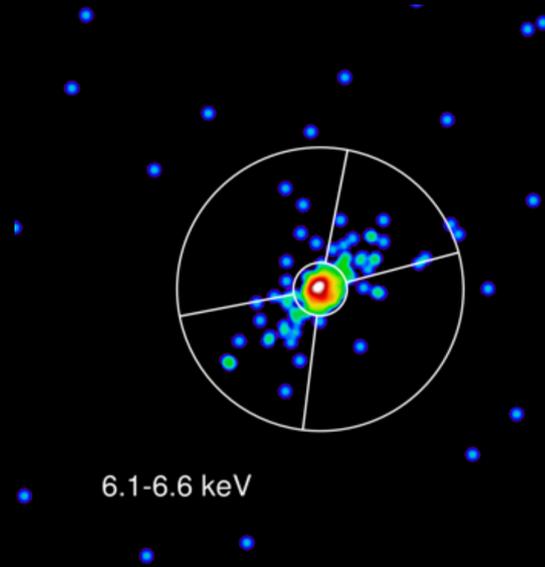


2.2'' (40 pc)

Ionized Fe

5'' (90 pc)

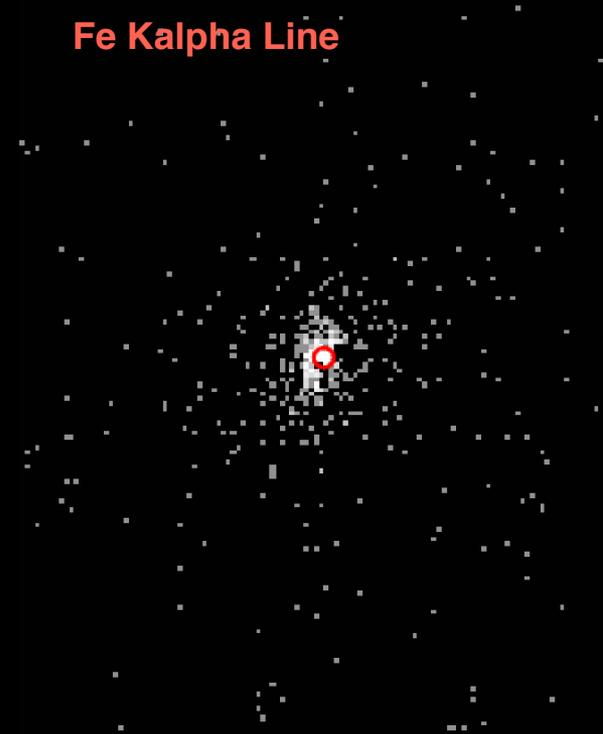
ESO428-G014: Fabbiano et al. (2017)



6.1-6.6 keV

NGC 2110: Bauer et al. (in prep.)

Fe K $\alpha$  Line

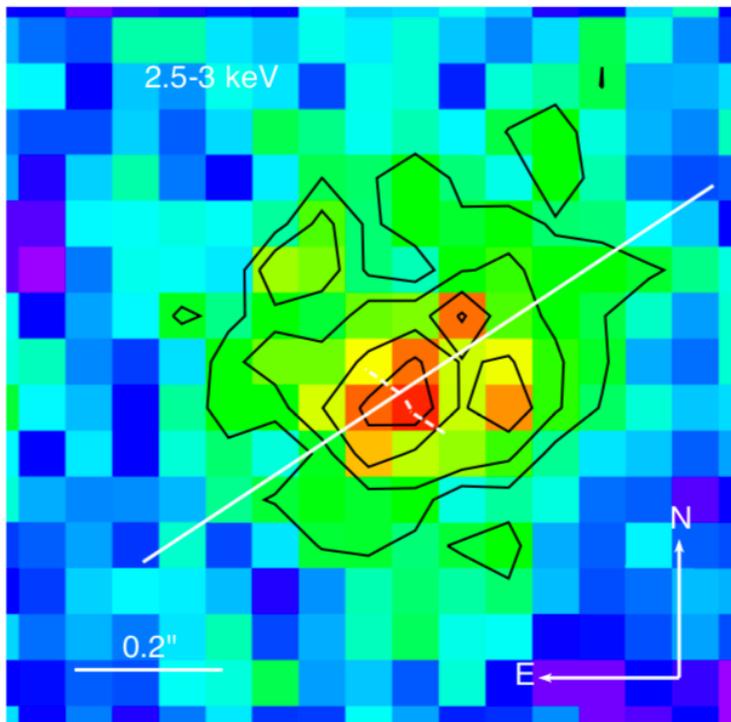


See talks by G. Fabbiano, M. Elvis, F. Bauer

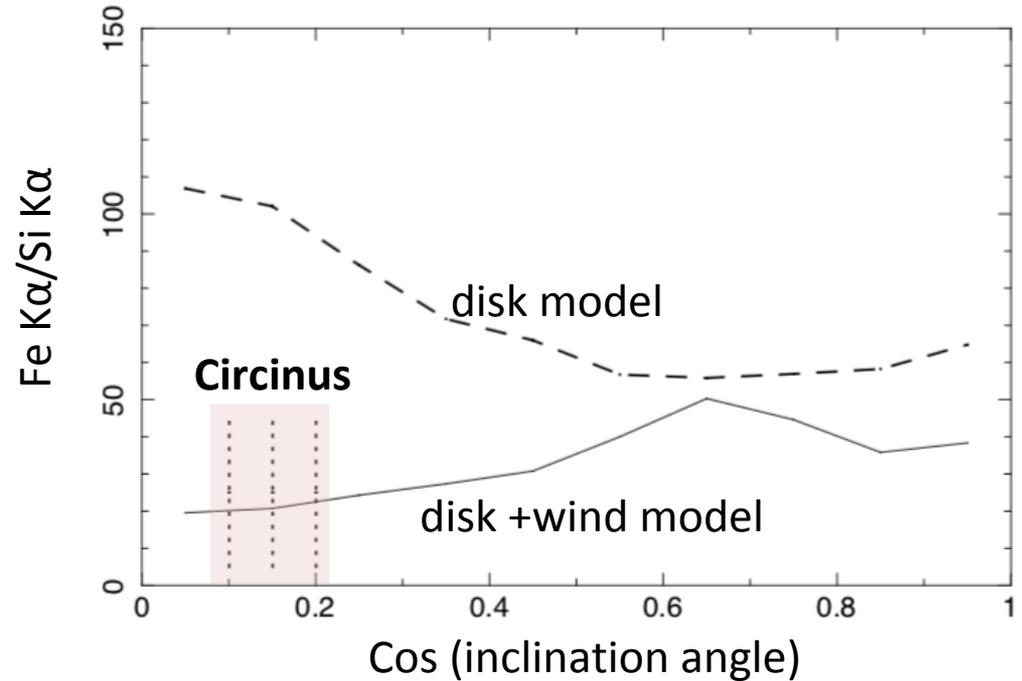
# Extended Si $K\alpha$ lines



Chandra 2.5-3 keV image of Circinus

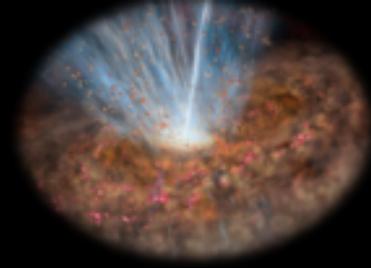


Simulations with two different models



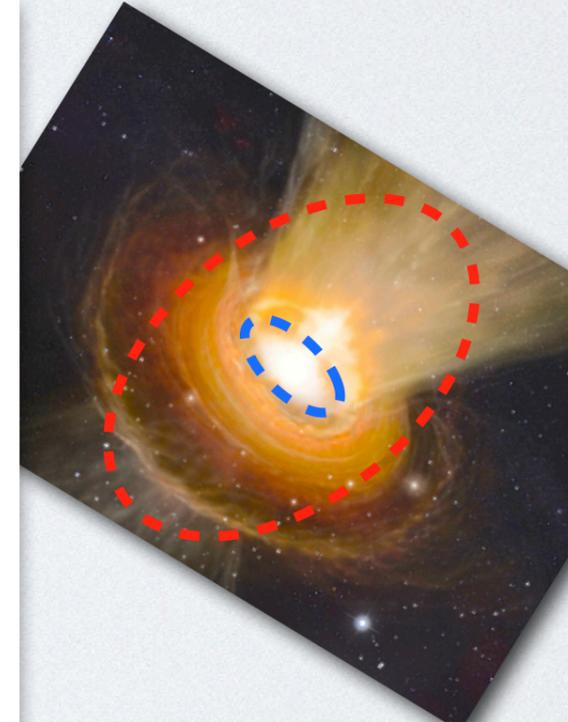
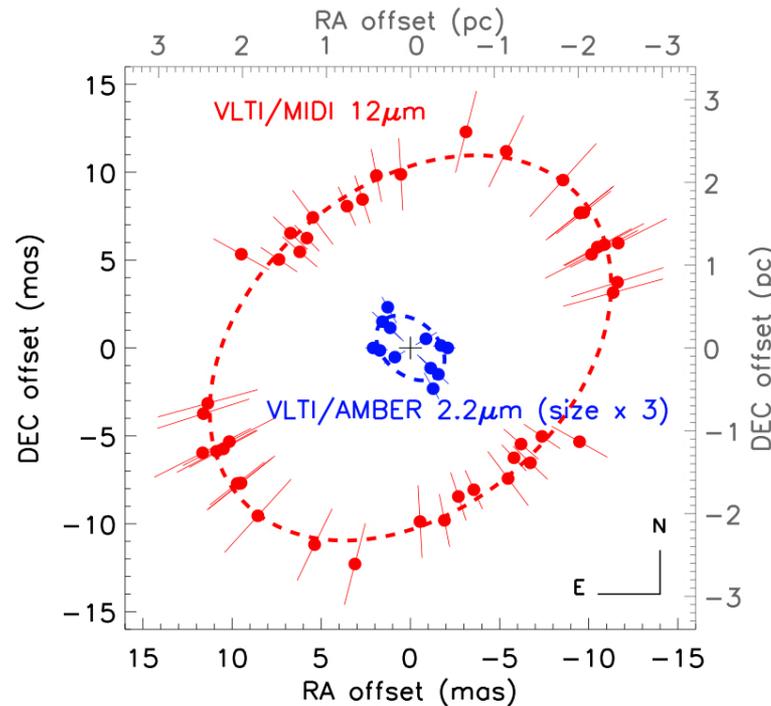
Liu, Hoenig, Ricci & Paltani (in prep.)

# Polar emission in the IR



polar direction /  
mid-IR

near-IR /  
accr. disk plane



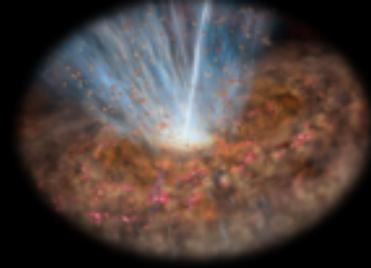
Hoening et al. (2013)

VLT: significant mid-IR emission from **polar region**

(Raban+09, Hoening+12,13, Burtscher+13, Tristram+14, Lopez-Gonzaga+16, Leftley+18)

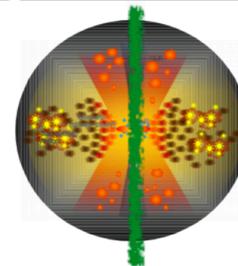
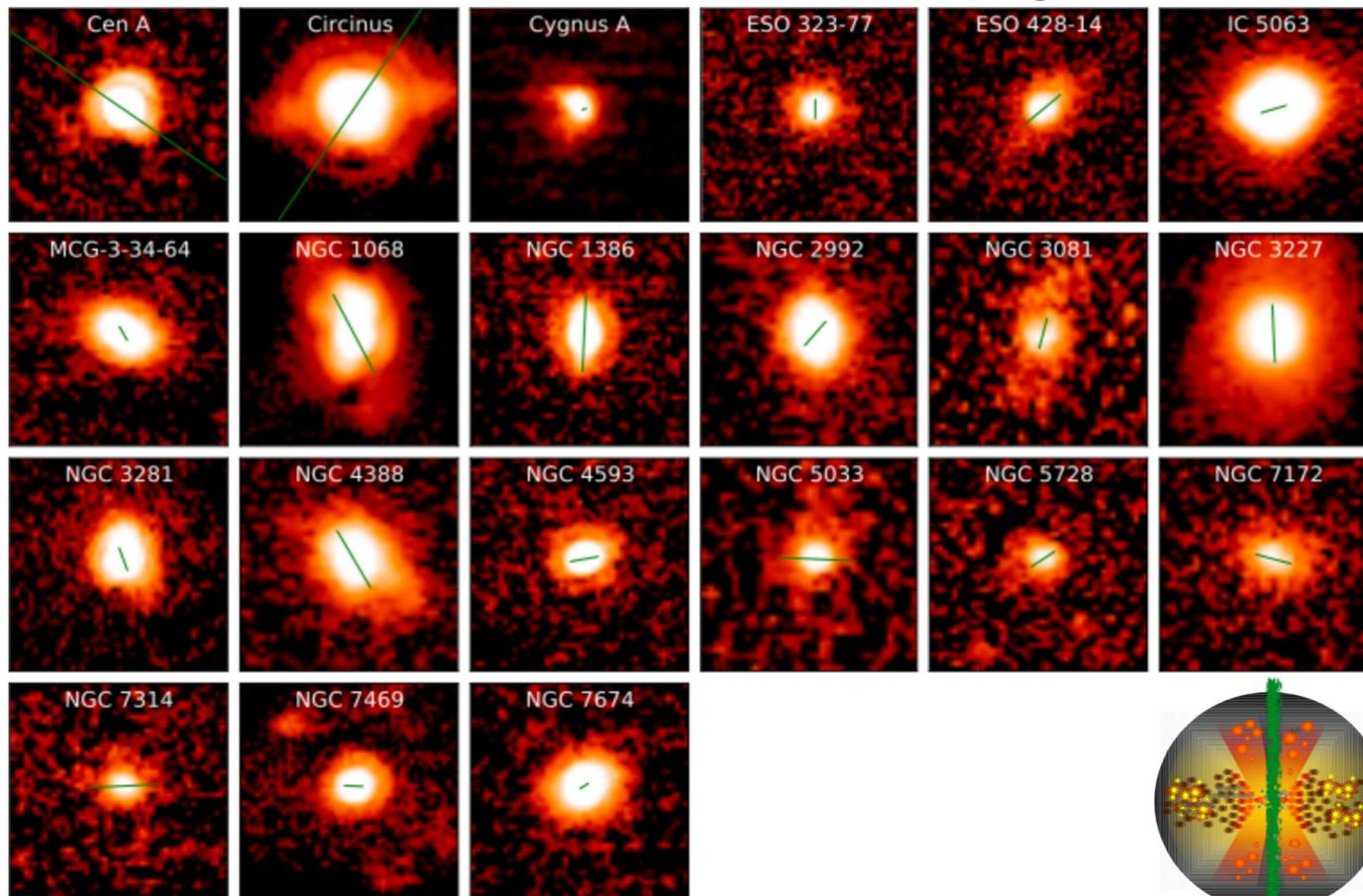
See also talks by D. Asmus, M. Stalevski; Poster by J. Leftley

# Polar emission in the IR

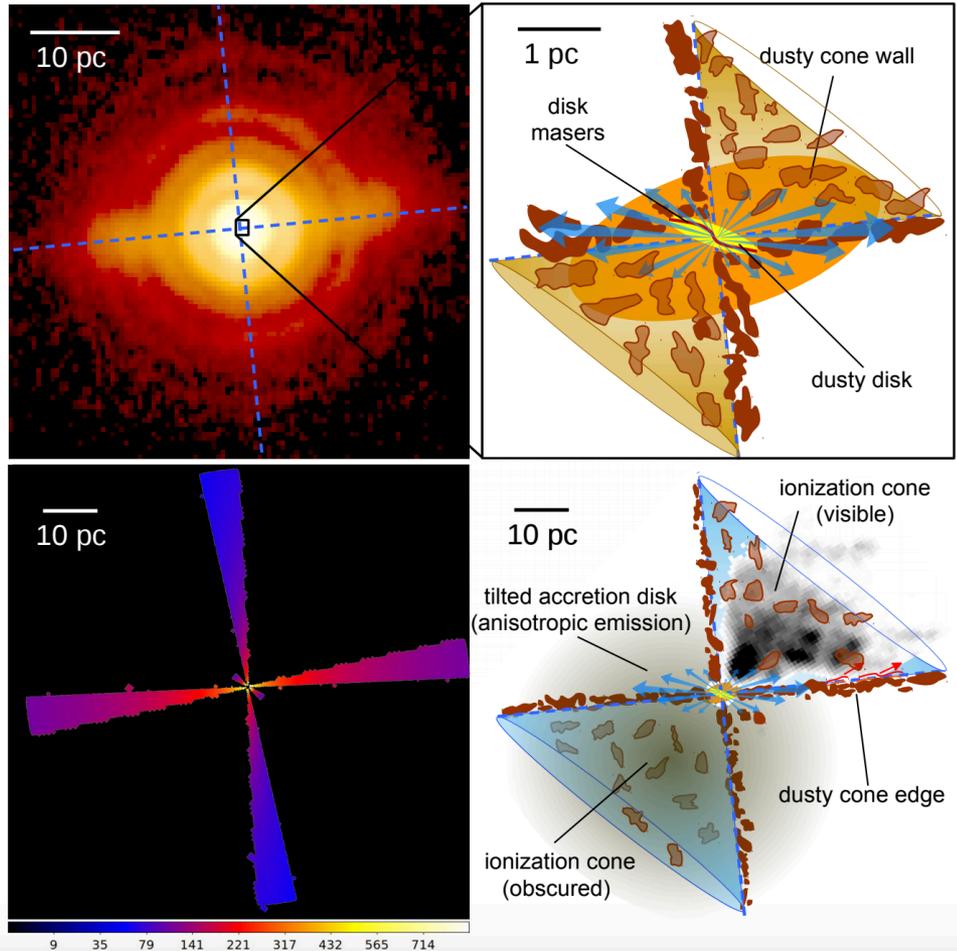
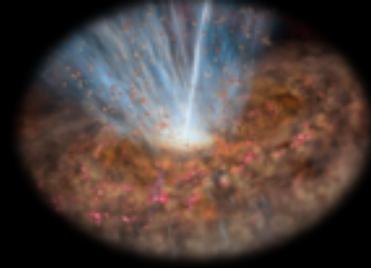


## VISIR images of local AGN

Asmus, Honig, Gandhi (2016)

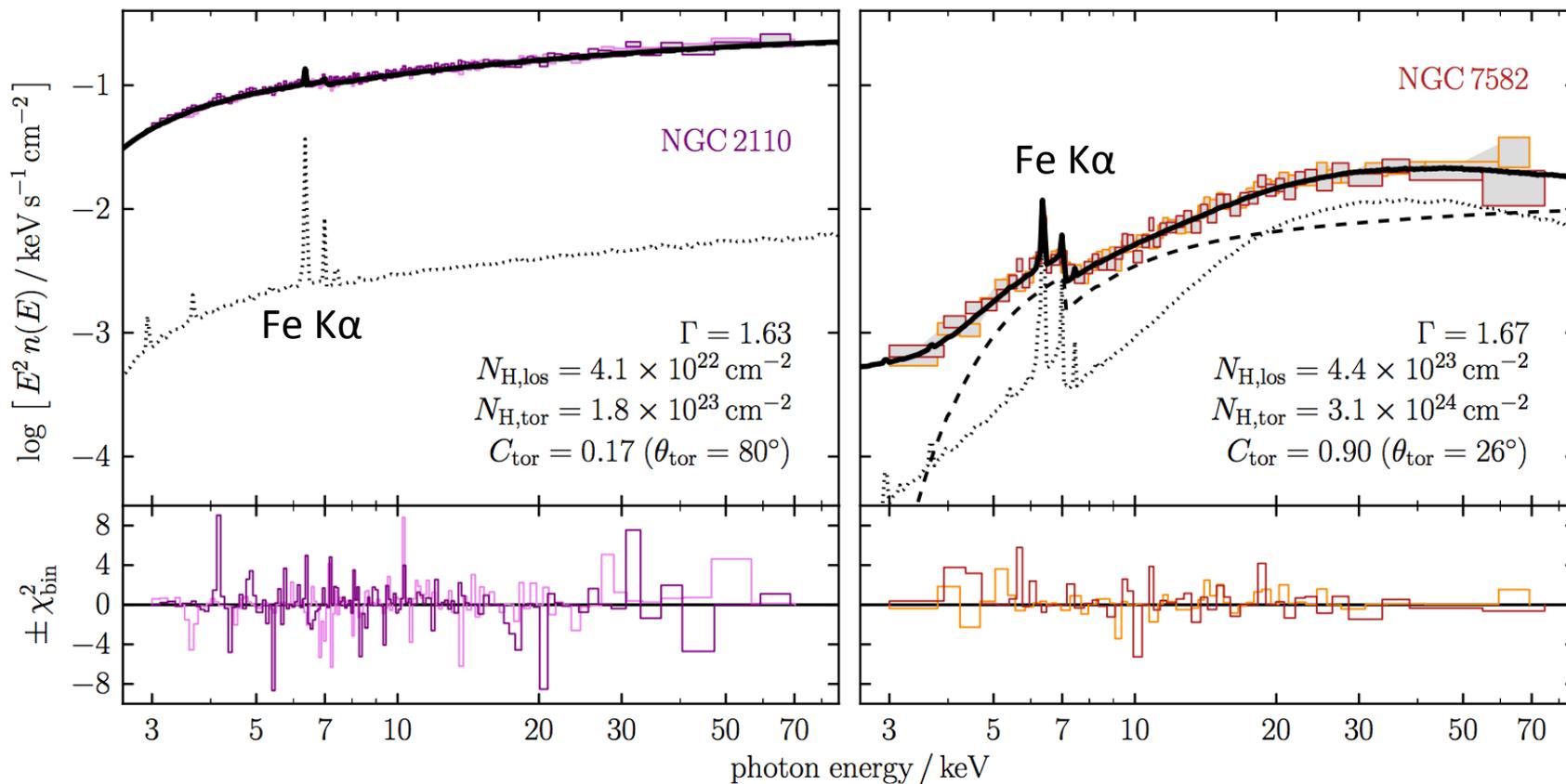


# Polar emission in the IR



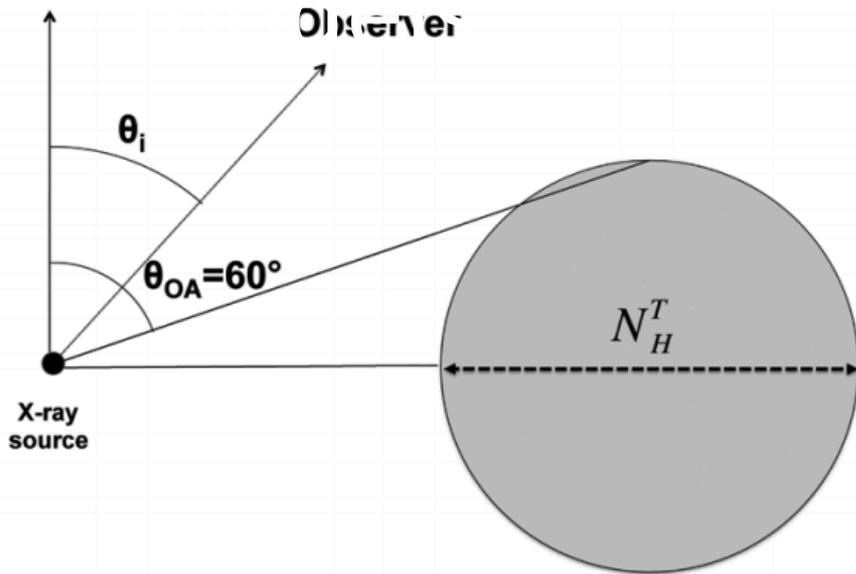
Stalevski, Asmus & Tristram (2017, 2019); see talks by D. Asmus, M. Stalevski; Poster by J. Lyu.

# Broad-band spectroscopy

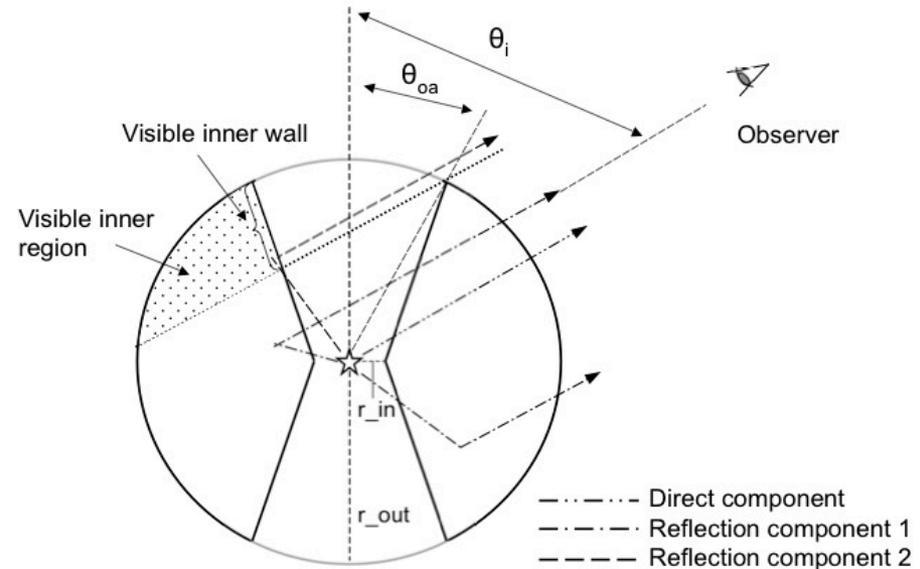


Balokovic et al. (2018); See talks by M. Balokovic, J. Buchner, N. Osorio-Clavijo; Poster by T. Kawamuro

# Reprocessed radiation: spectral models



*MyTORUS* (Murphy & Yaqoob 2009)

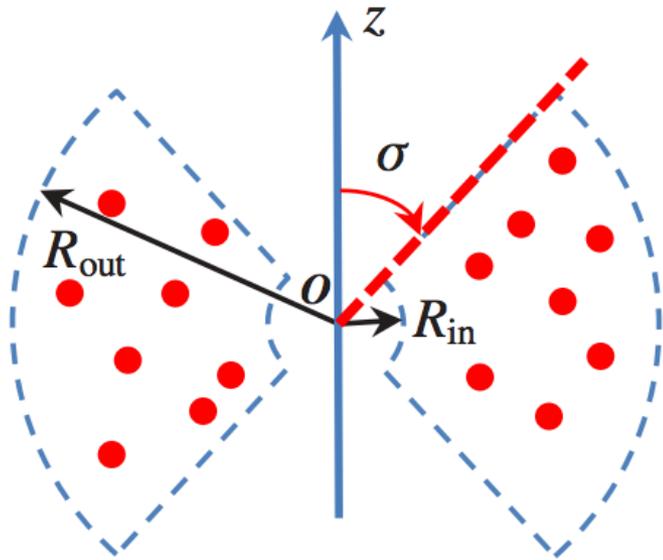


*Ikeda et al. (2009)*

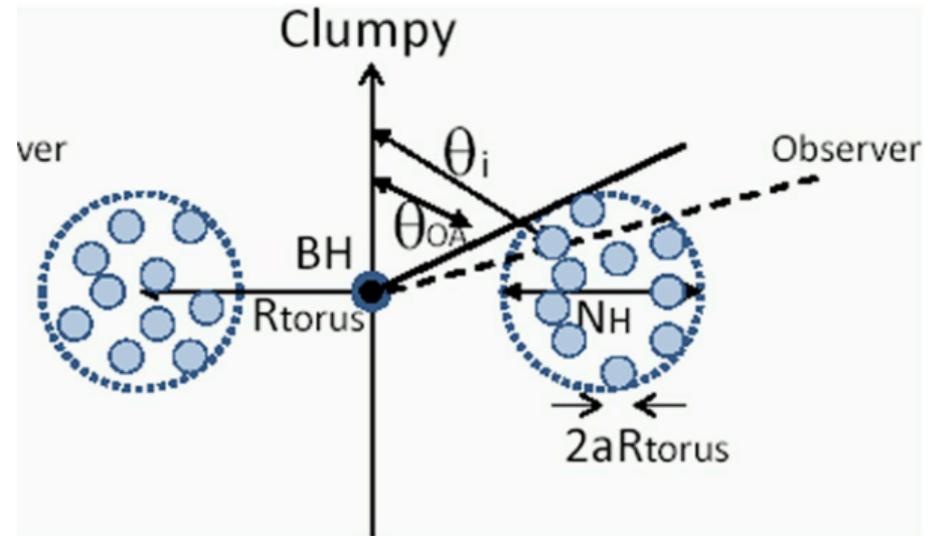
*BNTorus* (Brightman et al. 2011), *Borus* (Balokovic et al. 2017); Ikeda et al. (2009)

See talks by M. Balokovic, M. Elvis; Posters by A. Tanimoto, D. Esparza Arredondo

# Reprocessed radiation: spectral models



Liu & Li (2014)



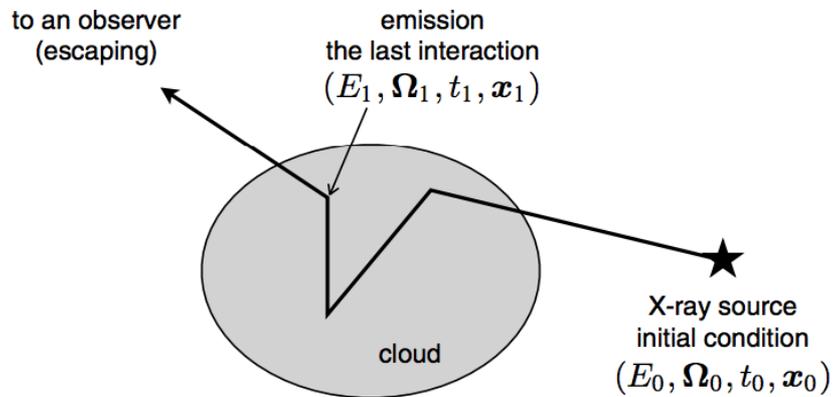
Furui et al. (2016)

See talks by M. Balokovic, M. Elvis; Posters by A. Tanimoto, D. Esparza Arredondo

# Reprocessed radiation: simulation platforms

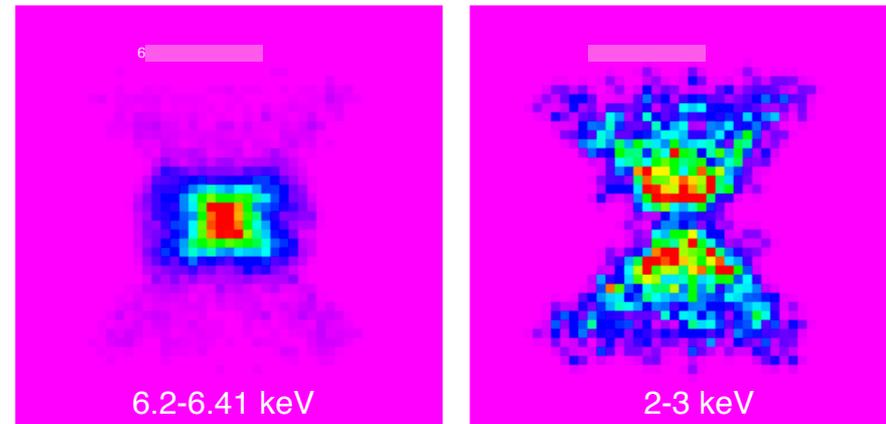
## MONACO (Geant4)

Odaka et al. (2011)



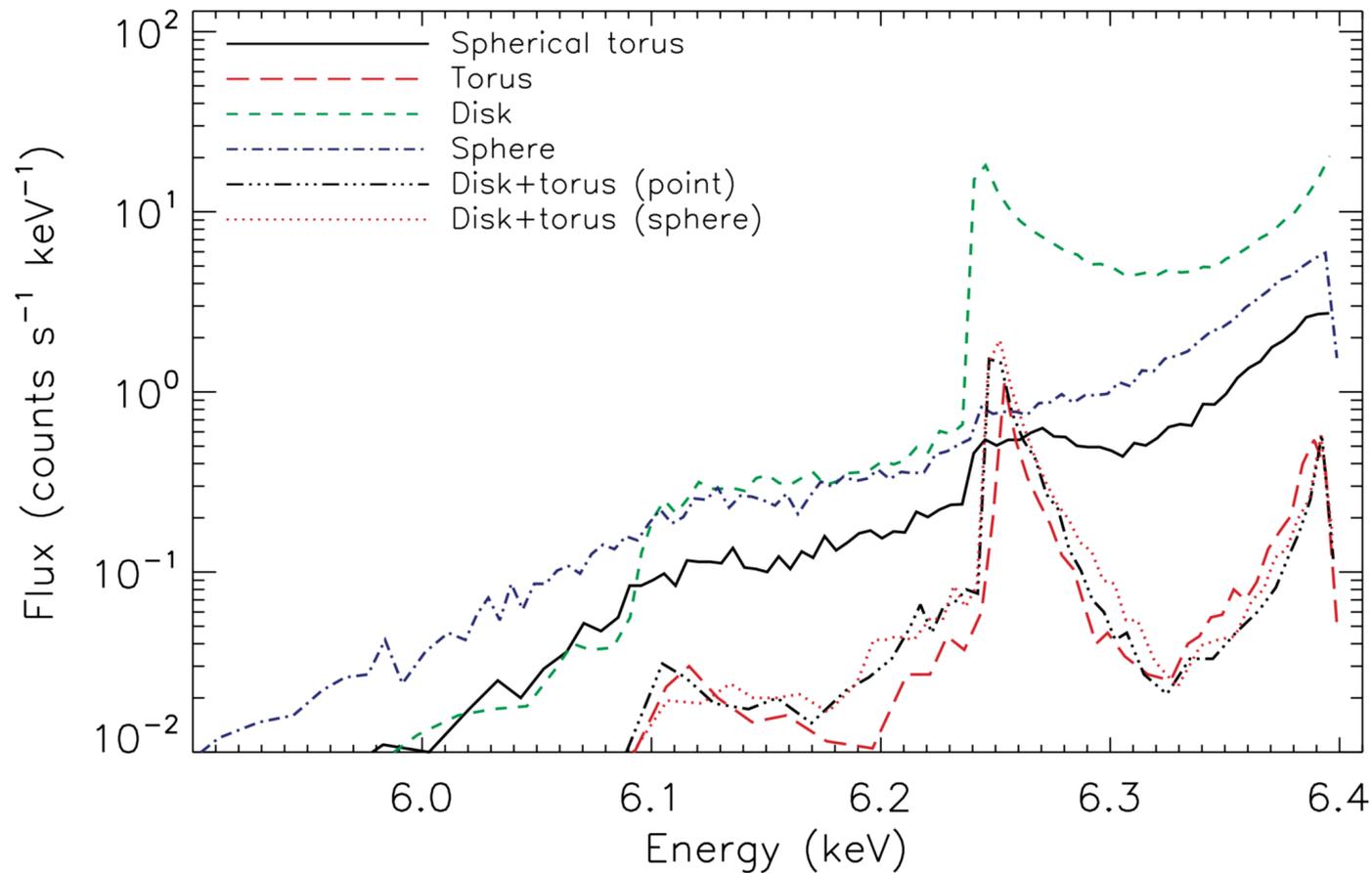
## RefleX

Paltani & Ricci (2017)



Liu et al. (in prep.)

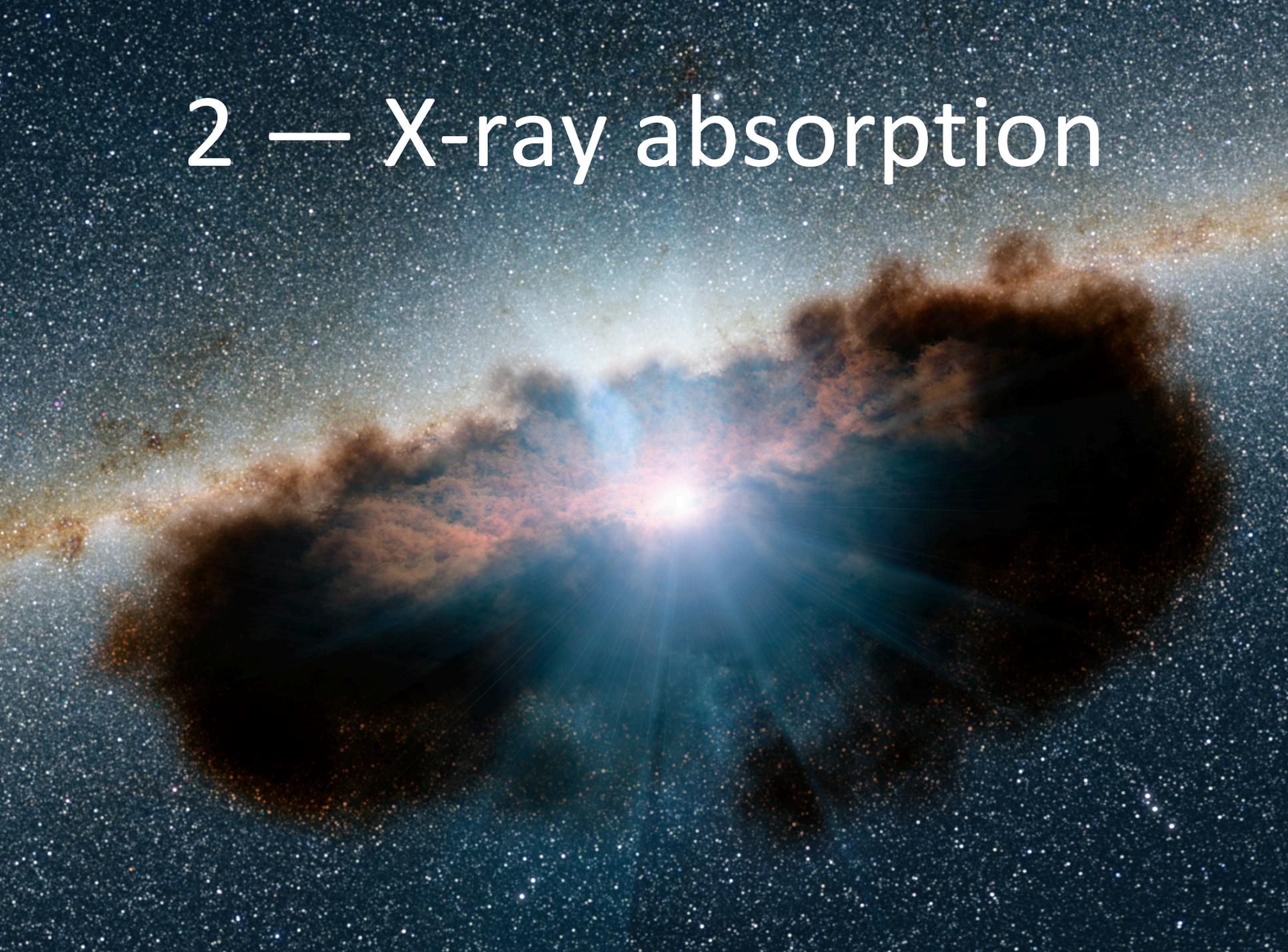
# Reprocessed radiation: simulation platforms



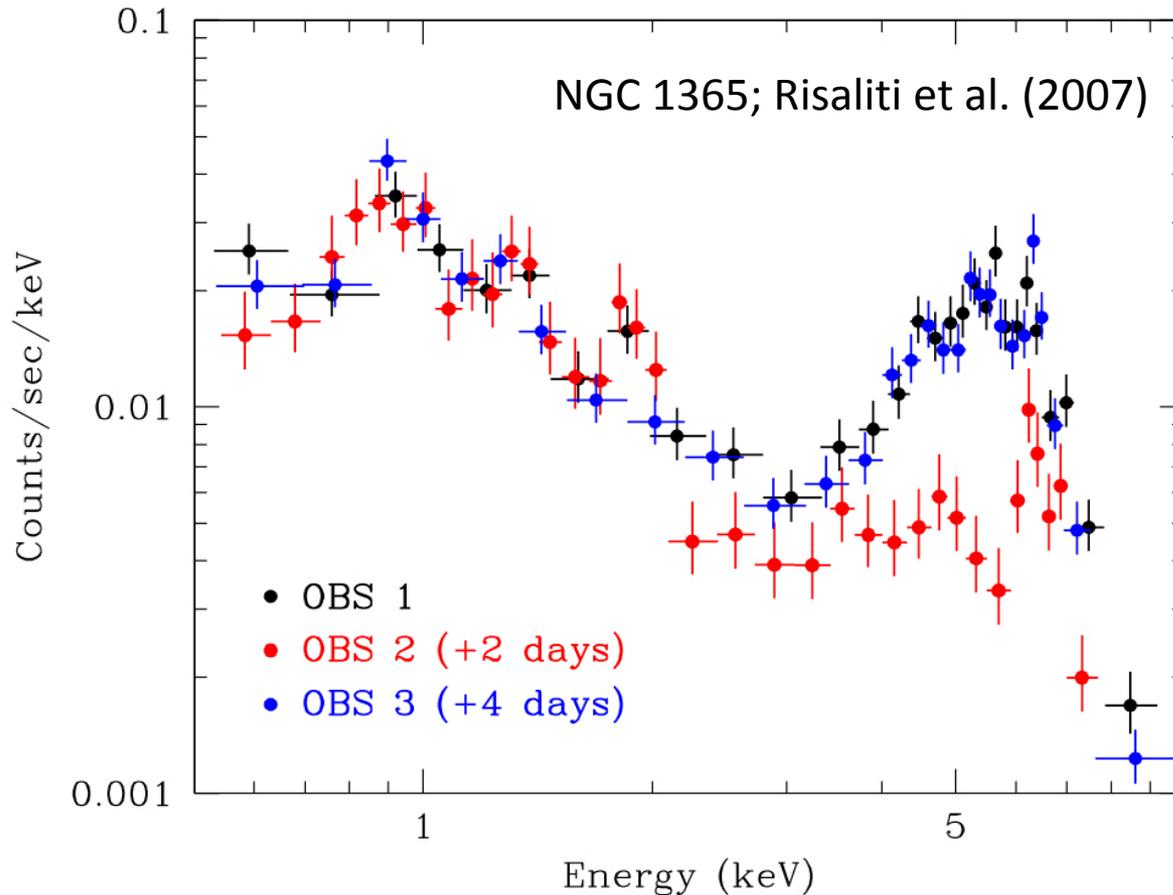
Ricci et al. (in prep.); see also Odaka et al. (2016)

See talk by C. Andonie

# 2 — X-ray absorption



# Obscuration variability

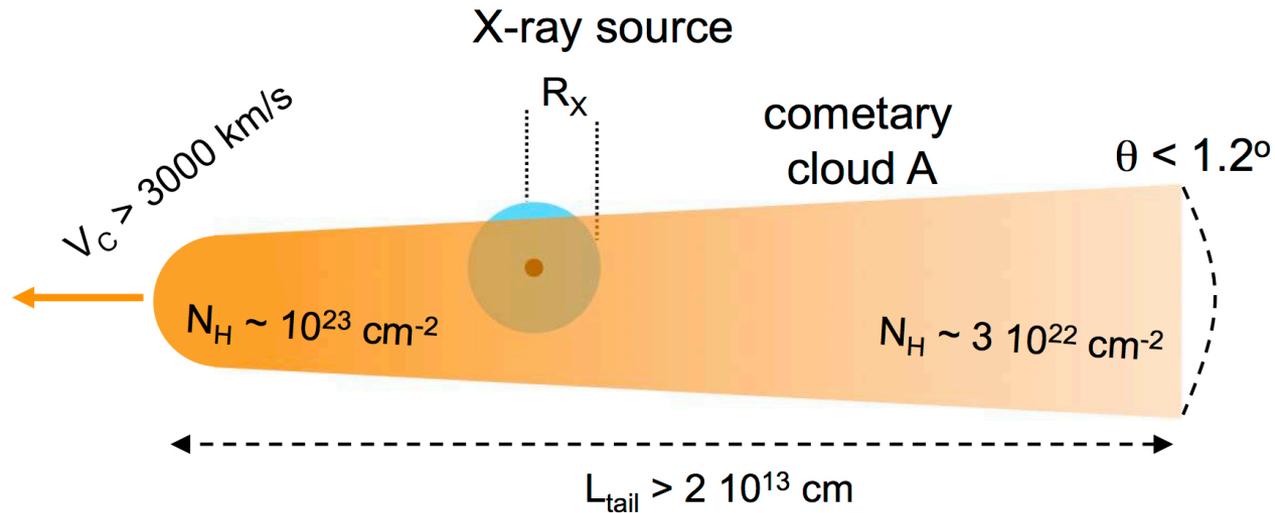


More than 20 AGN (*e.g.*, Risaliti+05, Bianchi+09, Marinucci+13, Miniutti+14, Rivers+15, Burtscher+15, Ricci+16b); See talk by S. Lamassa

# Obscuration variability



## Broad-line regions clouds?



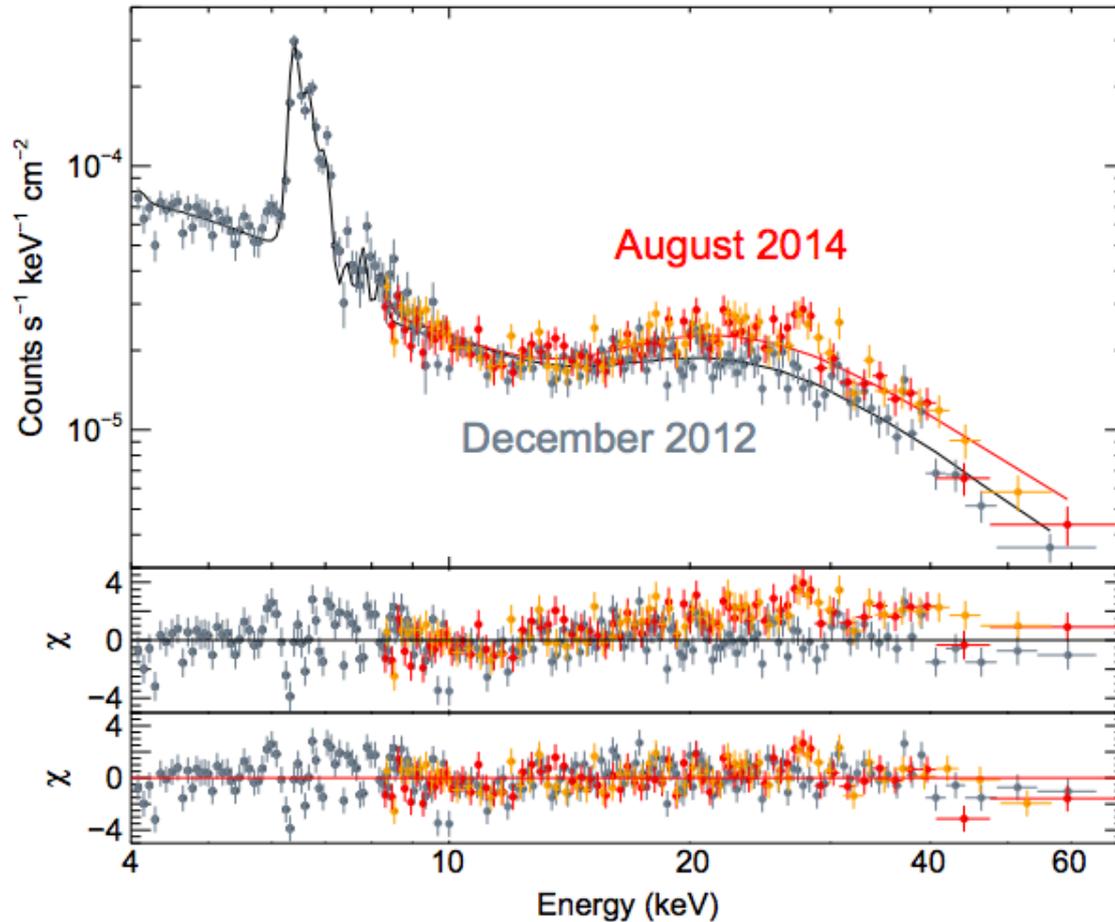
Maiolino et al. (2010)

See talk by S. Lamassa, E. Sturm, M. R. Stock, A. Laor, T. Waters

# Obscuration variability

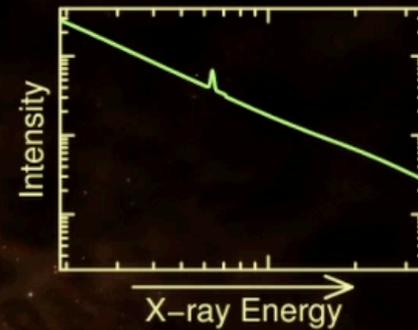


## NGC 1068

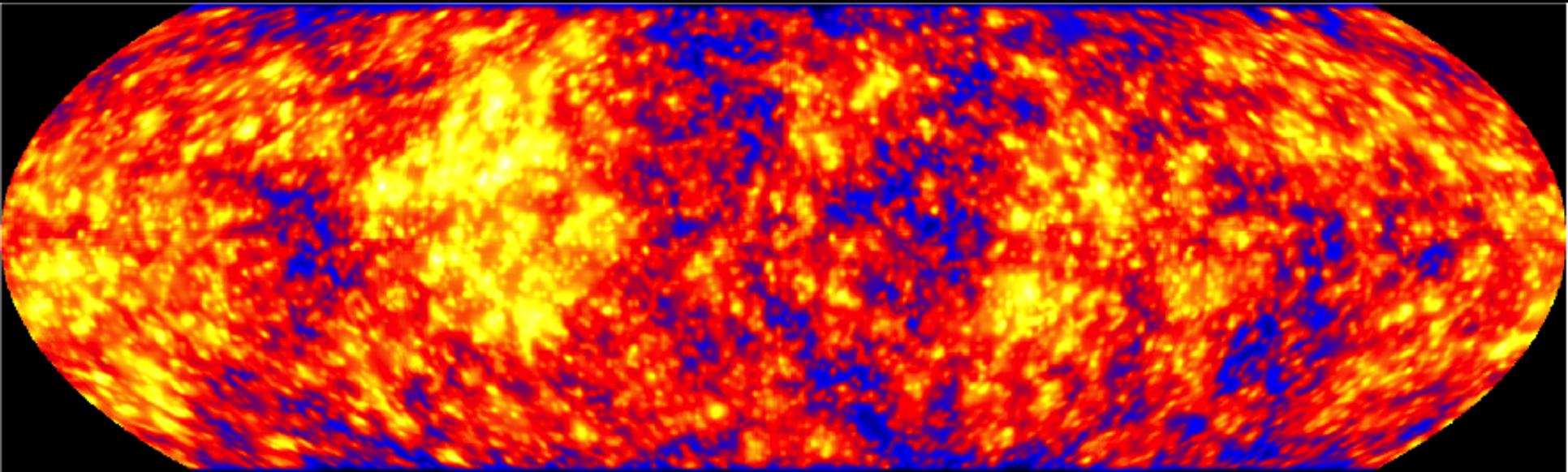


Marinucci et al. (2016); See talk by A. Zaino

# Obscuration variability



# The sky seen from the SMBH



Optical depth in the V band

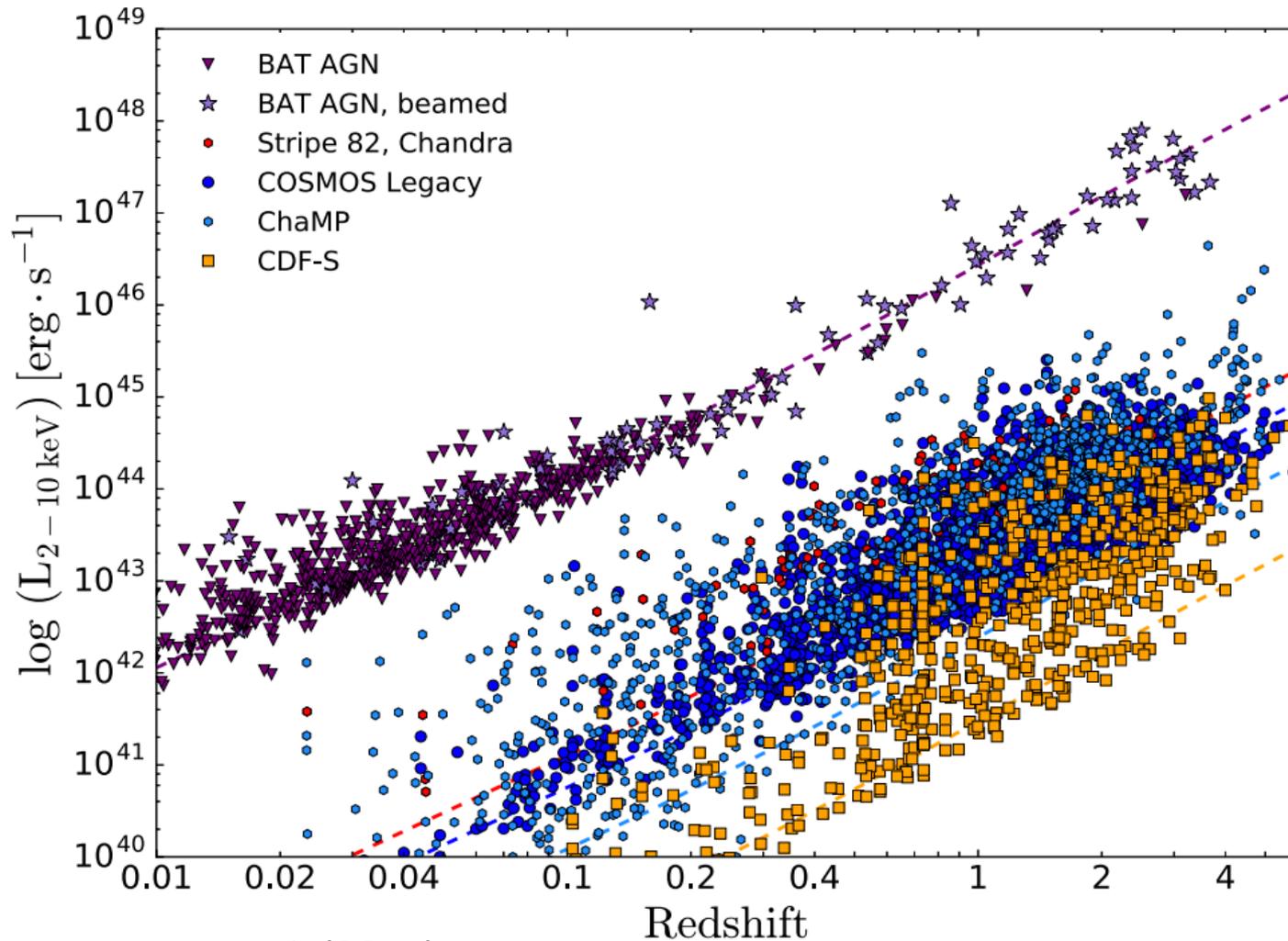
50

100

200

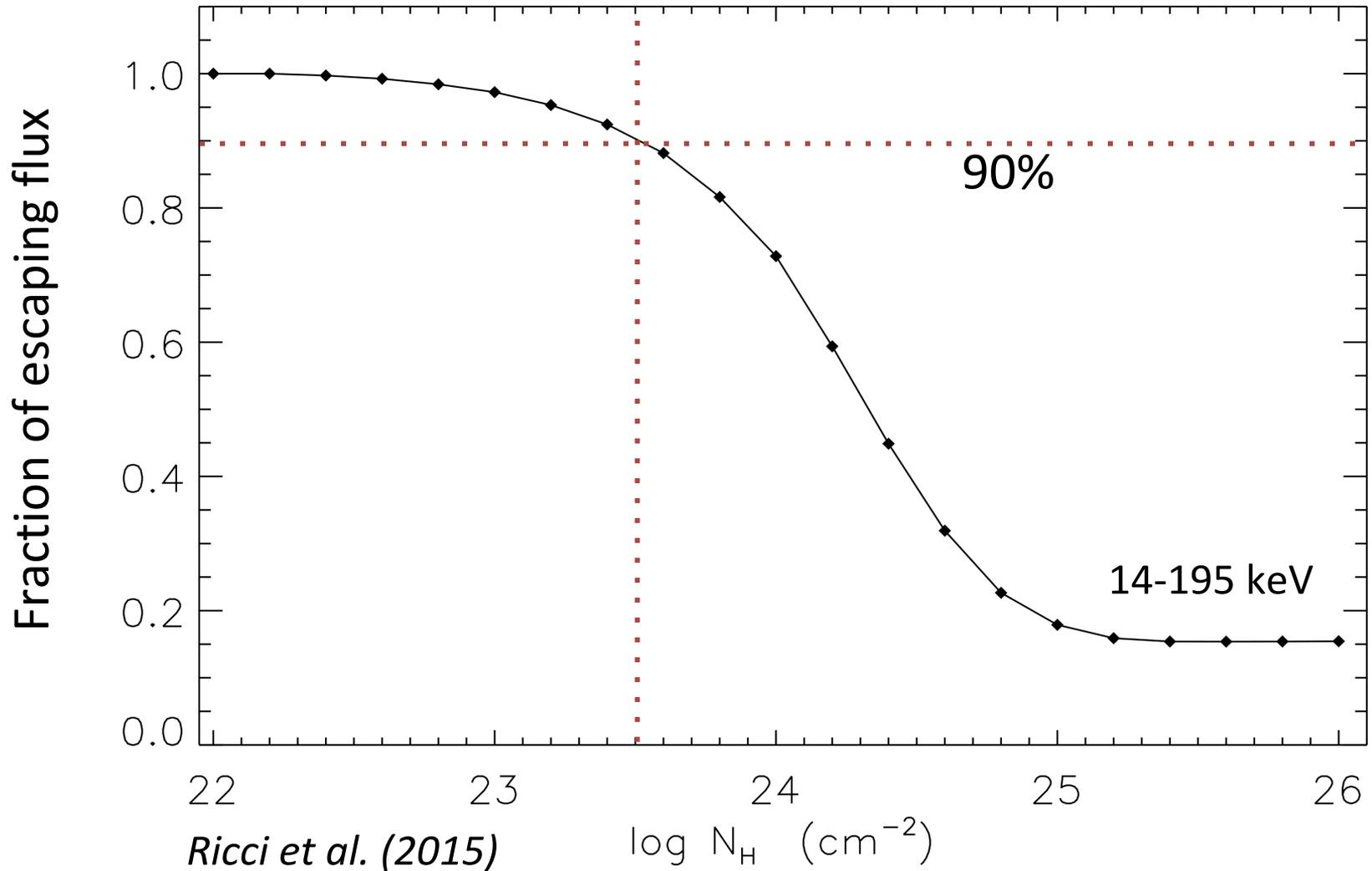
Stalevski et al. (2012); See talk by R. Nikutta, ; Poster by E. Lopez-Rodriguez

# X-ray Surveys

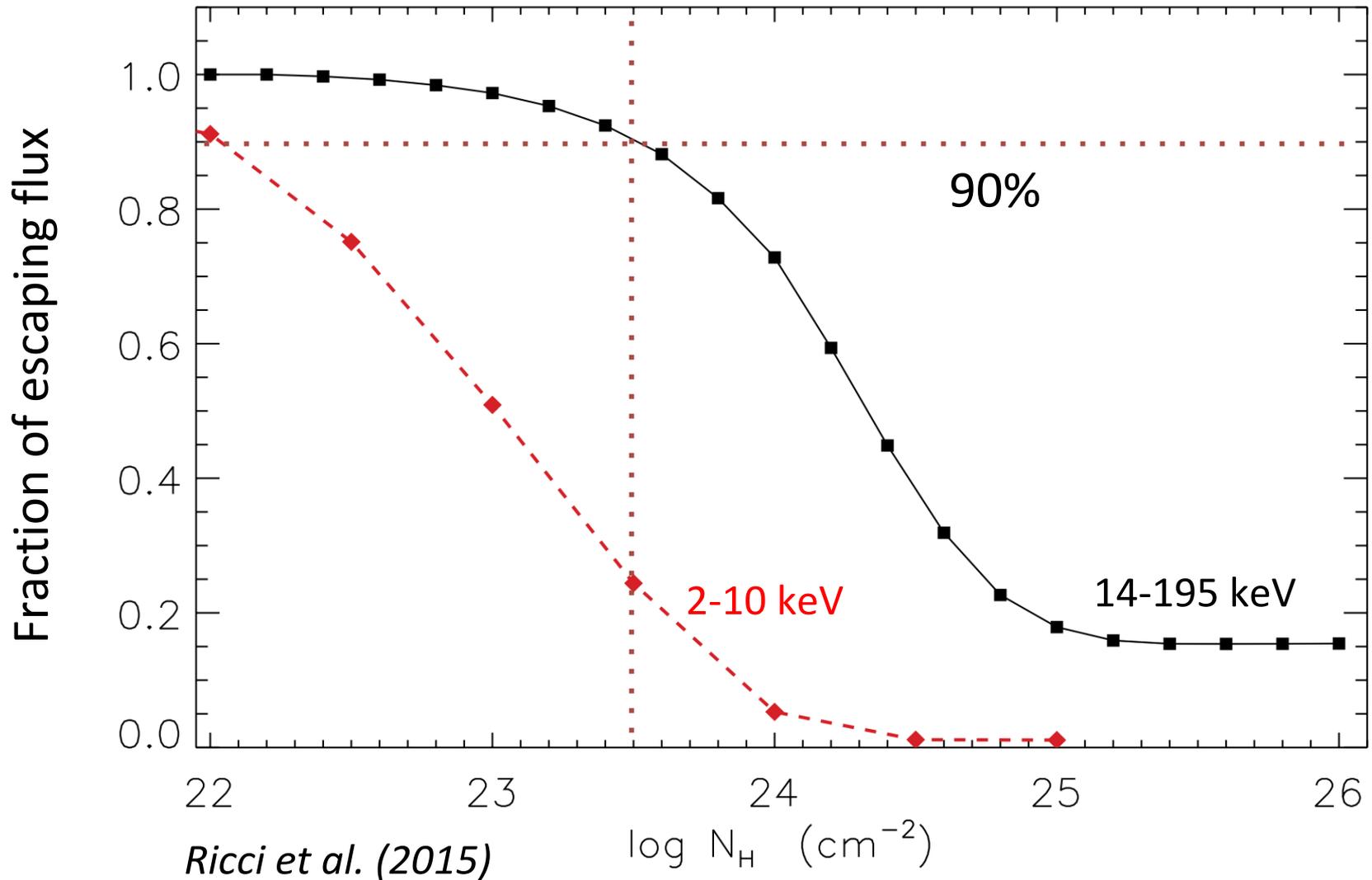


Koss et al. (2017)

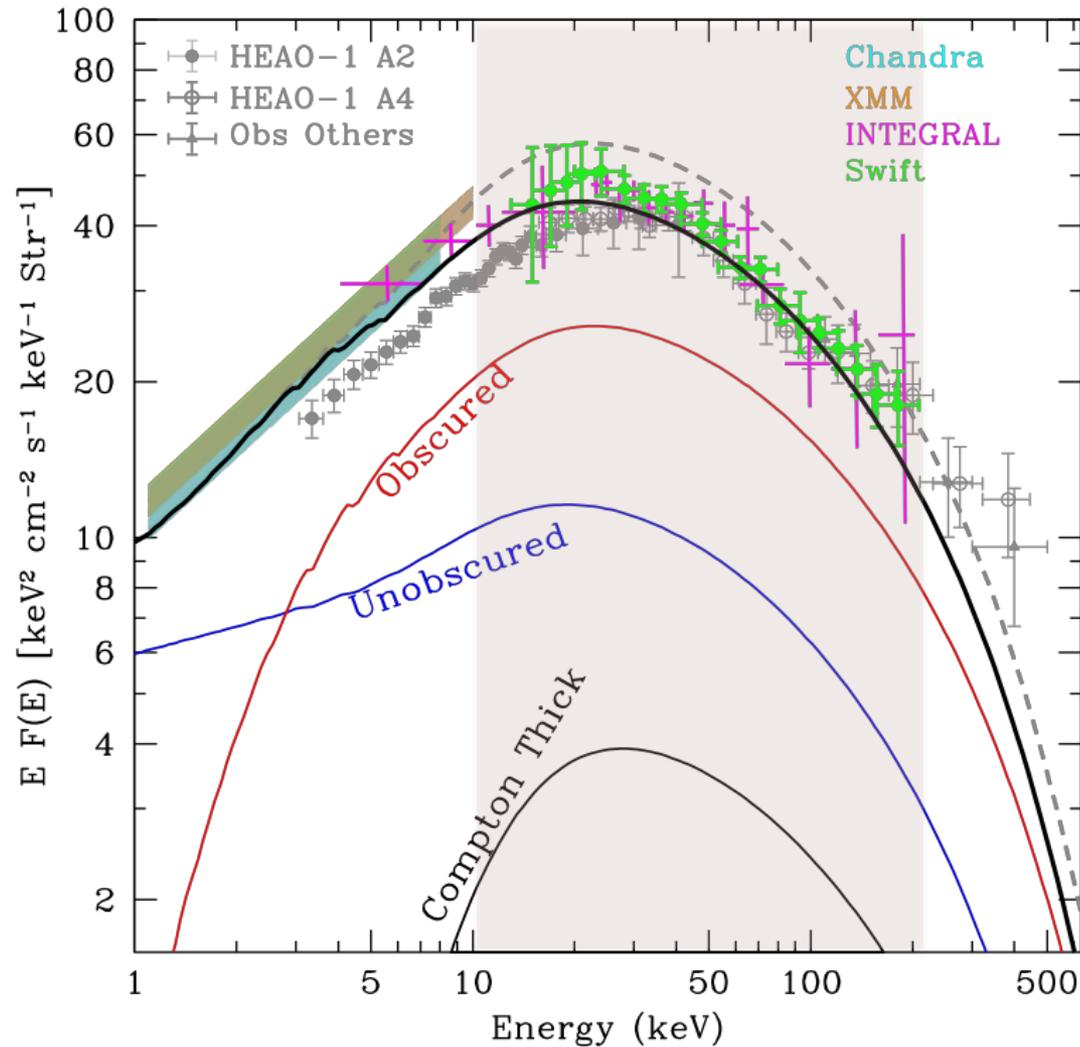
# Obscured AGN in the hard X-ray band



# Obscured AGN in the hard X-ray band



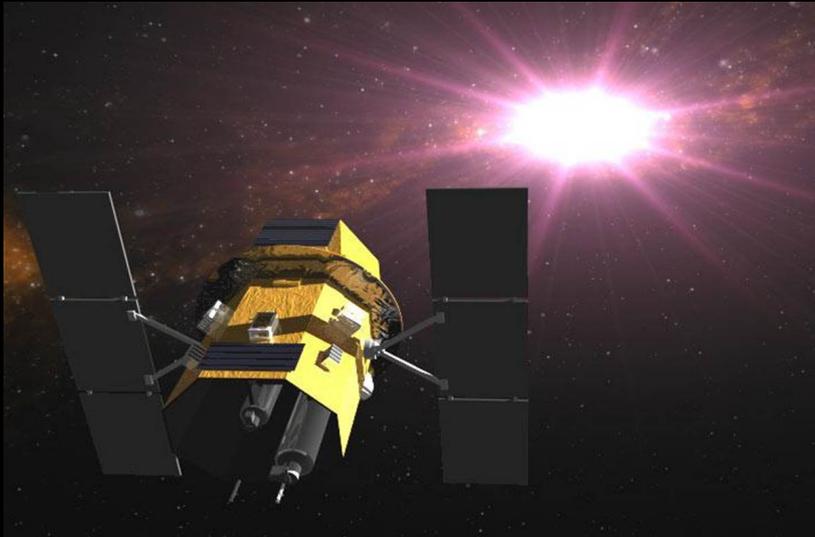
# Obscured AGN in the hard X-ray band



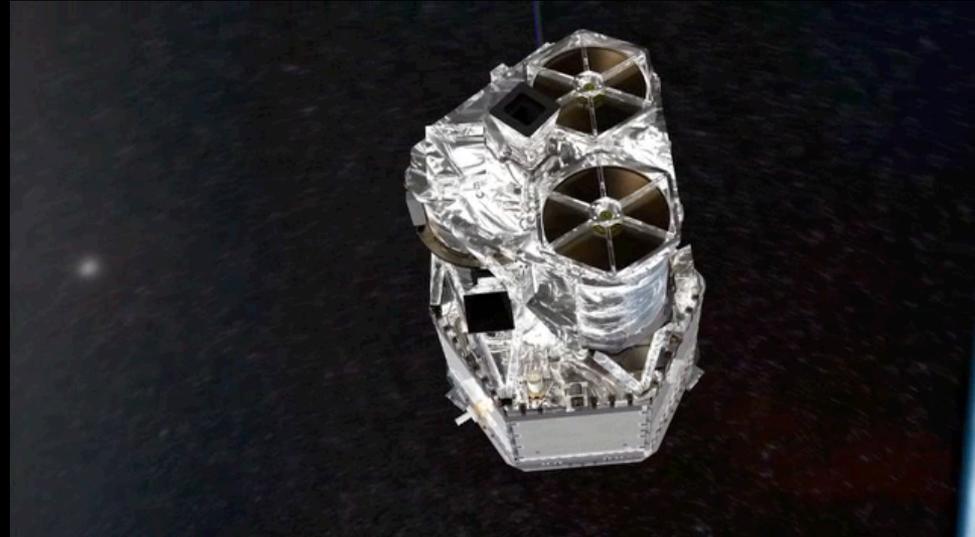
Treister et al. (2009)

# Obscured AGN in the hard X-ray band

Swift/BAT  
(14-195 keV)

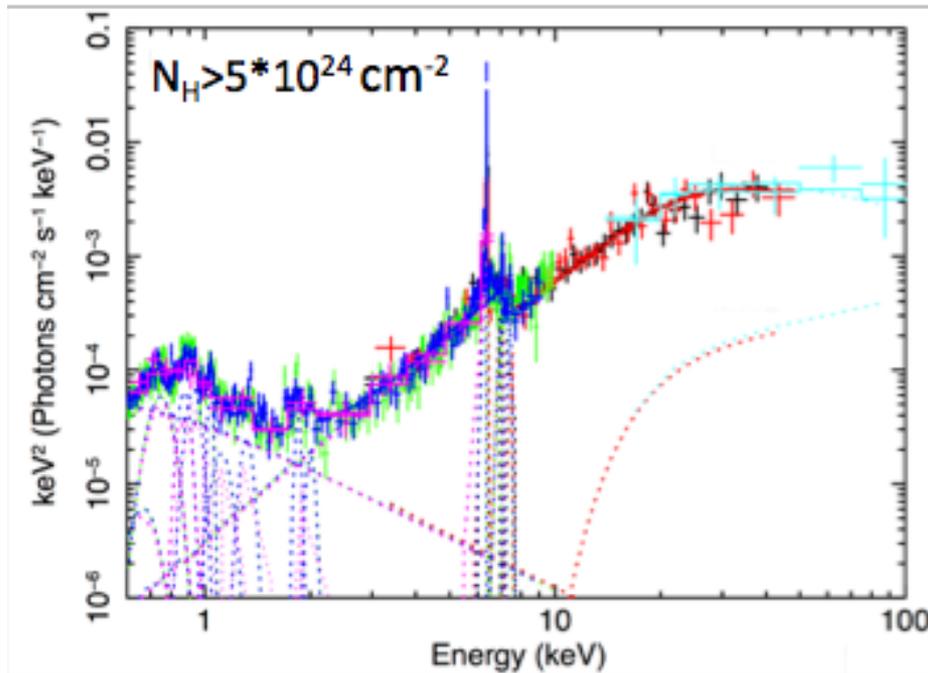


NuSTAR  
(3-80 keV)



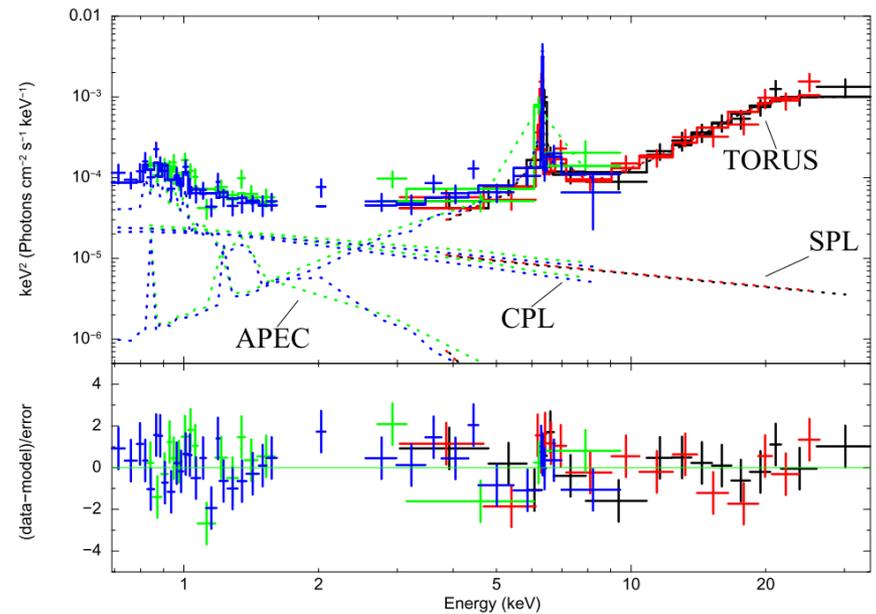
# Obscured AGN in the hard X-ray band

## NGC 5643



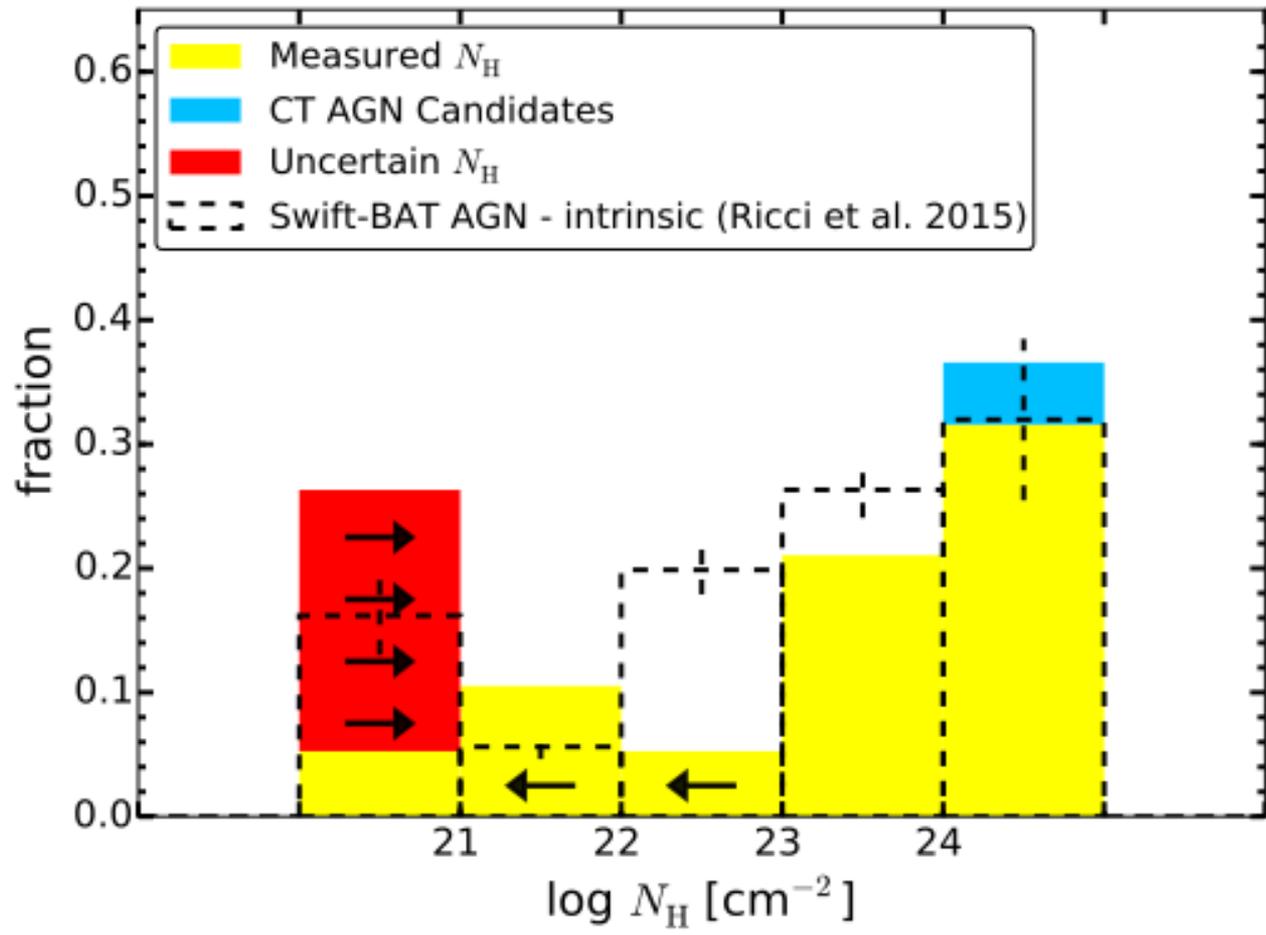
*Annuar et al. (2015)*

## IC 3639



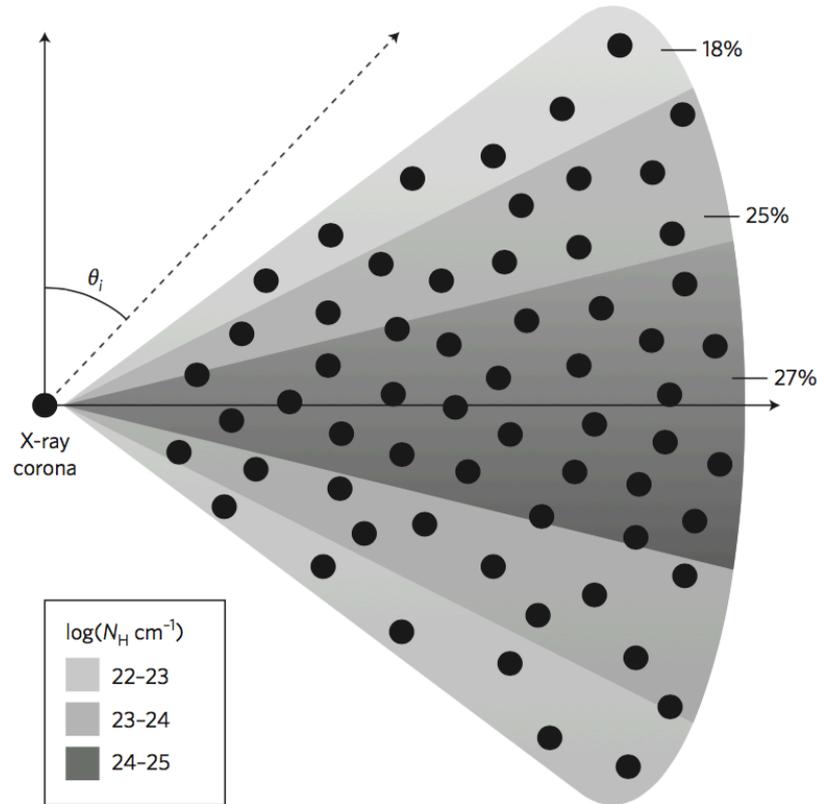
*Boorman et al. (2016)*

# The Covering factor of gas and dust



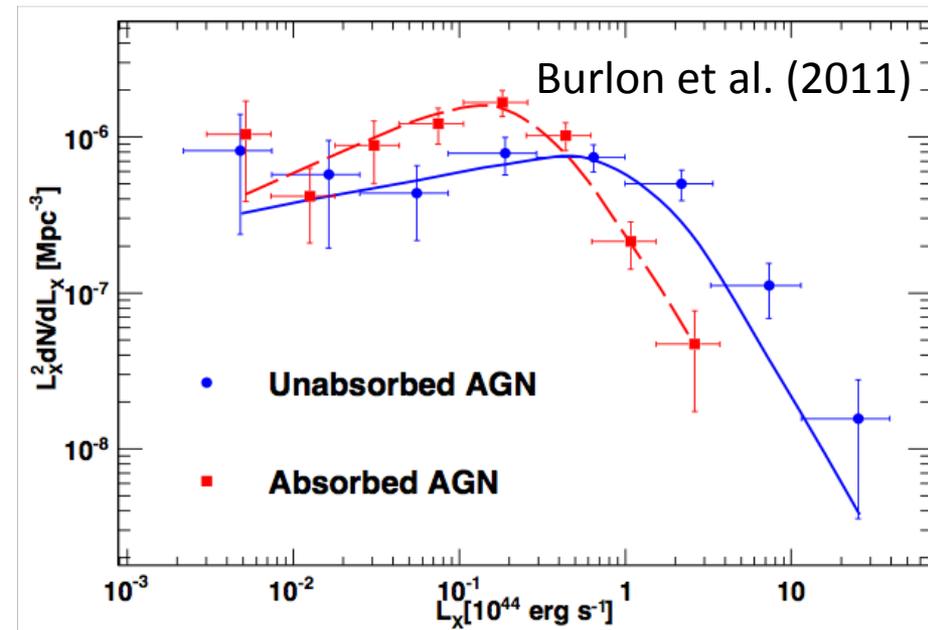
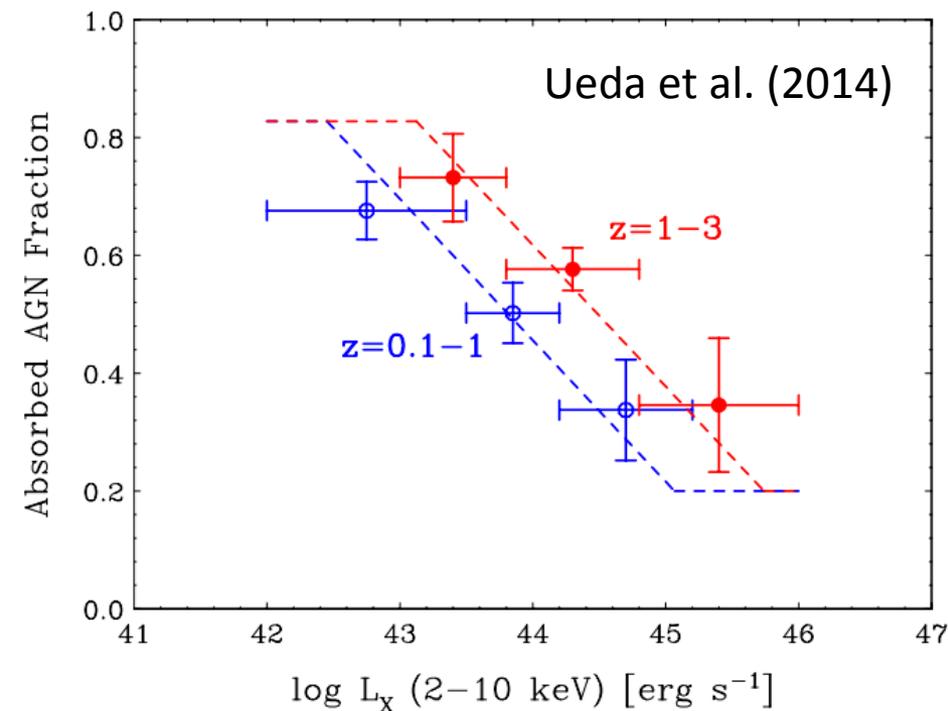
Annuar et al. (in prep.),

# The Covering factor of gas and dust



Ramos Almeida & Ricci, Nature Astronomy 2017, see also Ricci et al. (2015);  
talks by P. Boorman, K. Ichikawa, S. Mateos, L. Lanz

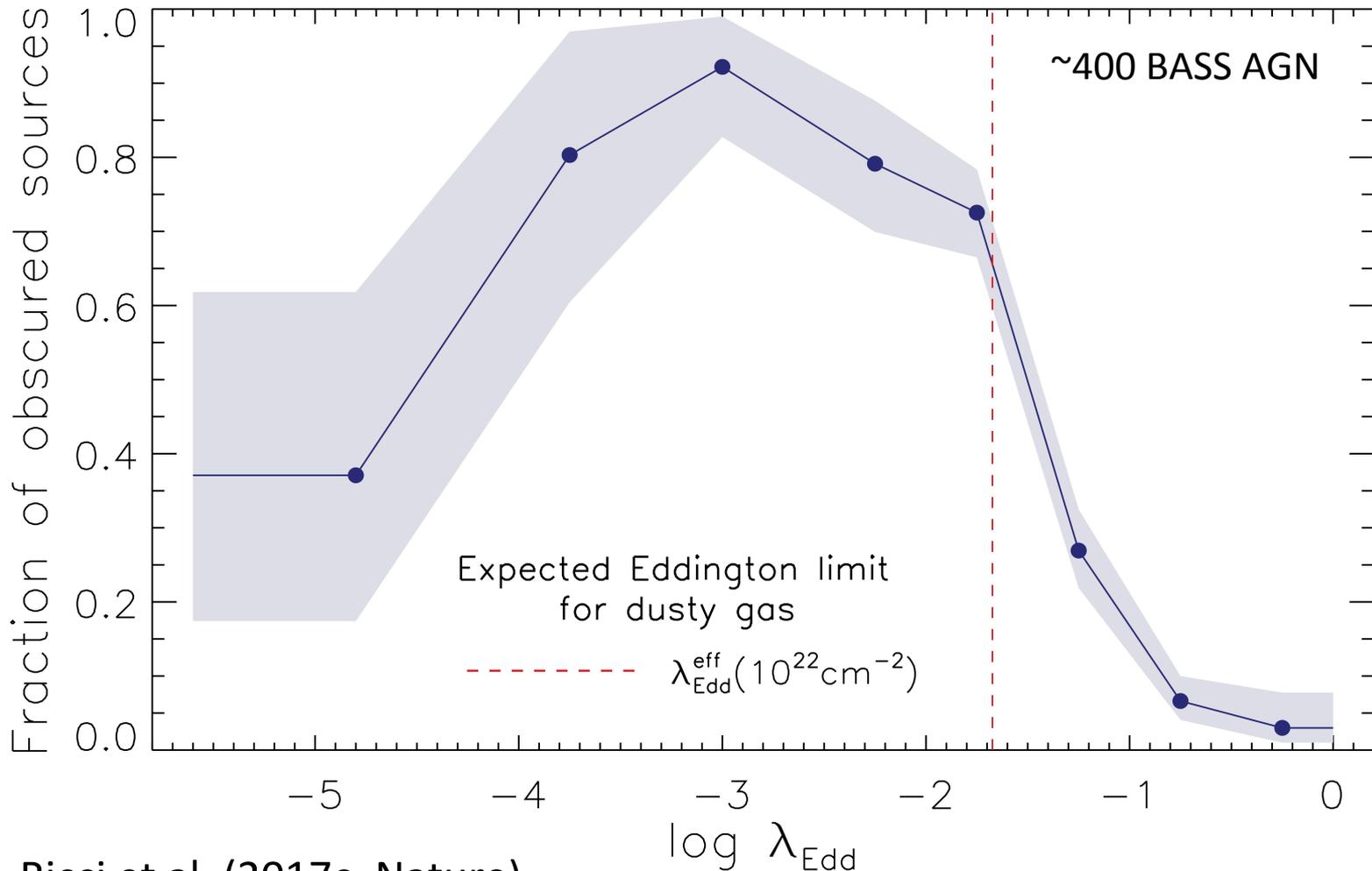
# Luminosity-dependent obscuration



see also Ueda et al. (2003, 2011), LaFranca+05, Sazonov+07, Hasinger 08, DellaCeca+08, Beckmann+09, Brightman+11, Merloni+14, Buchner+15, Aird+15, Sazonov+16, Georgakakis+17, Mateos+17, Ricci+17

See talks by C. Ramos Almeida, K. Ichikawa, S. Mateos, L. Lanz

# Obscuration vs Eddington ratio



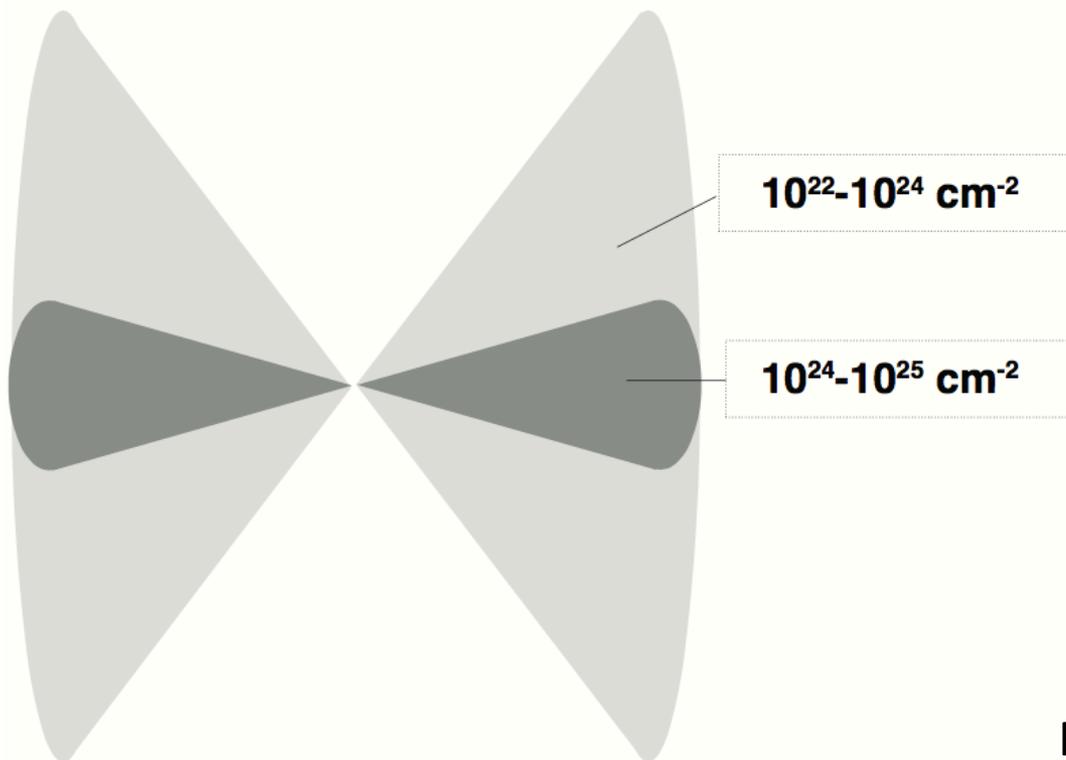
Ricci et al. (2017c, Nature)

# Radiation-regulated unification

## Low Eddington Ratio

$$(10^{-4} < \lambda_{\text{Edd}} < 10^{-1.5})$$

Covering factor  $\sim 85\%$

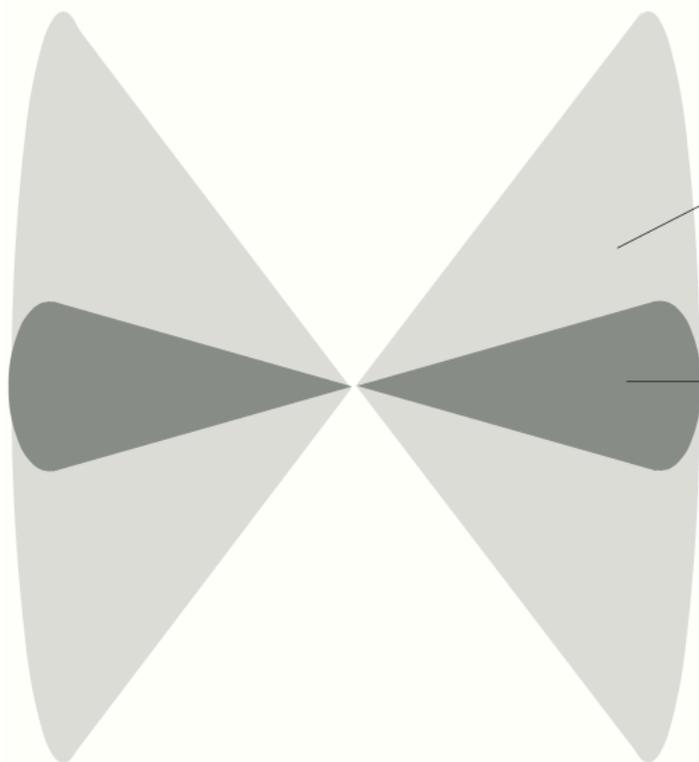


# Radiation-regulated unification

## Low Eddington Ratio

$$(10^{-4} < \lambda_{\text{Edd}} < 10^{-1.5})$$

Covering factor  $\sim 85\%$



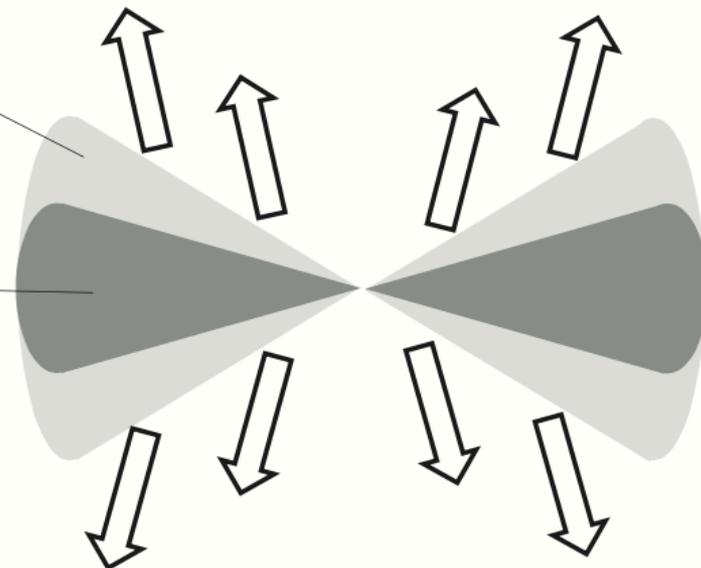
$10^{22}-10^{24} \text{ cm}^{-2}$

$10^{24}-10^{25} \text{ cm}^{-2}$

## High Eddington Ratio

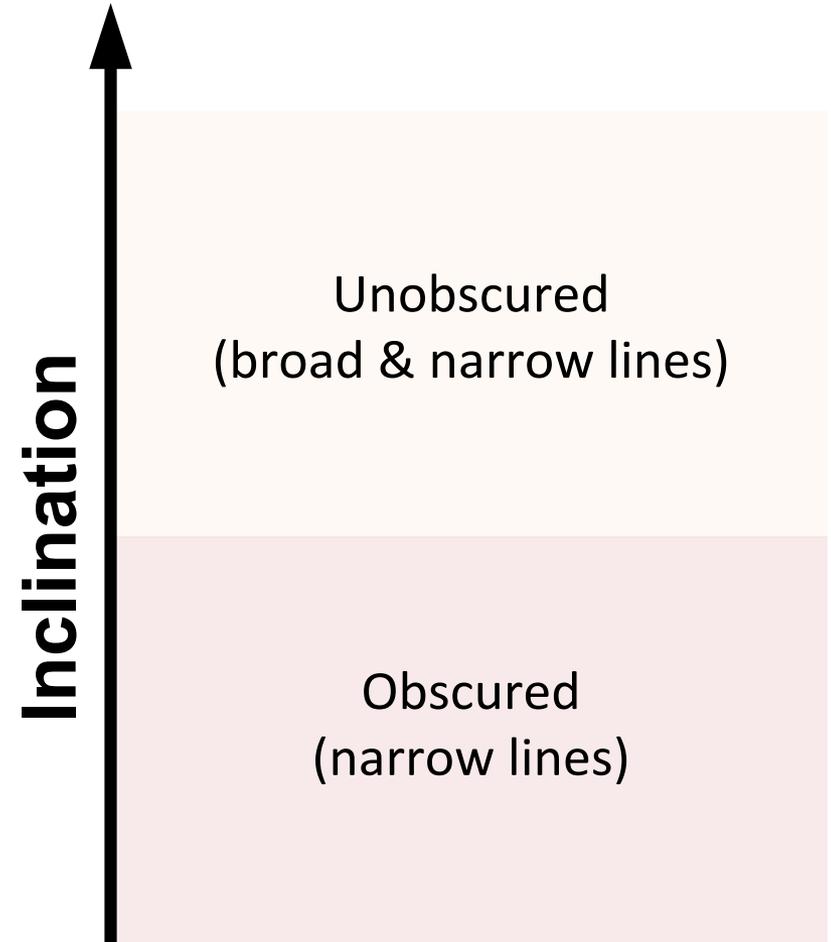
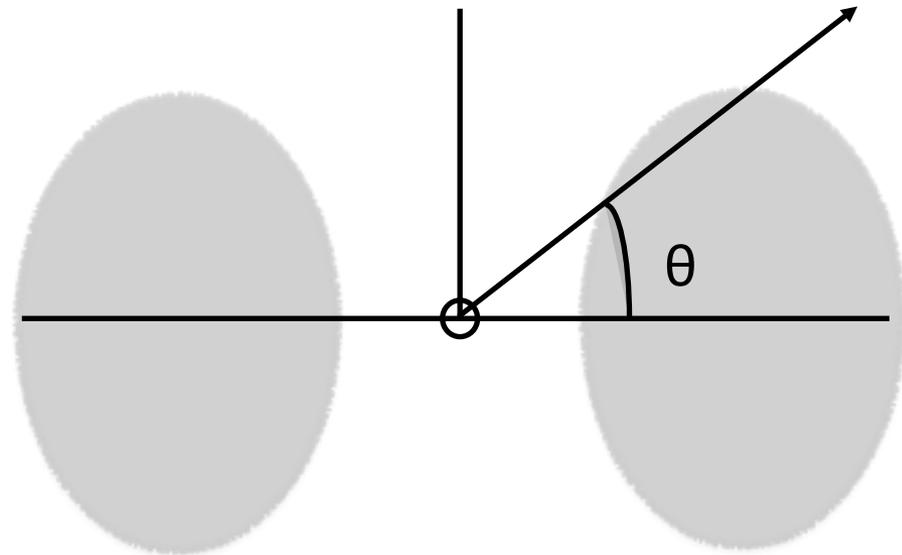
$$(10^{-1.5} < \lambda_{\text{Edd}} < 1)$$

Covering factor  $\sim 40\% + \text{outflows}$

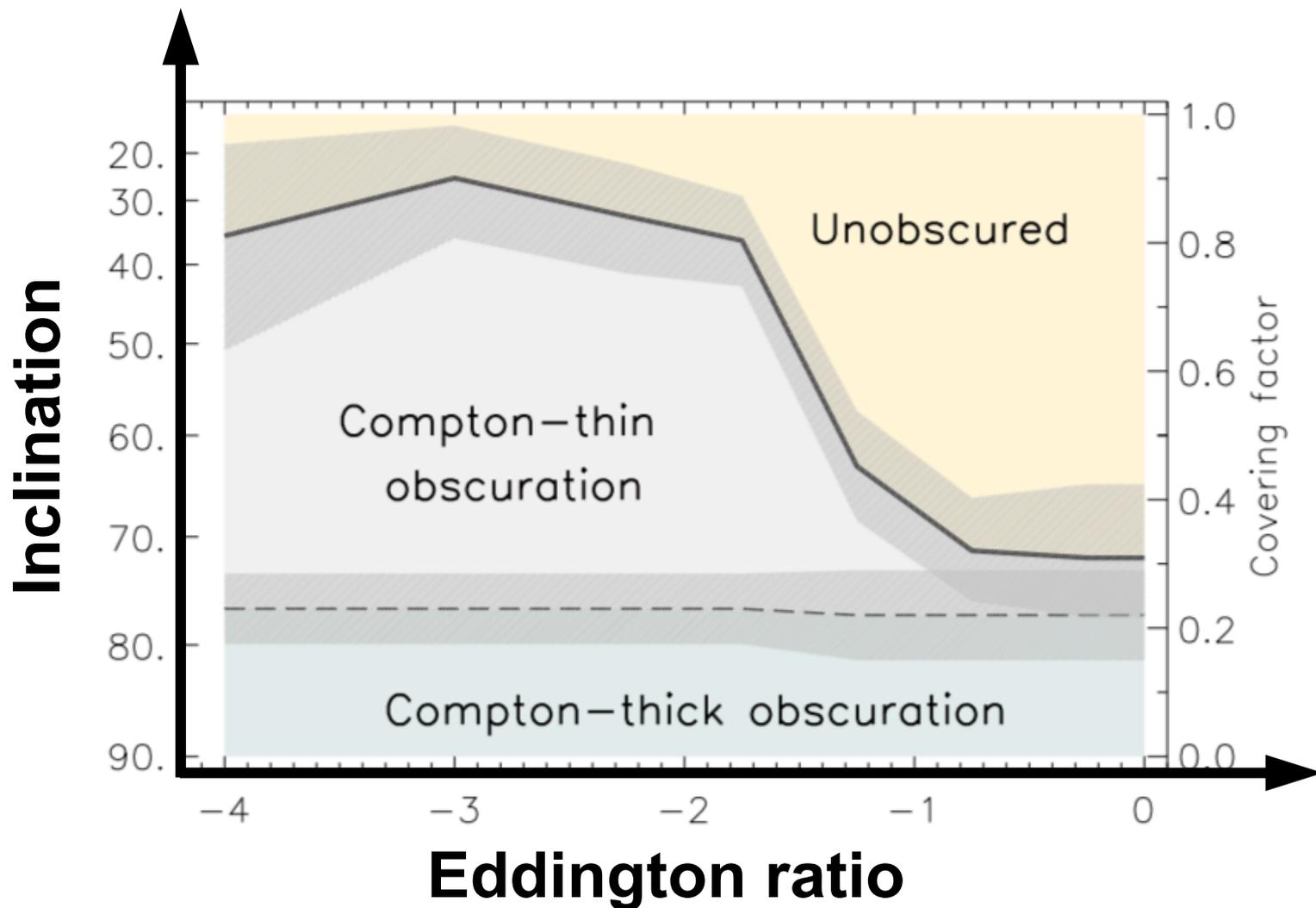


Ricci et al. (2017c, Nature)

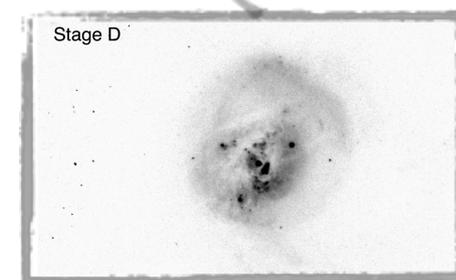
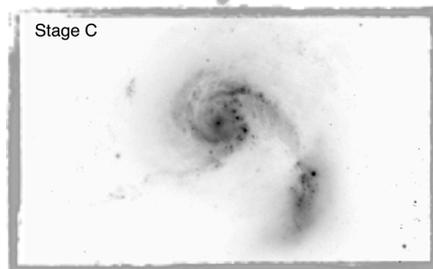
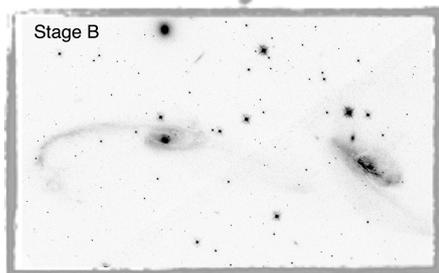
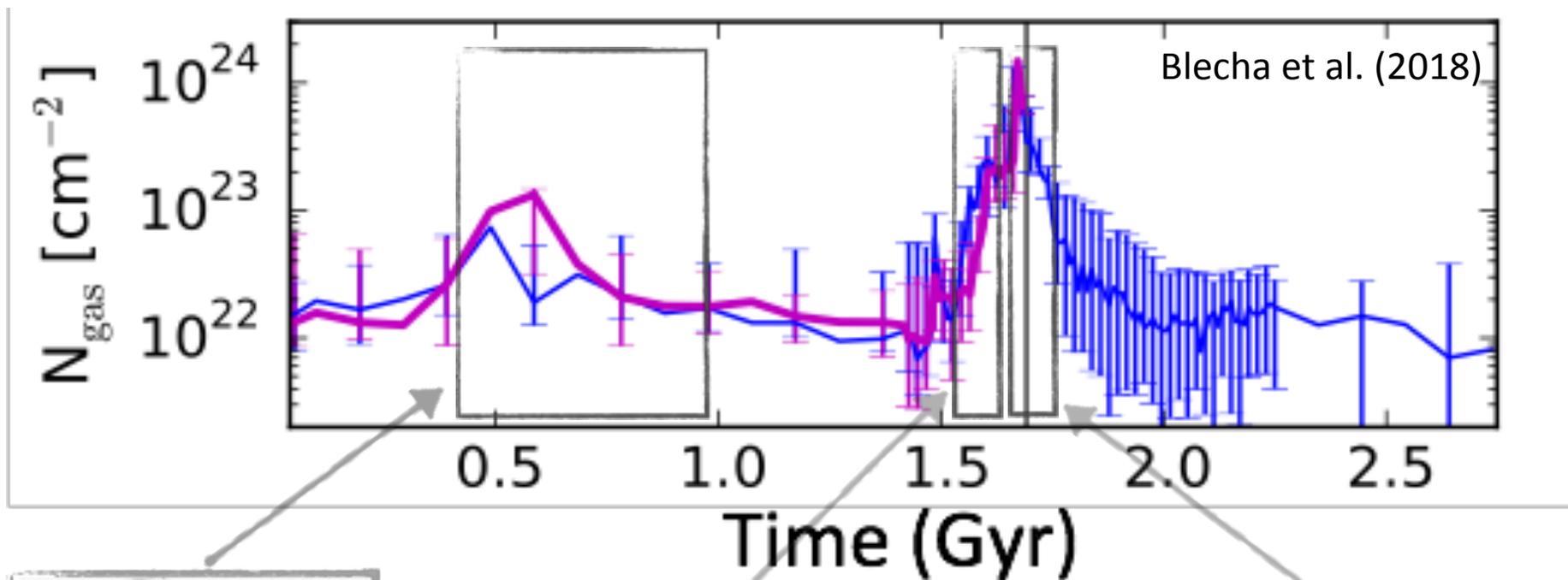
# AGN unification



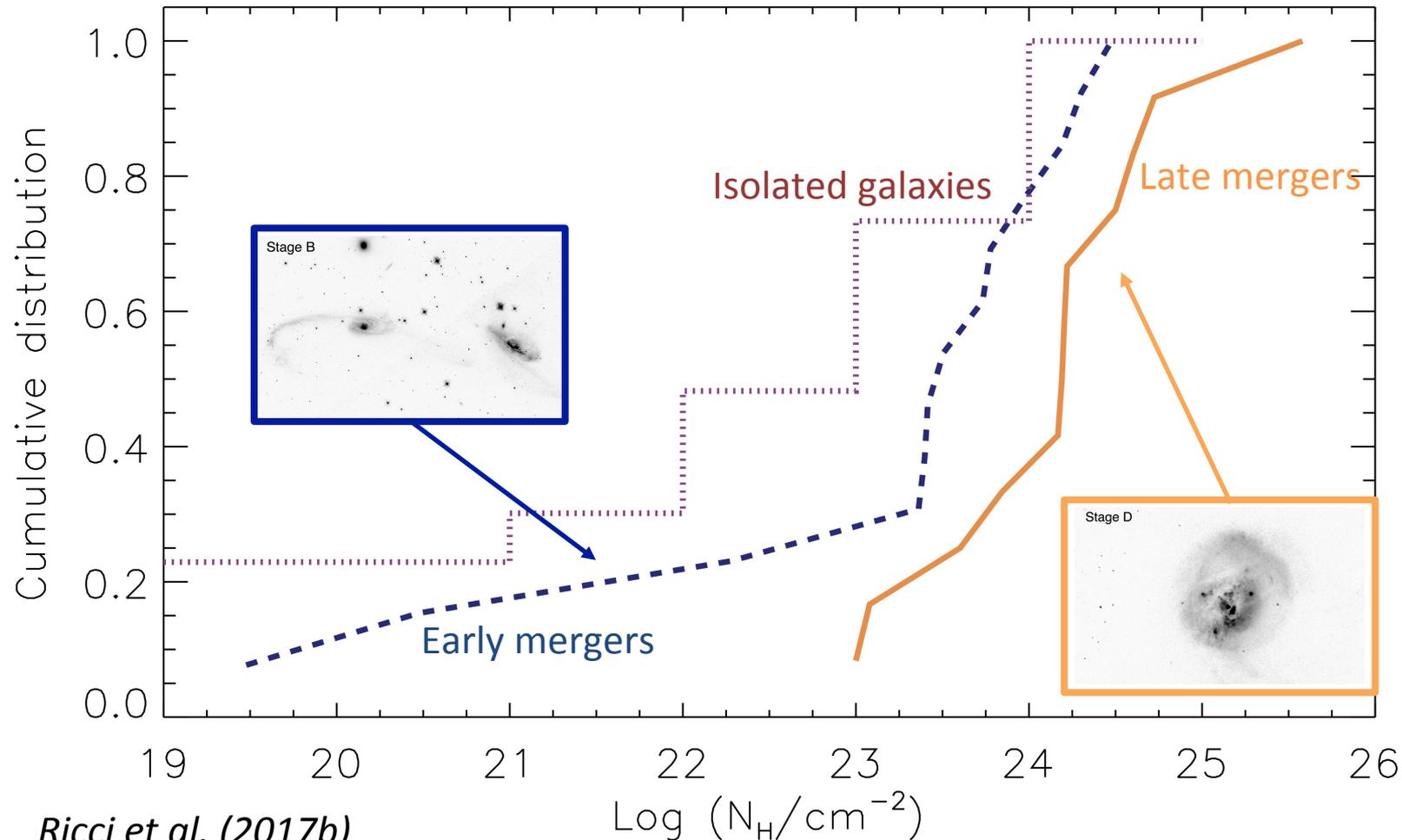
# Radiation-regulated unification



# Galaxy mergers and obscuration



# Galaxy mergers and obscuration

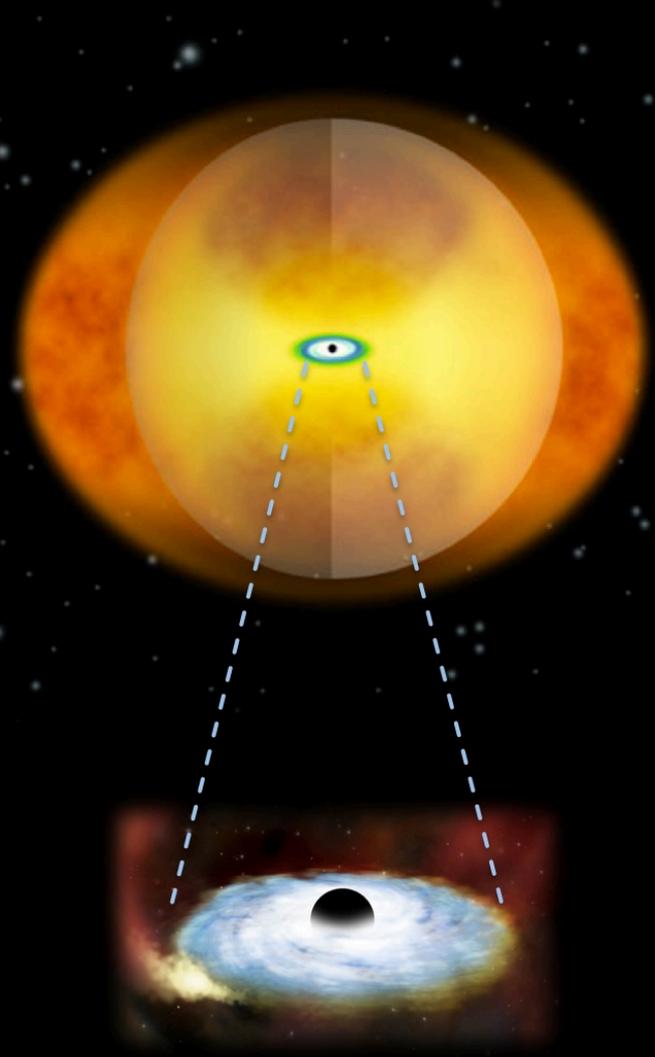
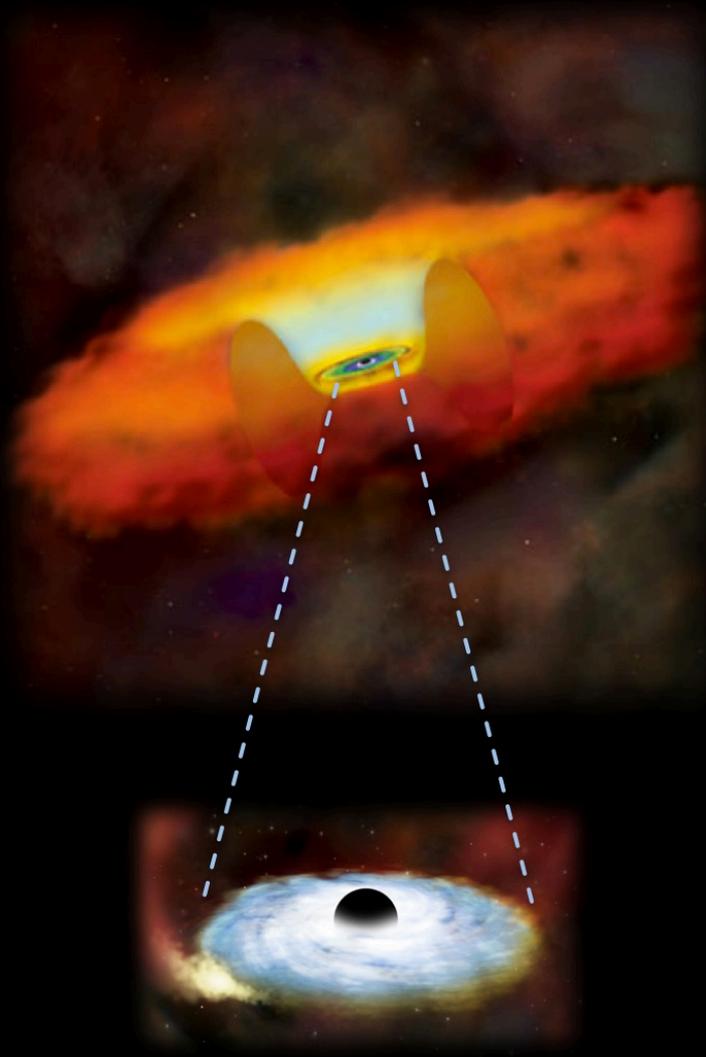


*Ricci et al. (2017b)*

See also Kocevski et al. (2015), Lanzuisi et al. (2015), Lansbury et al. (2017b), Del Moro et al. (2016), Koss et al. (2016), De Rosa et al. (2018), Koss et al. (2018); Talks by R. Pfeifle, E. Treister, C. Carroll

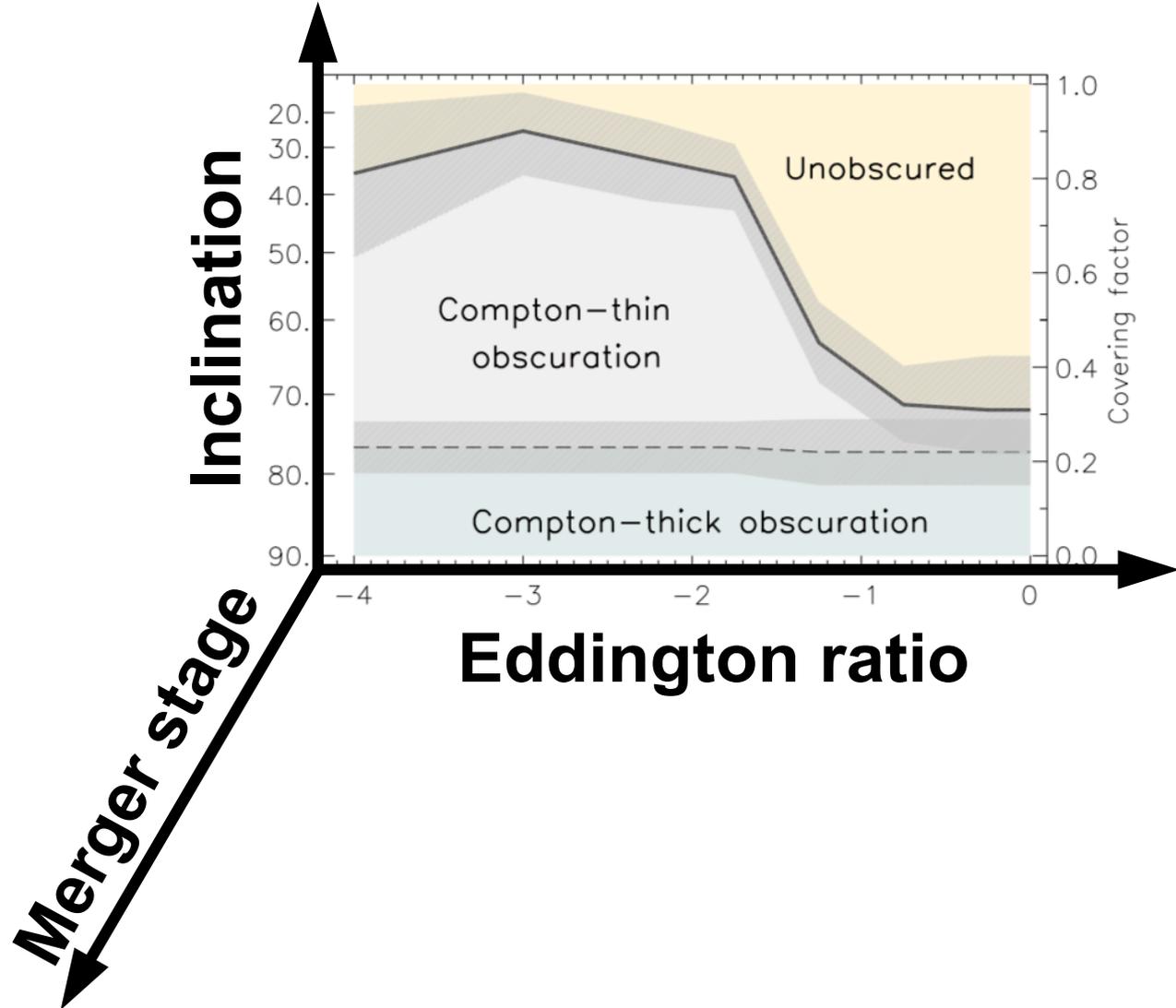
Isolated galaxies

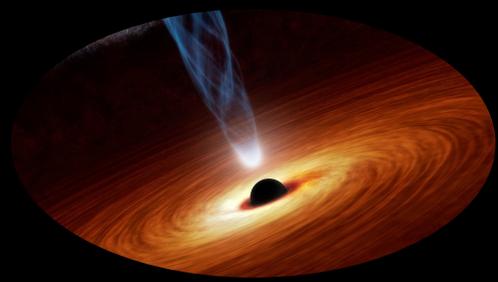
Late stages of merger



From the NASA press release; Credits: NAOJ/NASA/CXC/M. Weiss (Imanishi et al. 2006)

# The drivers of obscuration





# Summary



- X-rays are a great probe of the circumnuclear environment of AGN
- Studies of reprocessed X-ray radiation have shown elongated emission
- More and more complex X-ray spectral models are being developed
- Absorption variability show a dynamic environment
- Hard X-ray surveys have allowed to infer the typical covering factor of local AGN
- Covering factor depends on the Eddington ratio
- Probability of observing obscured AGN is very high in close mergers

# AGN in Santiago

[www.agnsantiago.cl](http://www.agnsantiago.cl)

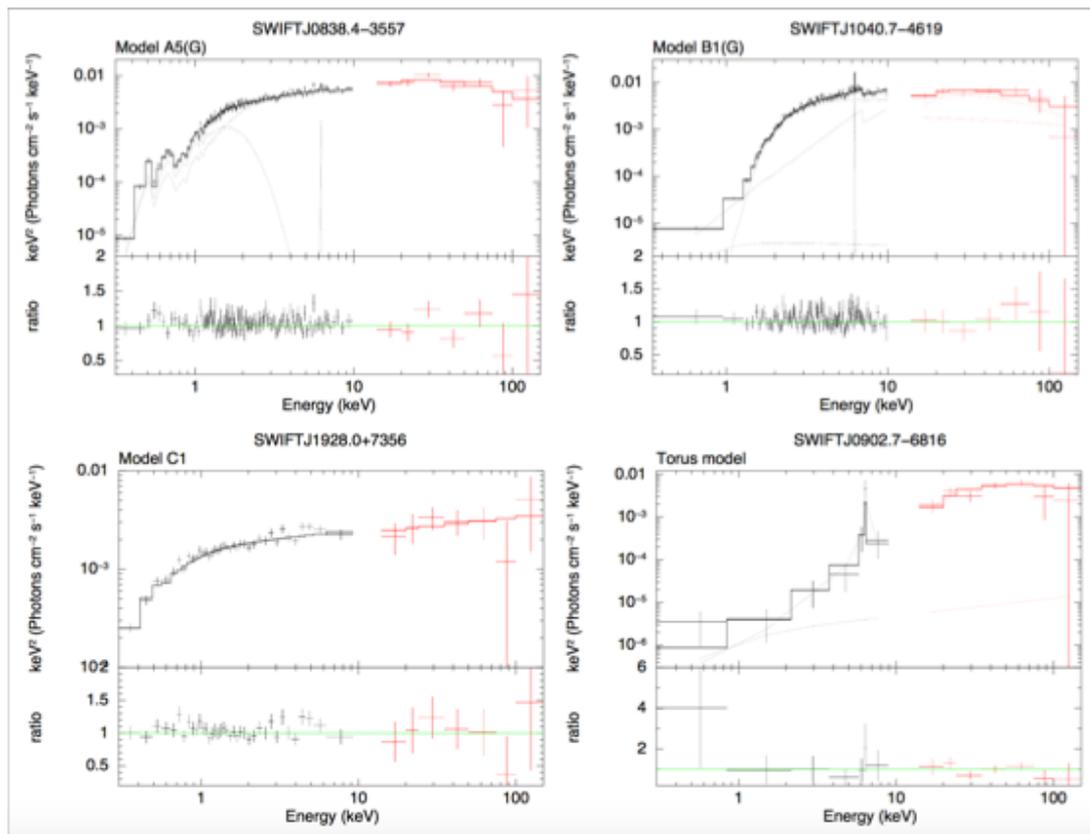


**BACK UP SLIDES**

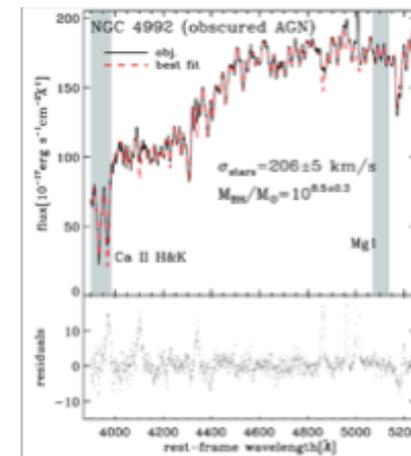
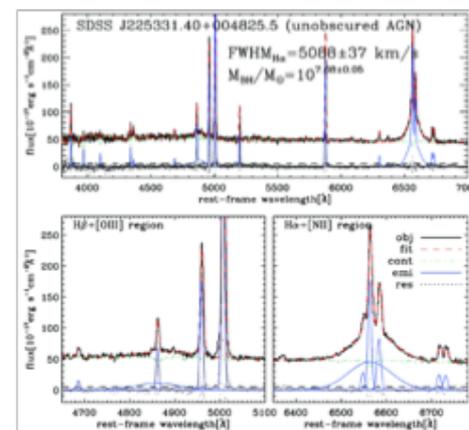
# The BAT AGN Spectroscopic Survey (BASS)



X-ray spectroscopy (Ricci et al. 2017d)

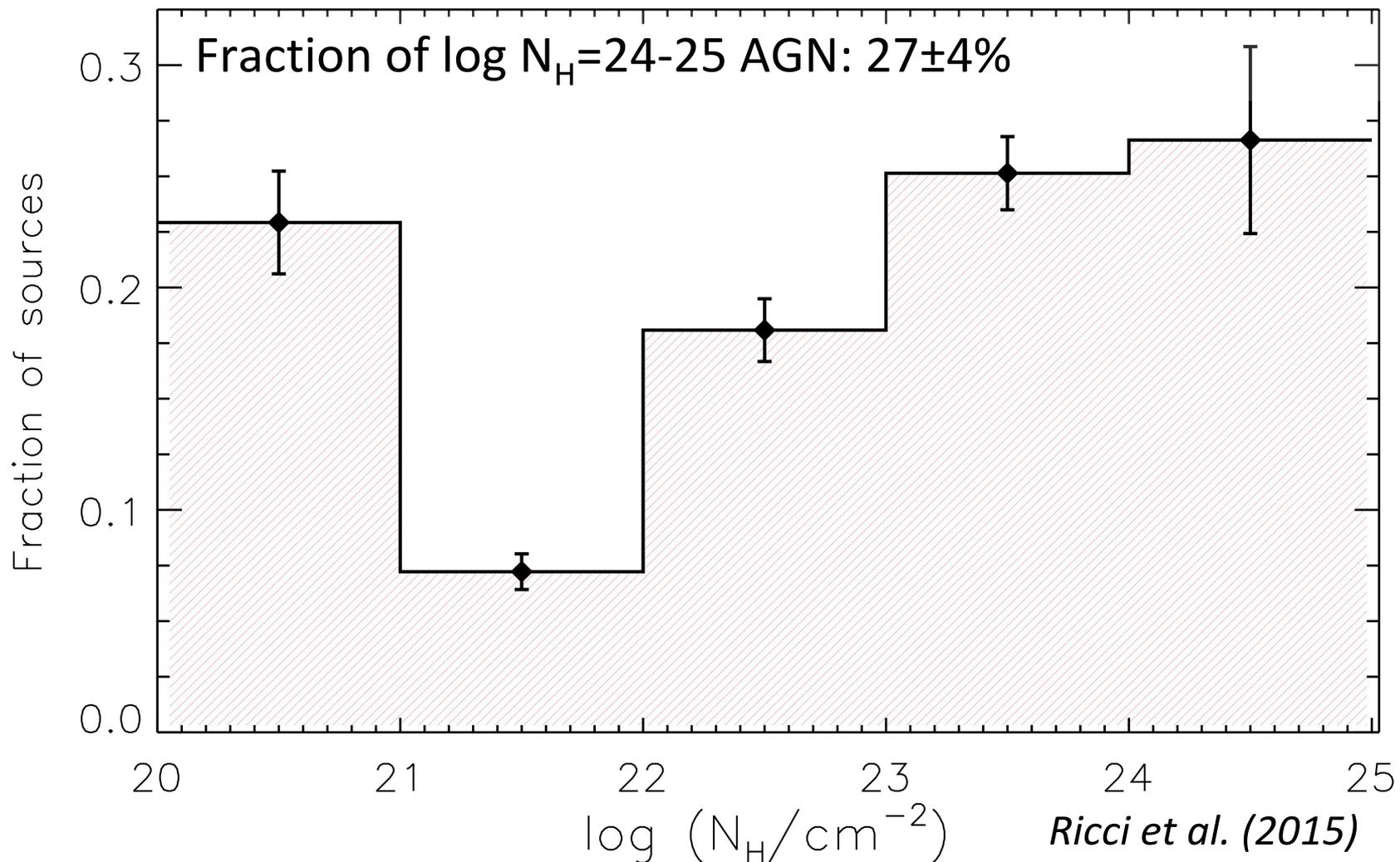


Optical spectroscopy (Koss et al. 2017)



Talk by M. Koss

# Intrinsic column density distribution



# Covering factor of gas and dust



## X-rays

### 1) Fraction of obscured AGN

(e.g., Ueda+03,11, Brightman+11,  
Merloni+14, Buchner+15, Aird+15)

### 2) Spectroscopy with torus models

(e.g., Brightman+15, Balokovic+ prep.)

# Covering factor of gas and dust



## X-rays

### 1) Fraction of obscured AGN

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## Infrared

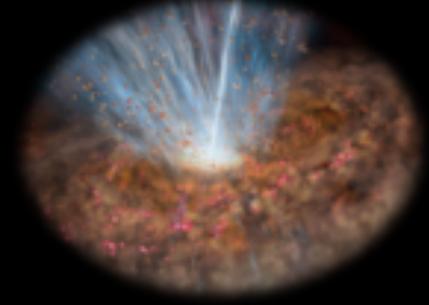
### 1) $L_{\text{IR}}/L_{\text{Bol}}$

(e.g., Maiolino+07, Treister+08,  
Assef+13, Mateos+16, Ichikawa+17)

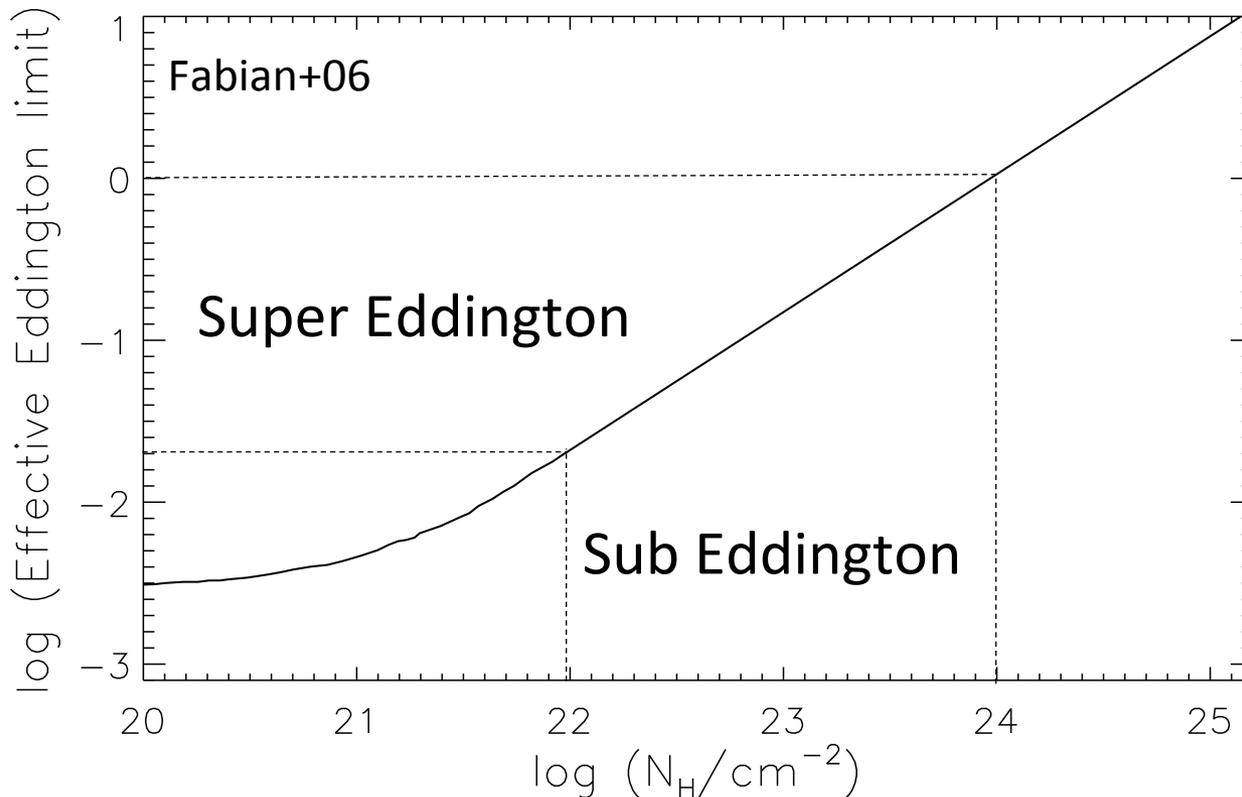
### 2) Spectroscopy with torus models

(e.g., Alonso Herrero+11, Ramos  
Almeida+11)

# Radiation pressure on dusty gas



Effective Eddington limit  $\lambda_{Edd}^{eff} = \sigma_T / \sigma_i (N_H; \xi)$



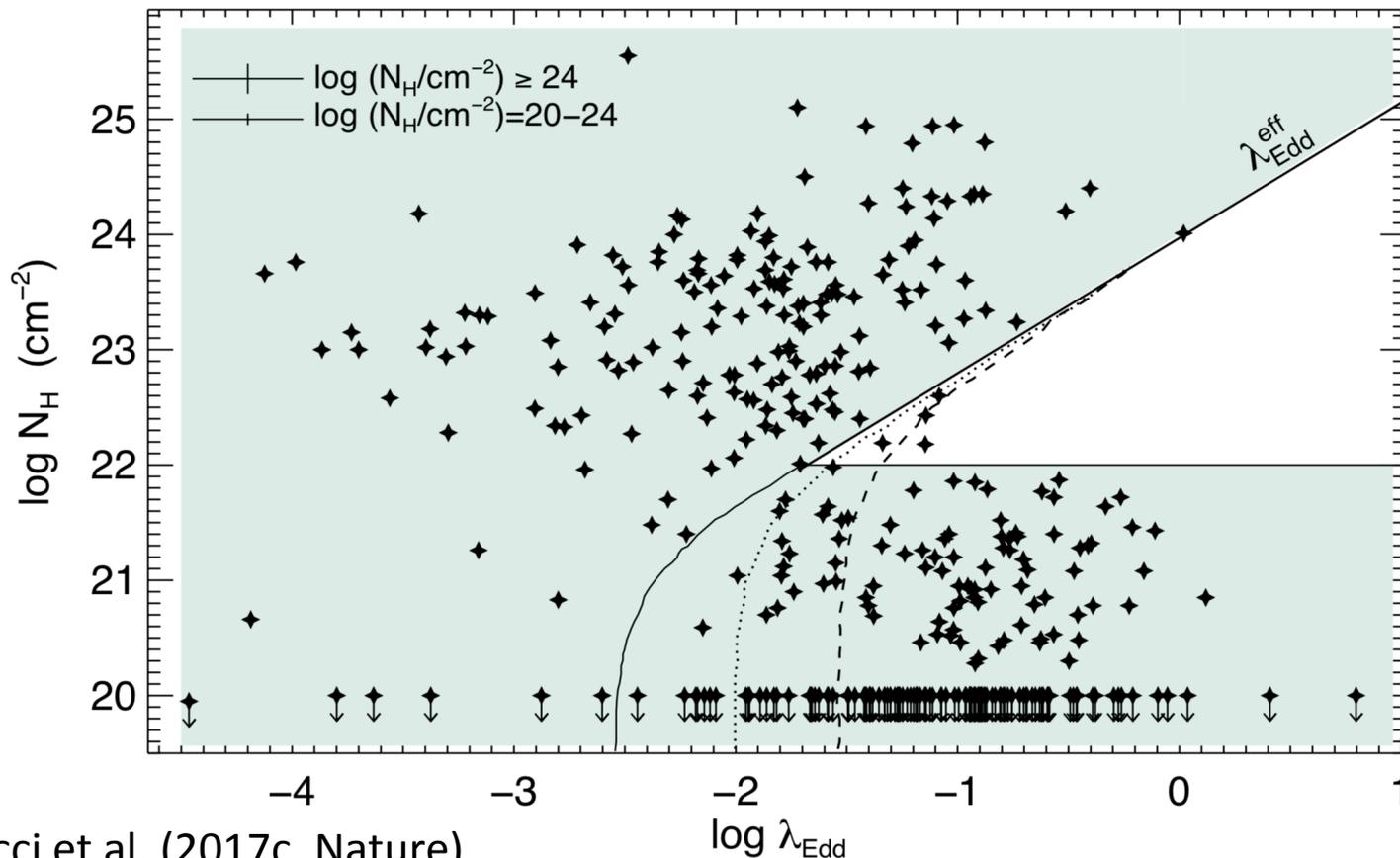
see also [Fabian+08](#); [Fabian+09](#); [Raimundo+10](#); [Vasudevan+13](#),  
[Ishibashi+18](#)

# Radiation pressure on dusty gas



Effective Eddington limit

$$\lambda_{\text{Edd}}^{\text{eff}} = \sigma_T / \sigma_i(N_H; \xi)$$



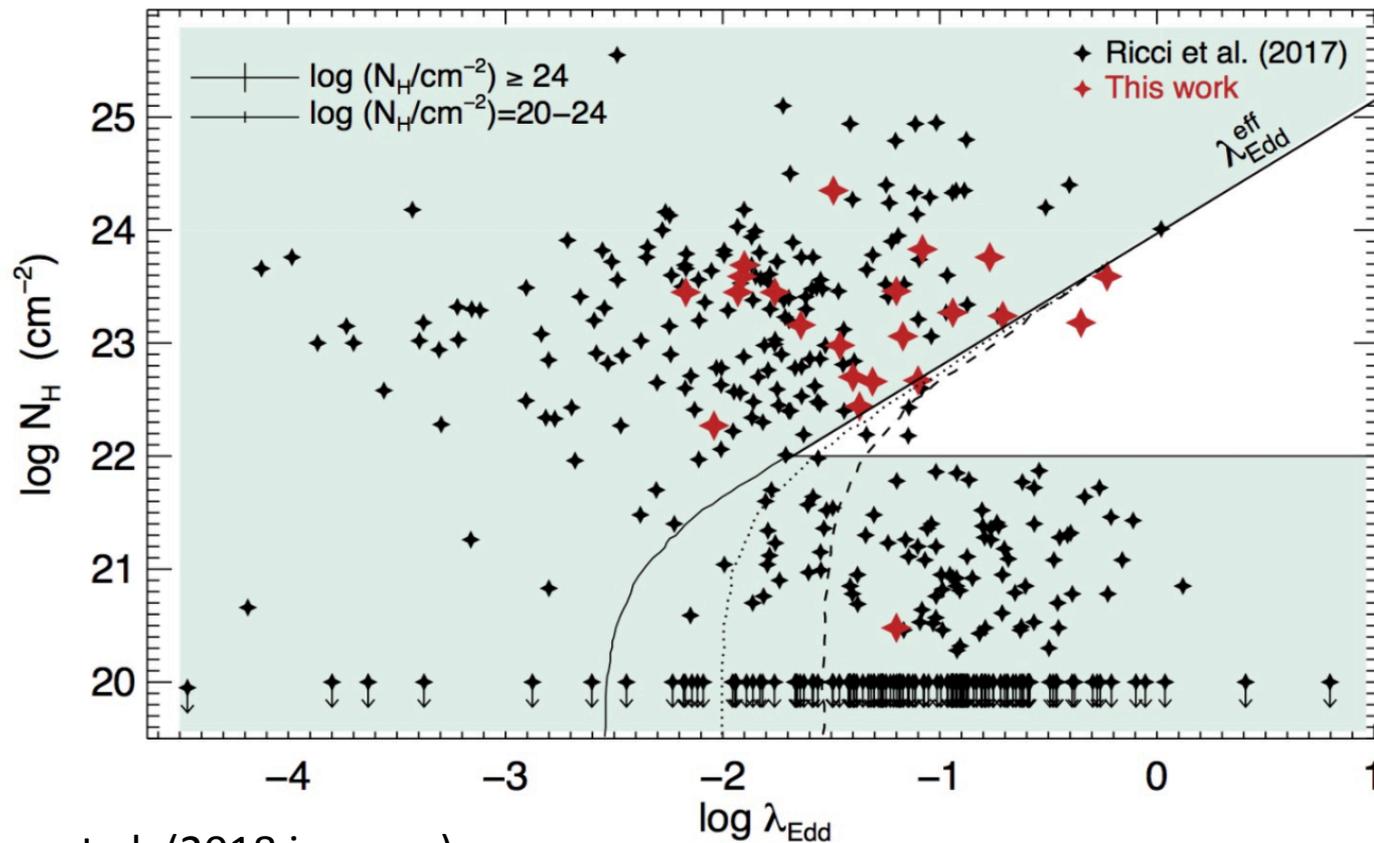
Ricci et al. (2017c, Nature)

# Radiation pressure on dusty gas



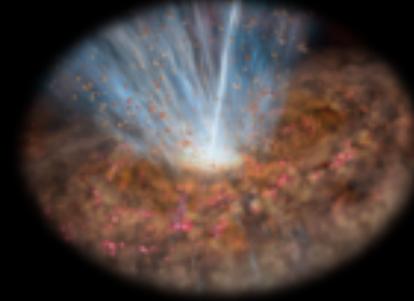
Effective Eddington limit

$$\lambda_{\text{Edd}}^{\text{eff}} = \sigma_T / \sigma_i(N_H; \xi)$$

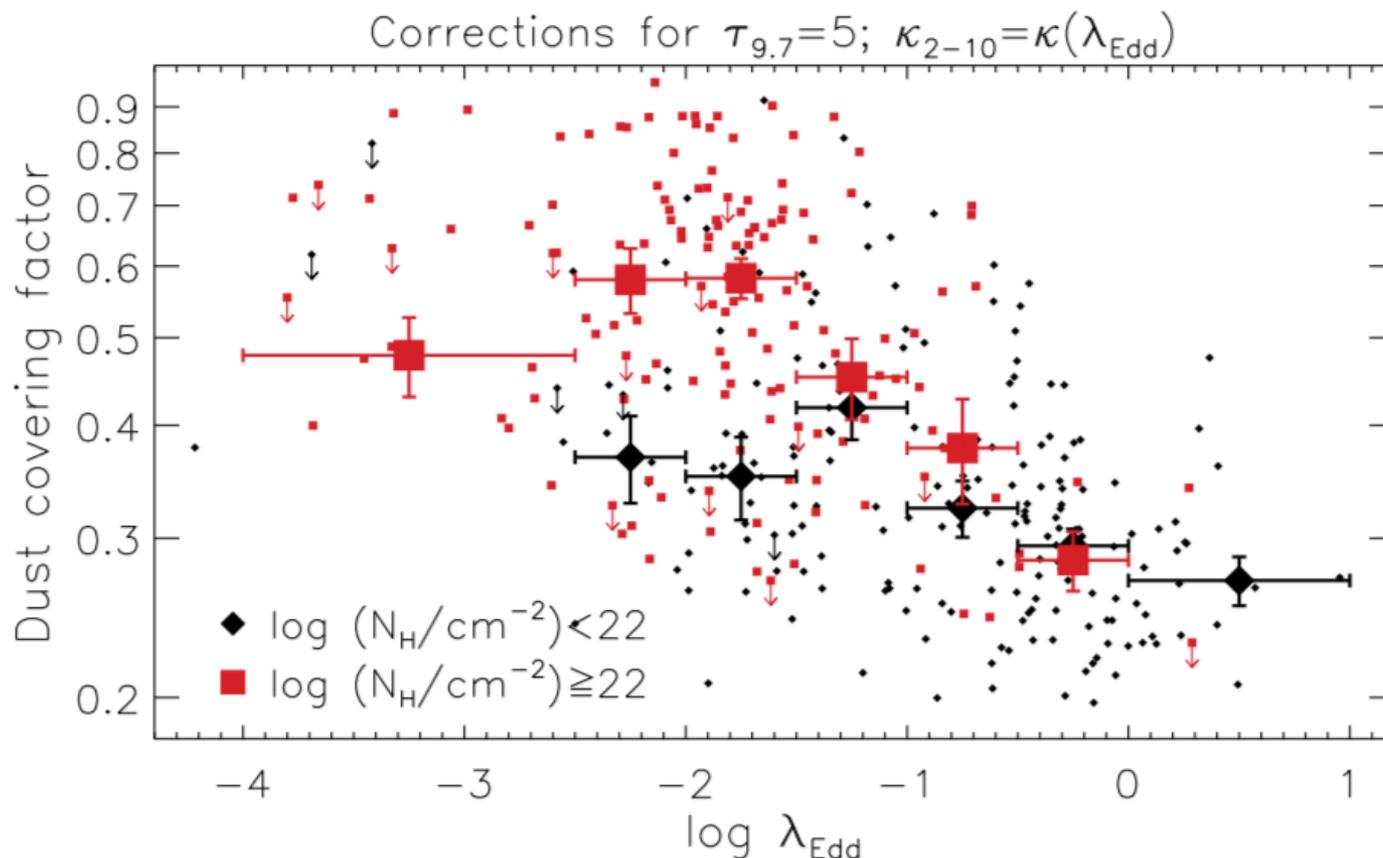


Baer et al. (2018 in prep.)

# The covering factor of dust

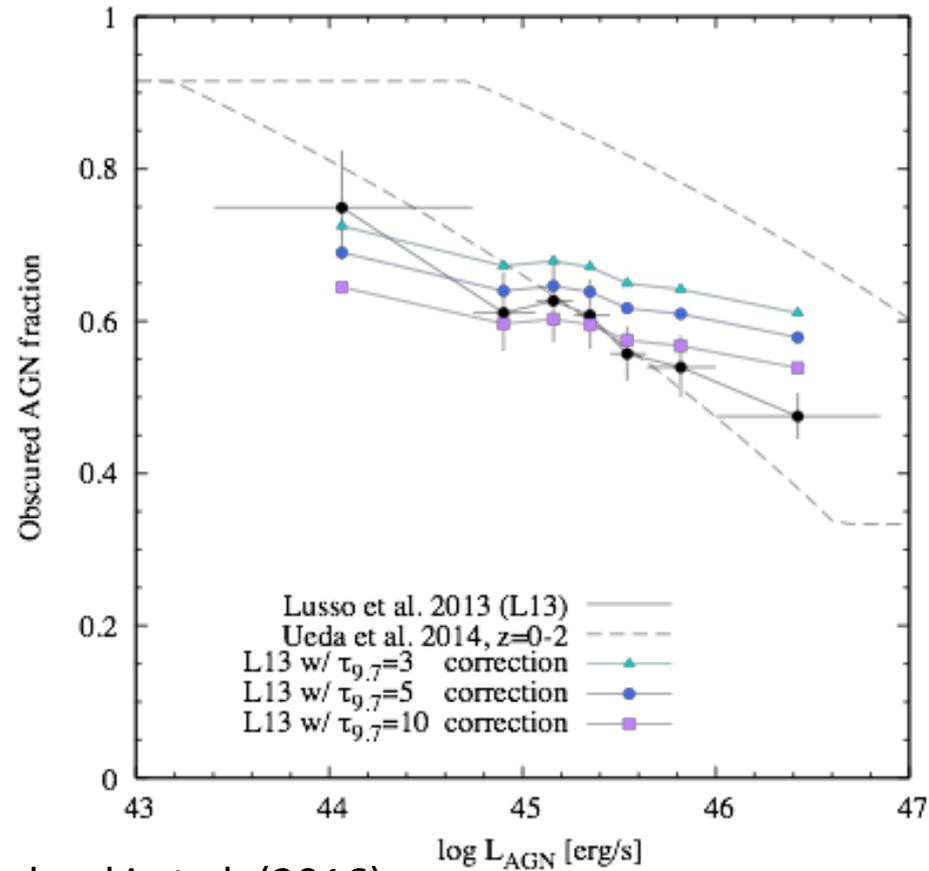


Covering factor from  $L_{\text{IR}}/L_{\text{Bol}}$ ;  $\sim 360$  BASS AGN



Ricci et al. (2018, in prep.); see also Ezhikode et al. (2017), Zhuang et al. (2018)

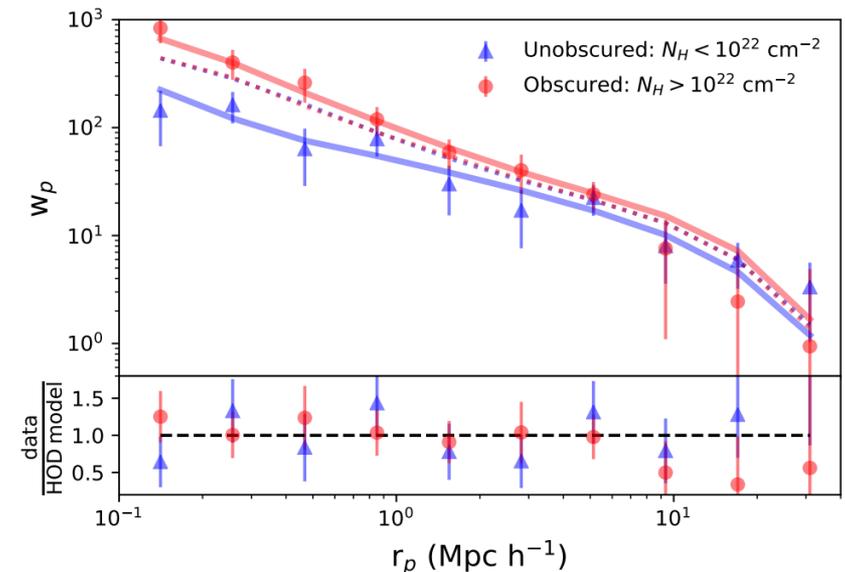
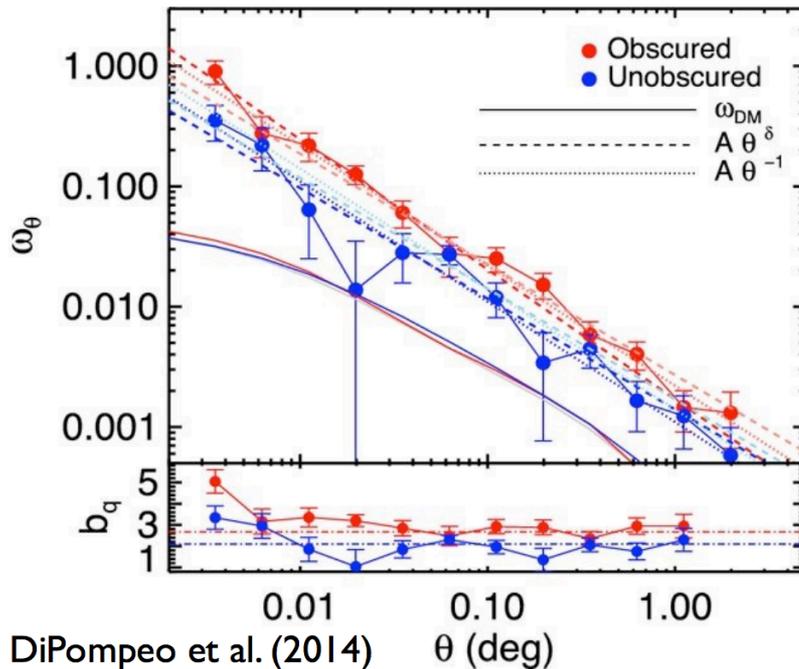
# Luminosity dependence of obscuration - IR



Stalevski et al. (2016)

see also Maiolino et al. (2007), Treister et al. (2008), Assef et al. (2013), Ichikawa et al. (2017), talk by S. Mateos

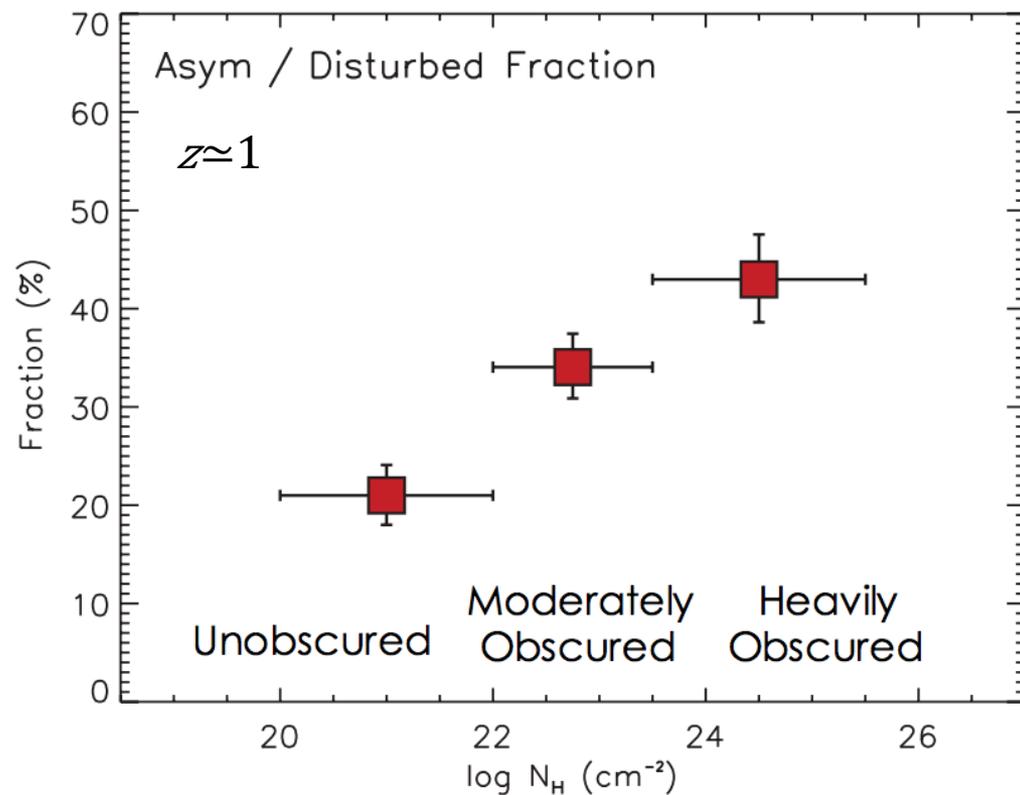
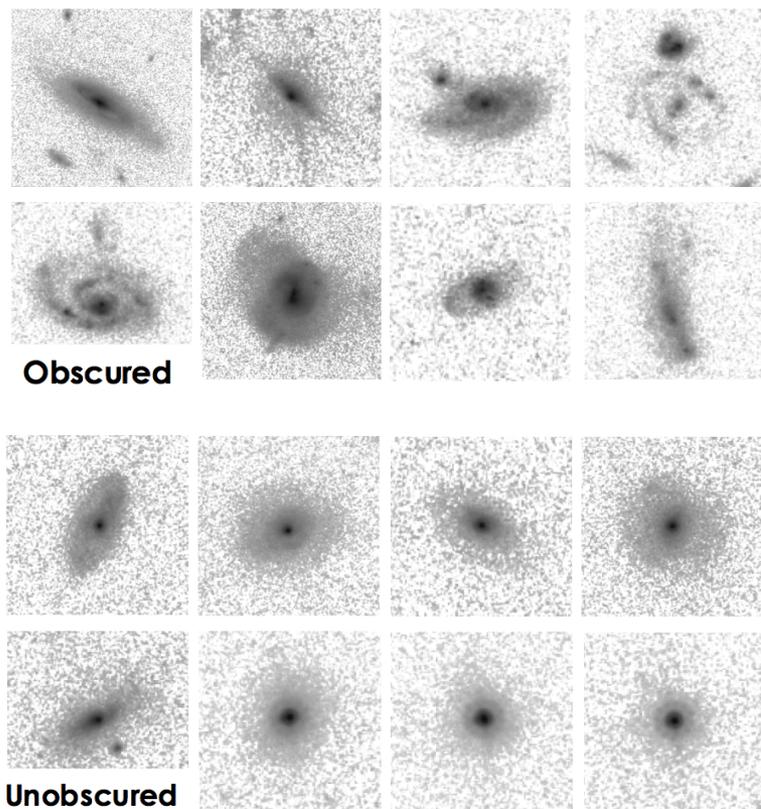
# Clustering and obscuration



See also Hickox et al. (2009), Allevato et al. (2011, 2014), Donoso et al. (2014), Di Pompeo et al. (2014, 2017a,b). See talk by V. Allevato, A. Masini

**Obscured AGN tend to reside in denser environments than unobscured AGN**

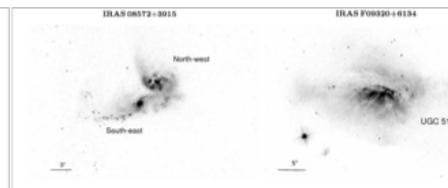
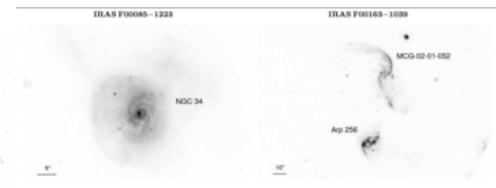
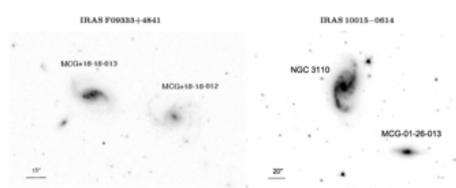
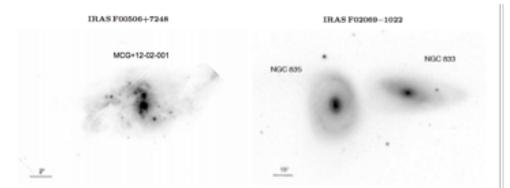
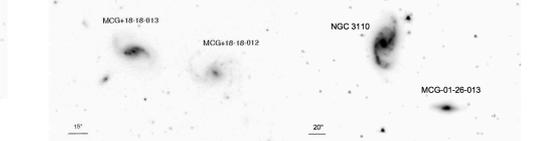
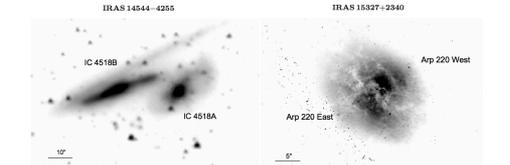
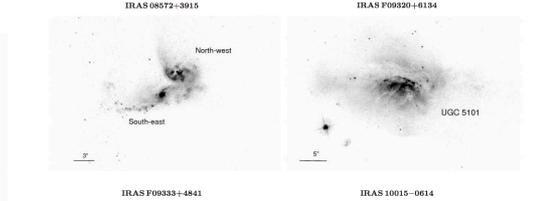
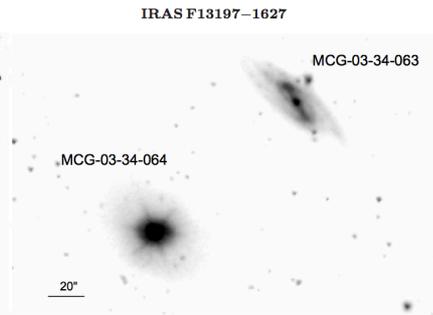
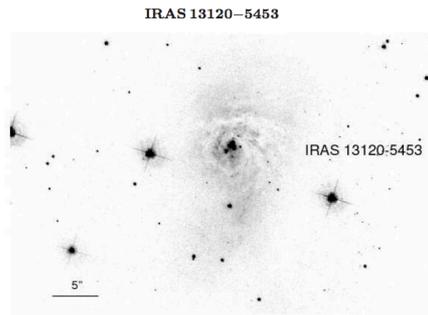
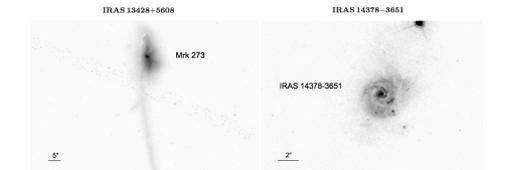
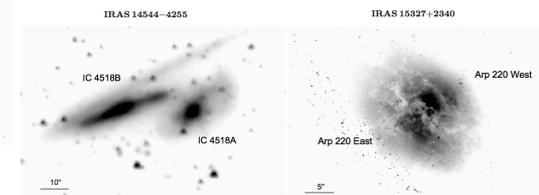
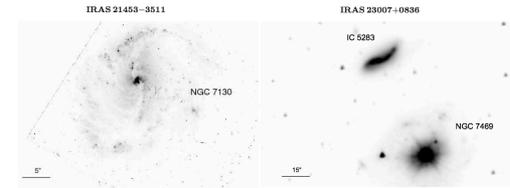
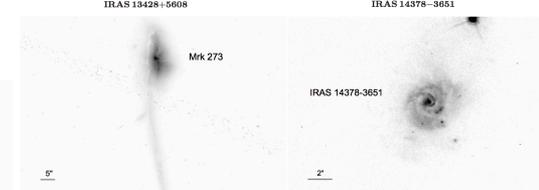
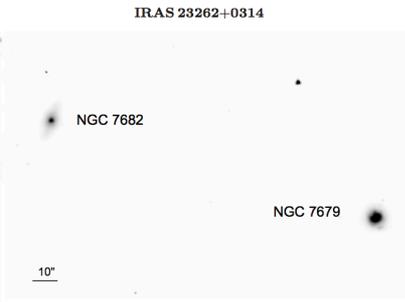
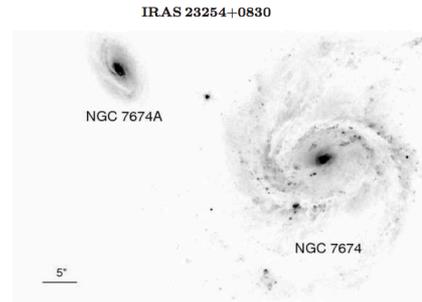
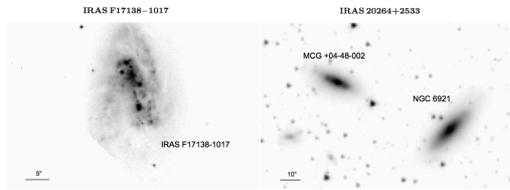
# Galaxy mergers and obscuration



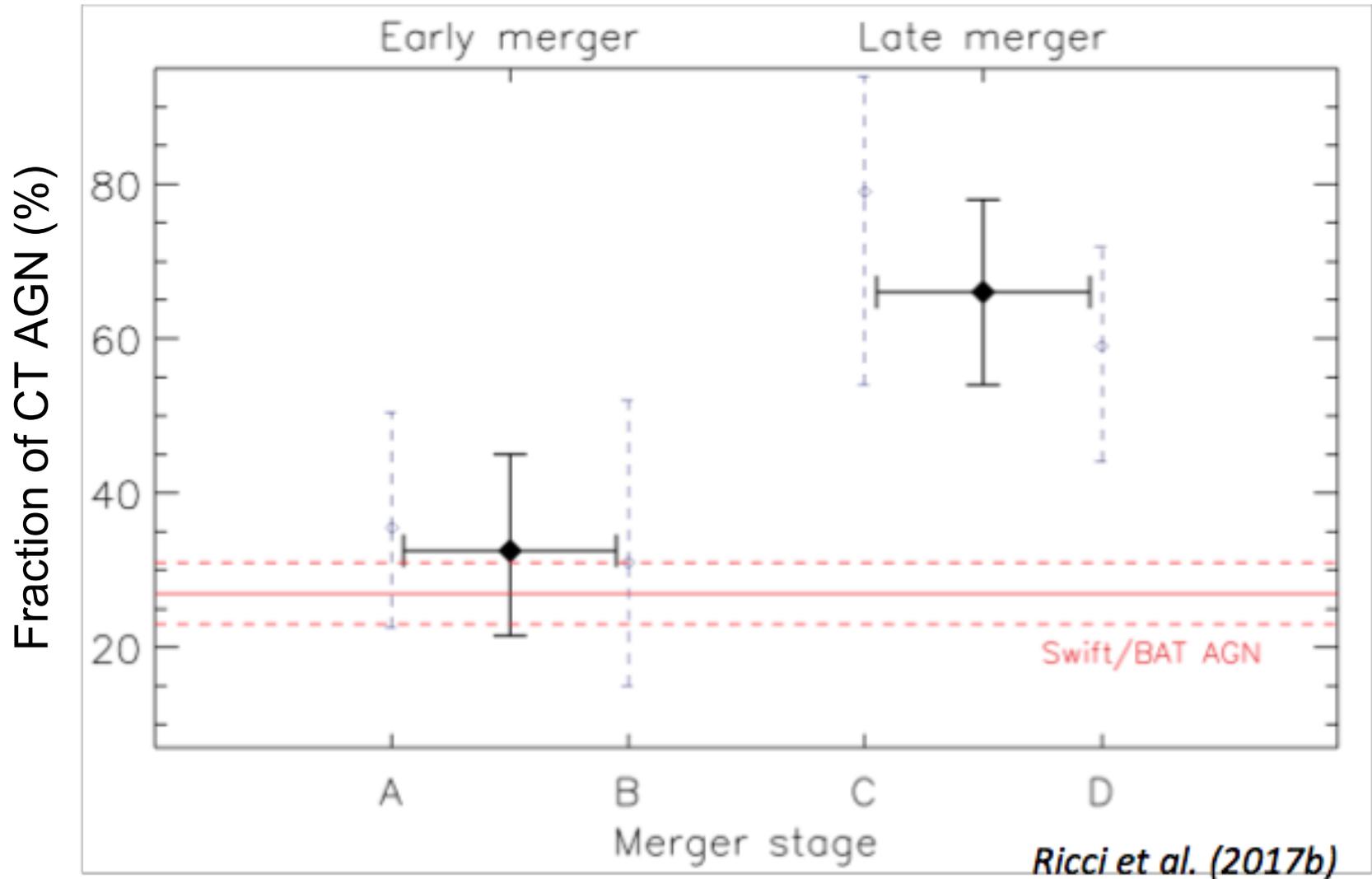
Kocevski et al. (2015)

See also Lanzuisi et al. (2015), Lansbury et al. (2017b), Del Moro et al. (2016), Koss et al. (2016), De Rosa et al. (2018), Koss et al. (2018, submitted)

# Obscuration properties of mergers



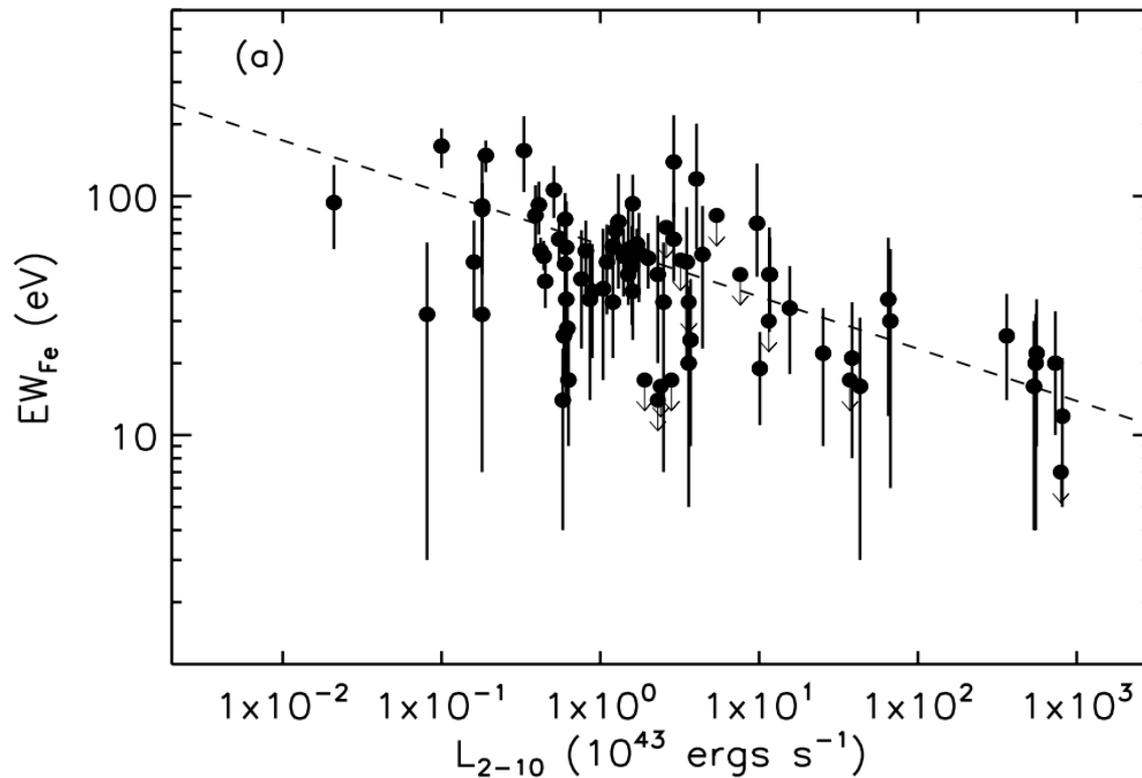
# Obscuration properties of mergers



# The Iwasawa-Taniguchi effect

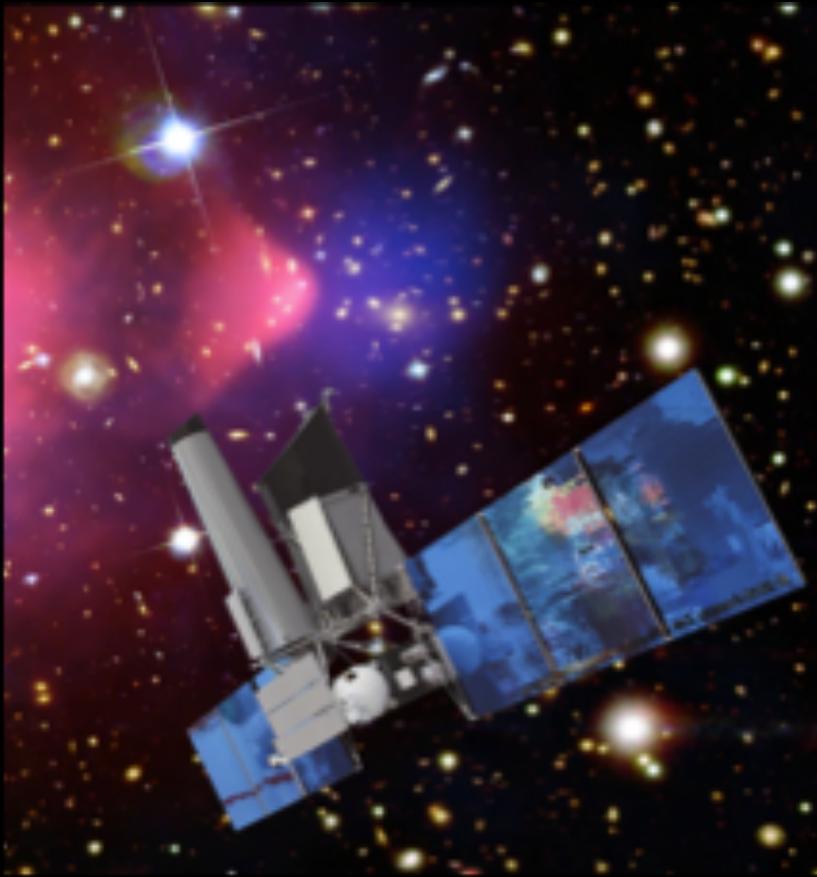


Iwasawa & Taniguchi (1993), Bianchi et al. (2007), Shu et al. (2010, 2011), Ricci et al. (2013,2014), Boorman et al. (2018)



Shu et al. (2010), see P. Boorman's talk

# X-ray studies of obscured AGN: the future



Talk by M. Salvato

# X-ray studies of obscured AGN: the future

